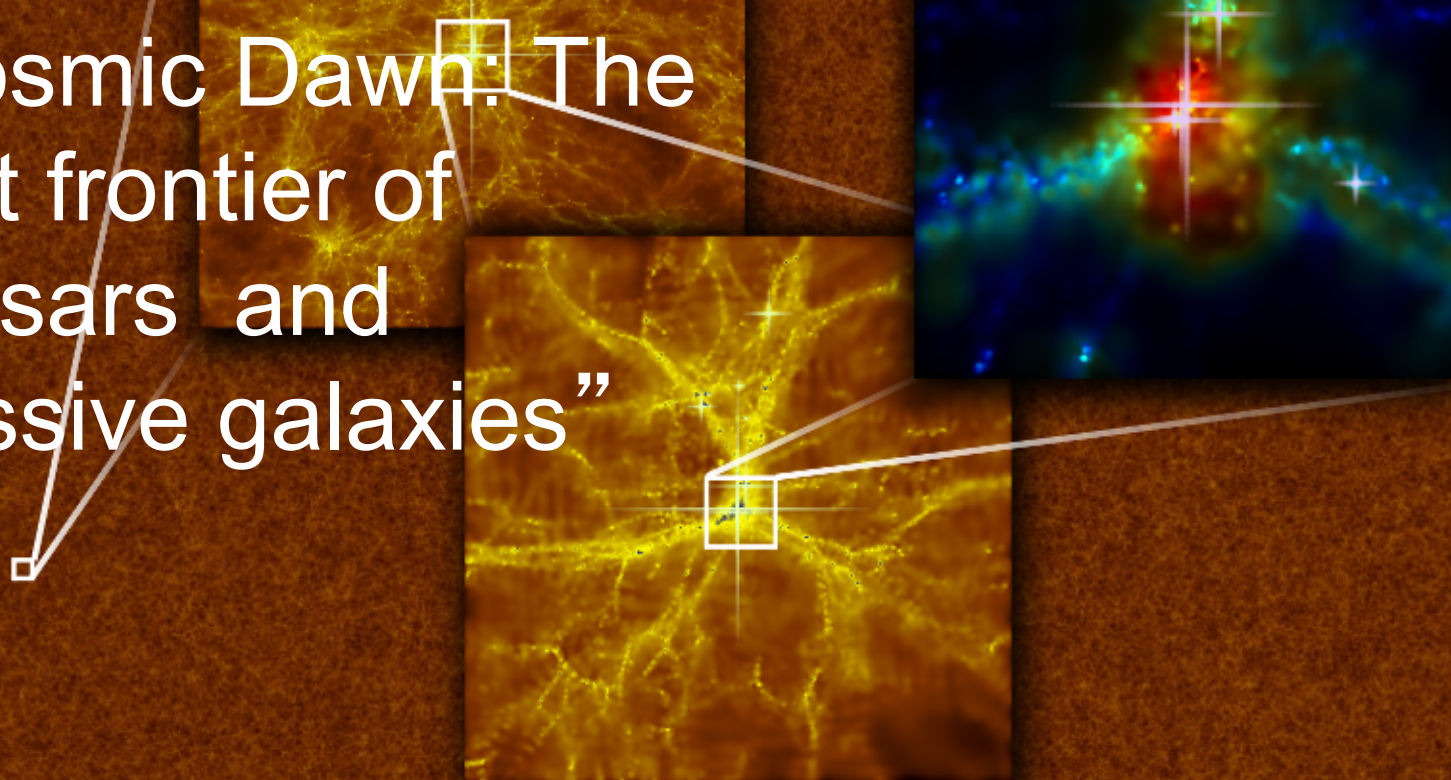


# “Cosmic Dawn: The next frontier of quasars and massive galaxies”



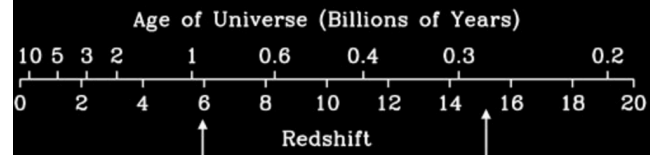
**Tiziana DiMatteo (CMU)**

Yu Feng (Berkeley), Dacen Waters (CMU), Rupert Croft (CMU)  
Simeon Bird (CMU), Steve Wilkins (Sussex), Ananth Tennesi (CMU)  
Nick Battaglia (Princeton), Mark Straka (NCSA)

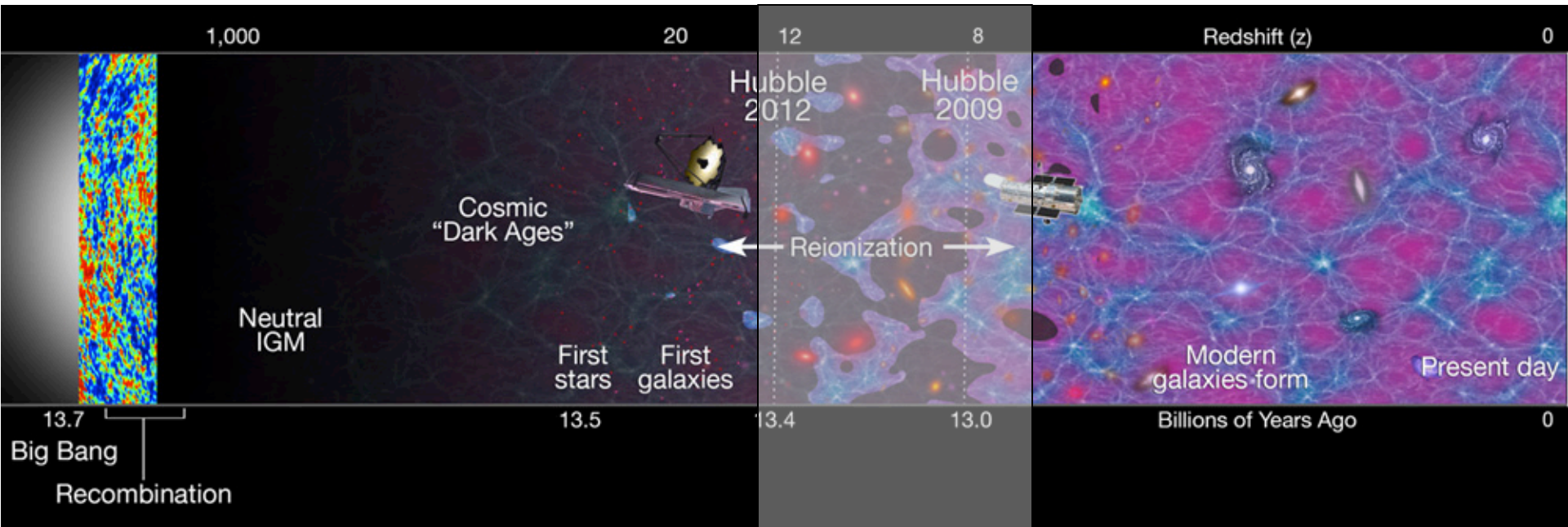
<http://bluetides-project.org>

100 MILLION  
LIGHT YEARS





# A sketch of Cosmic History



300Myrs

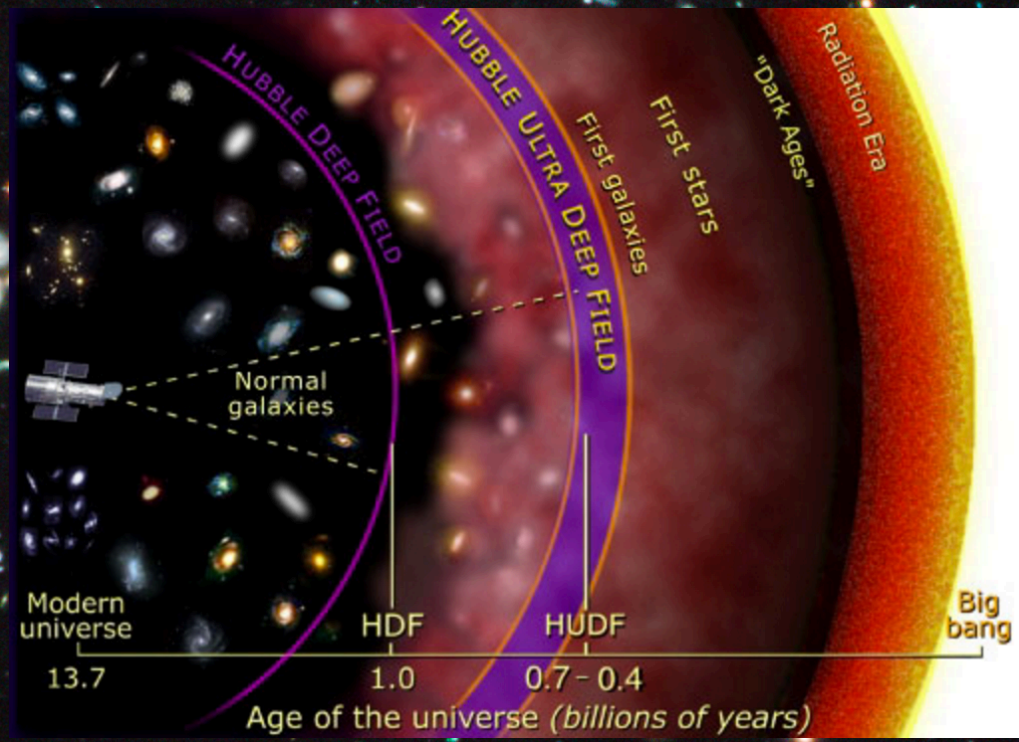
800Myrs



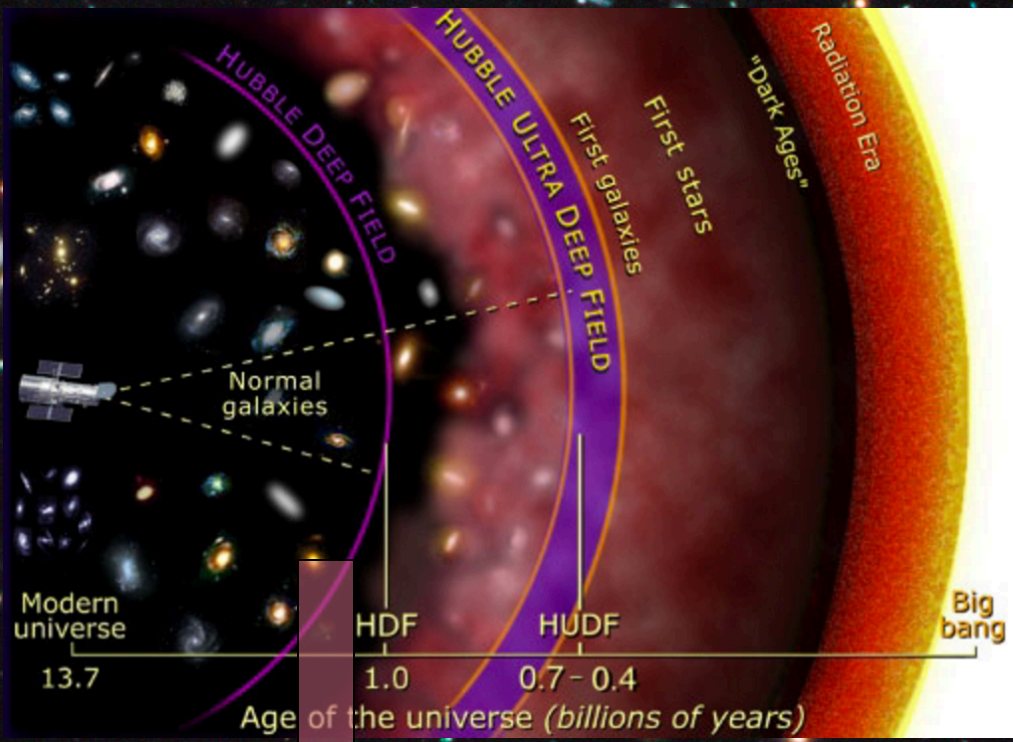
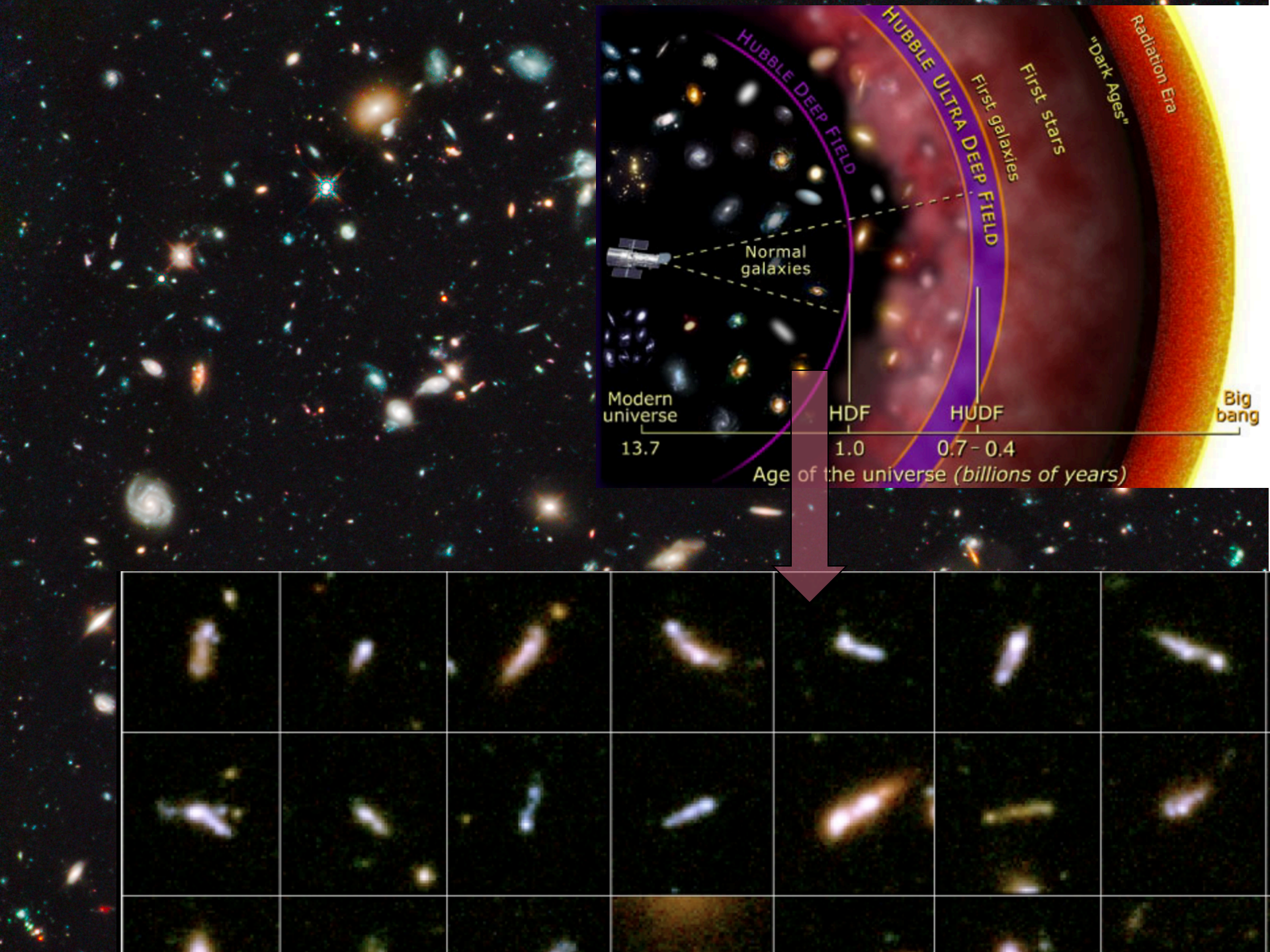
The first 800 million years ( $z=7+$ ):

Observations:

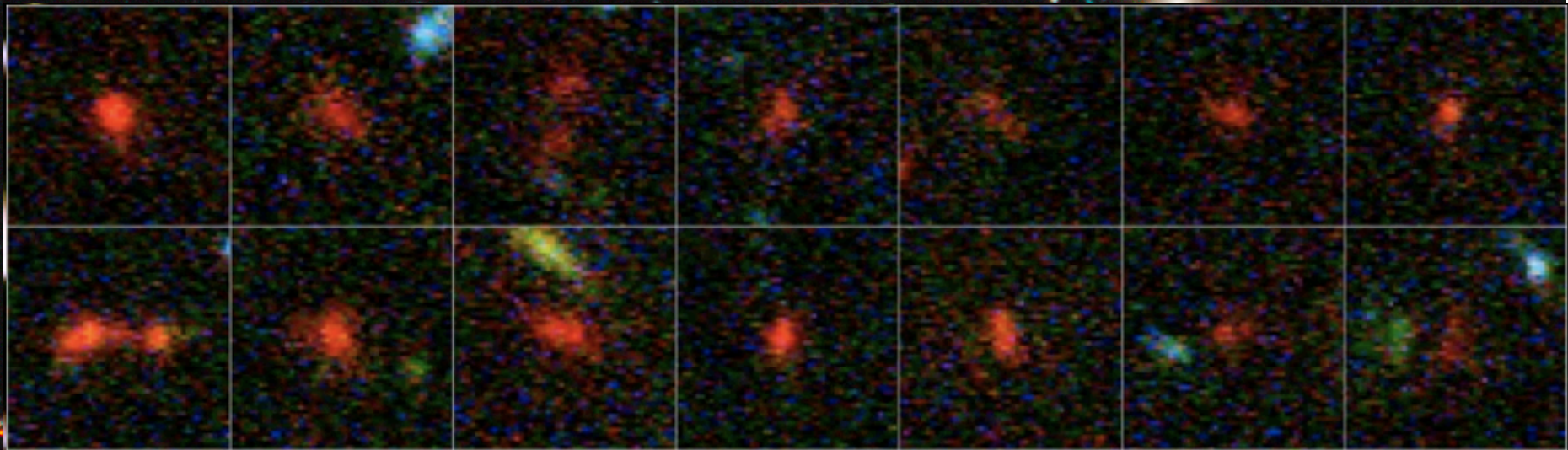
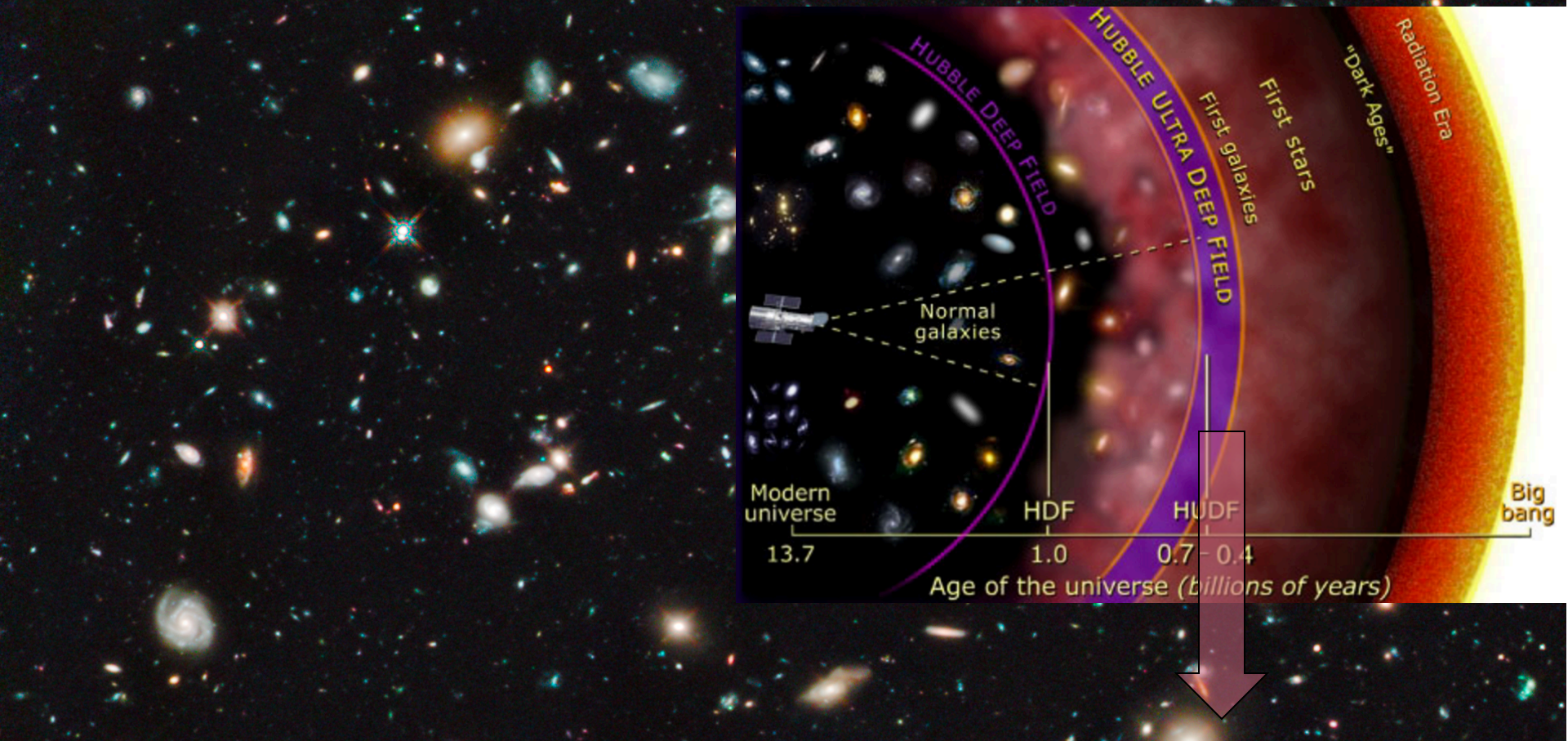




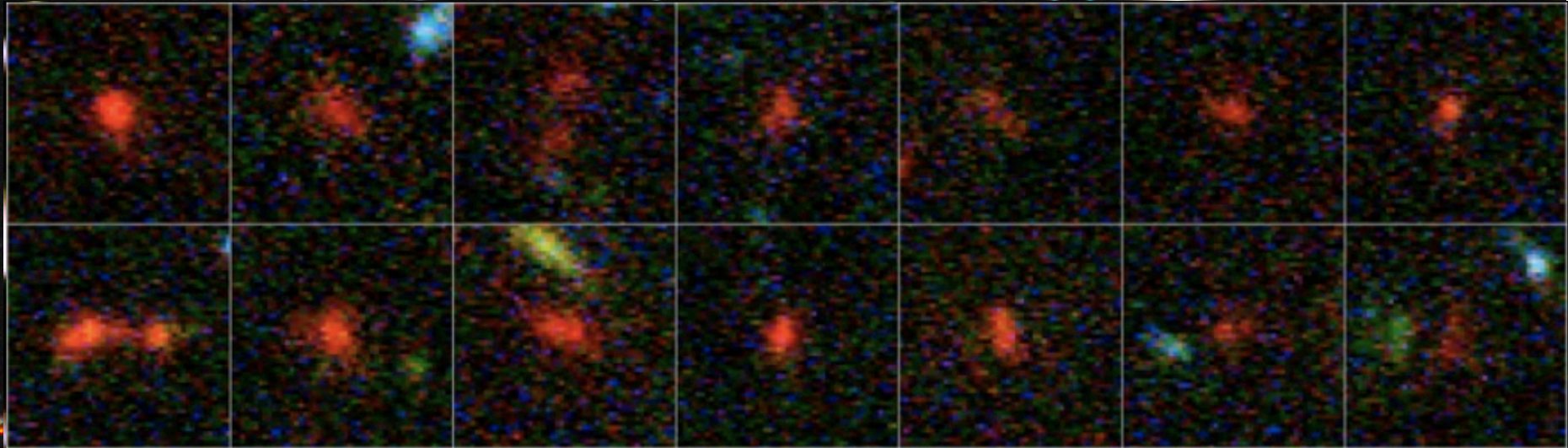
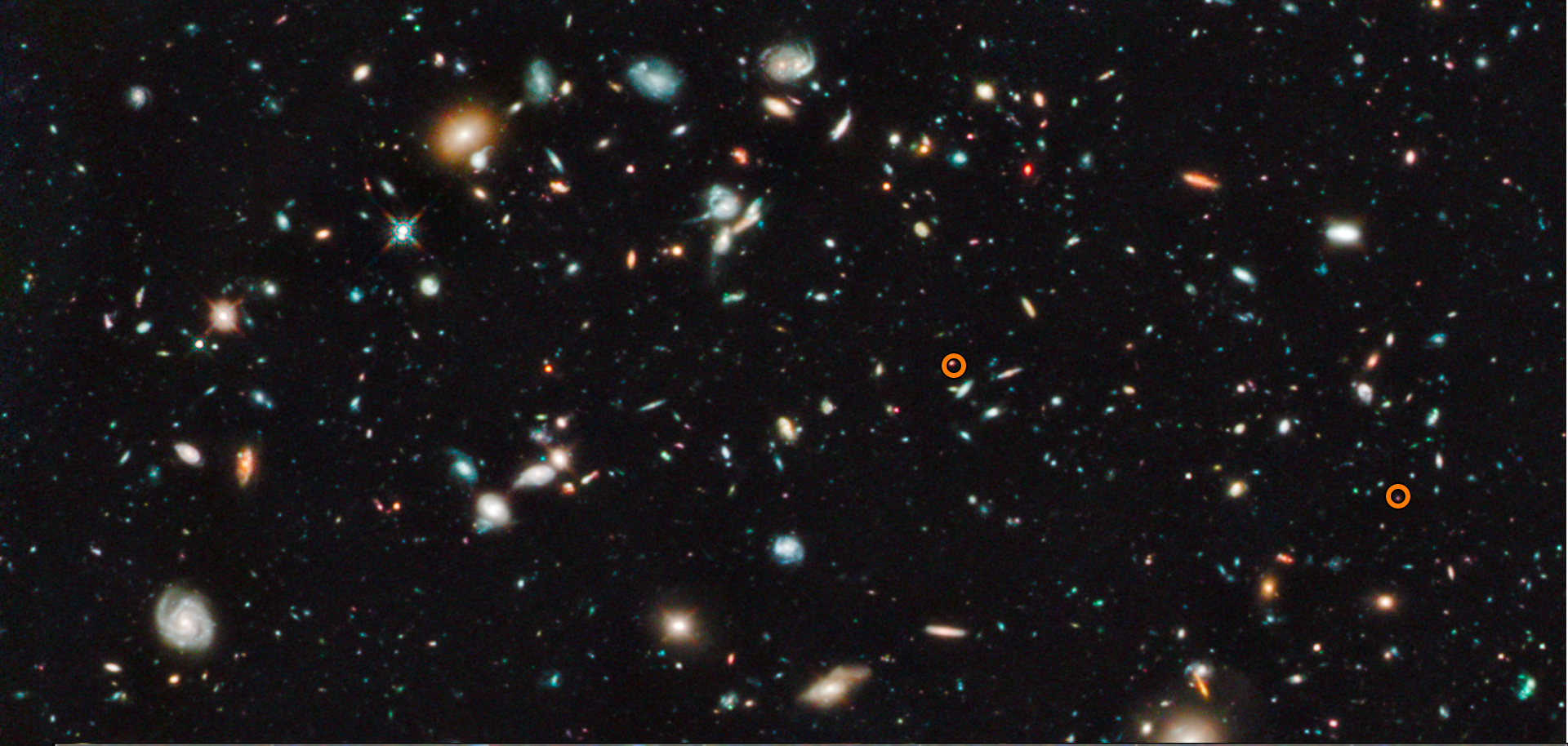










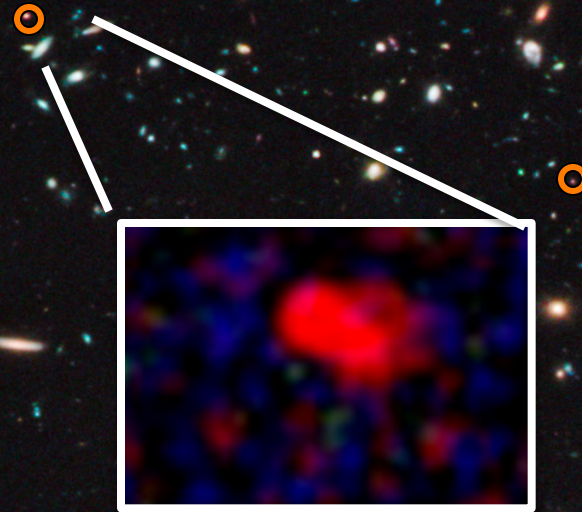




500-700 Million years: Faint, compact smudges



2 million light years away



13 billion light years away

Tiny (few arcmin<sup>2</sup>) **Hubble Legacy Ultra Deep Fields** show compact, clumpy and irregular morphologies

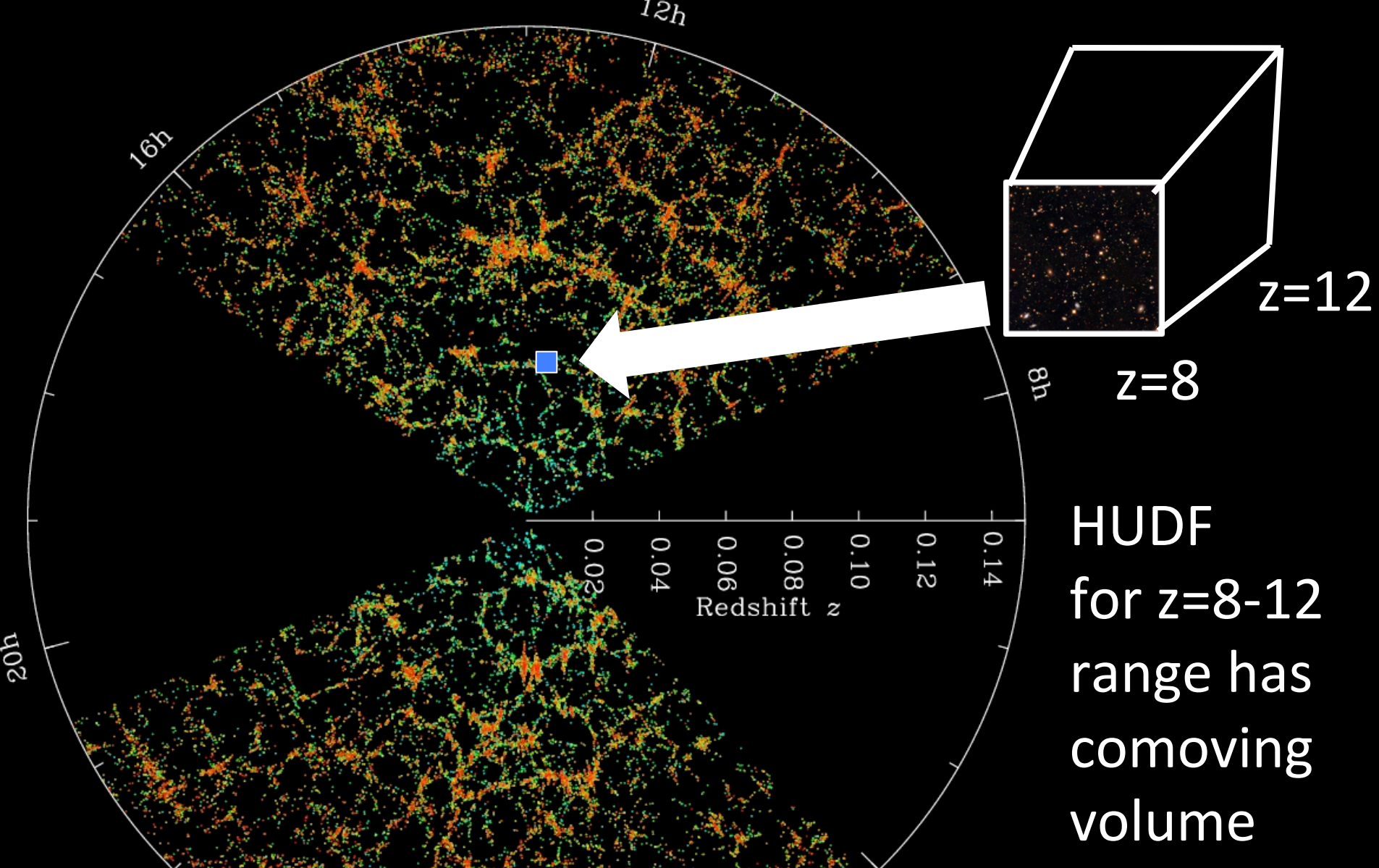


The first 600 million years ( $z=7+$ ):

Is this what the first galaxies are like?

Problem 1:





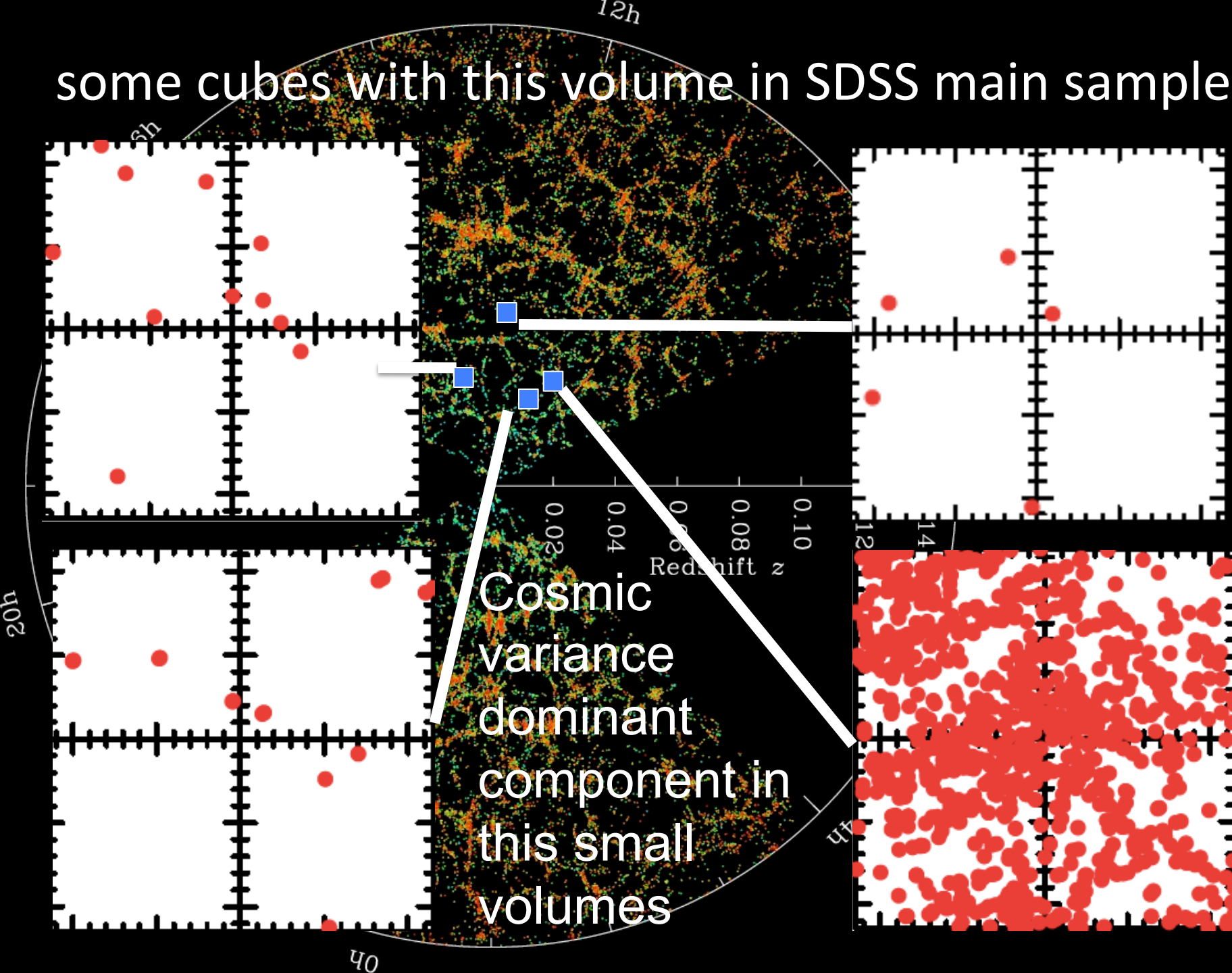
**1 parsec = 3.26 light years**

- 1 light year = distance a photon of light can travel in a year – about 5.68 **trillion** miles!!

HUDF  
for  $z=8-12$   
range has  
comoving  
volume  
 $(23 \text{ Mpc}/h)^3$



some cubes with this volume in SDSS main sample:





# First 600 million years: The first and only 1 quasar ...

LETTER TO NATURE

At  $z=7$ , 1 in a  $\text{Gpc}^3$

## A luminous quasar at a redshift of $z = 7.085$

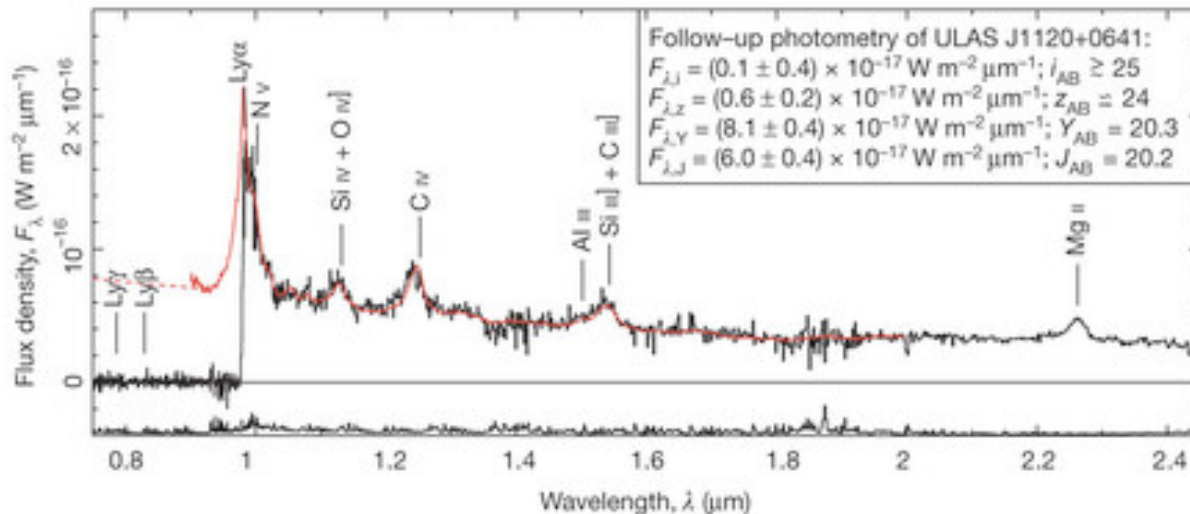
Daniel J. Mortlock<sup>1</sup>, Stephen J. Warren<sup>1</sup>, Bram P. Venemans<sup>2</sup>, Mitesh Patel<sup>1</sup>, Paul C. Hewett<sup>3</sup>, Richard G. McMahon<sup>3</sup>, Chris Simpson<sup>4</sup>, Tom Theuns<sup>5,6</sup>, Eduardo A. González-Solares<sup>3</sup>, Andy Adamson<sup>7</sup>, Simon Dye<sup>8</sup>, Nigel C. Hambly<sup>9</sup>, Paul Hirst<sup>10</sup>, Mike J. Irwin<sup>3</sup>, Ernst Kuiper<sup>11</sup>, Andy Lawrence<sup>9</sup> & Huub J. A. Röttgering<sup>11</sup>

$M_{\text{BH}} = 10^9 M_{\text{sun}}$

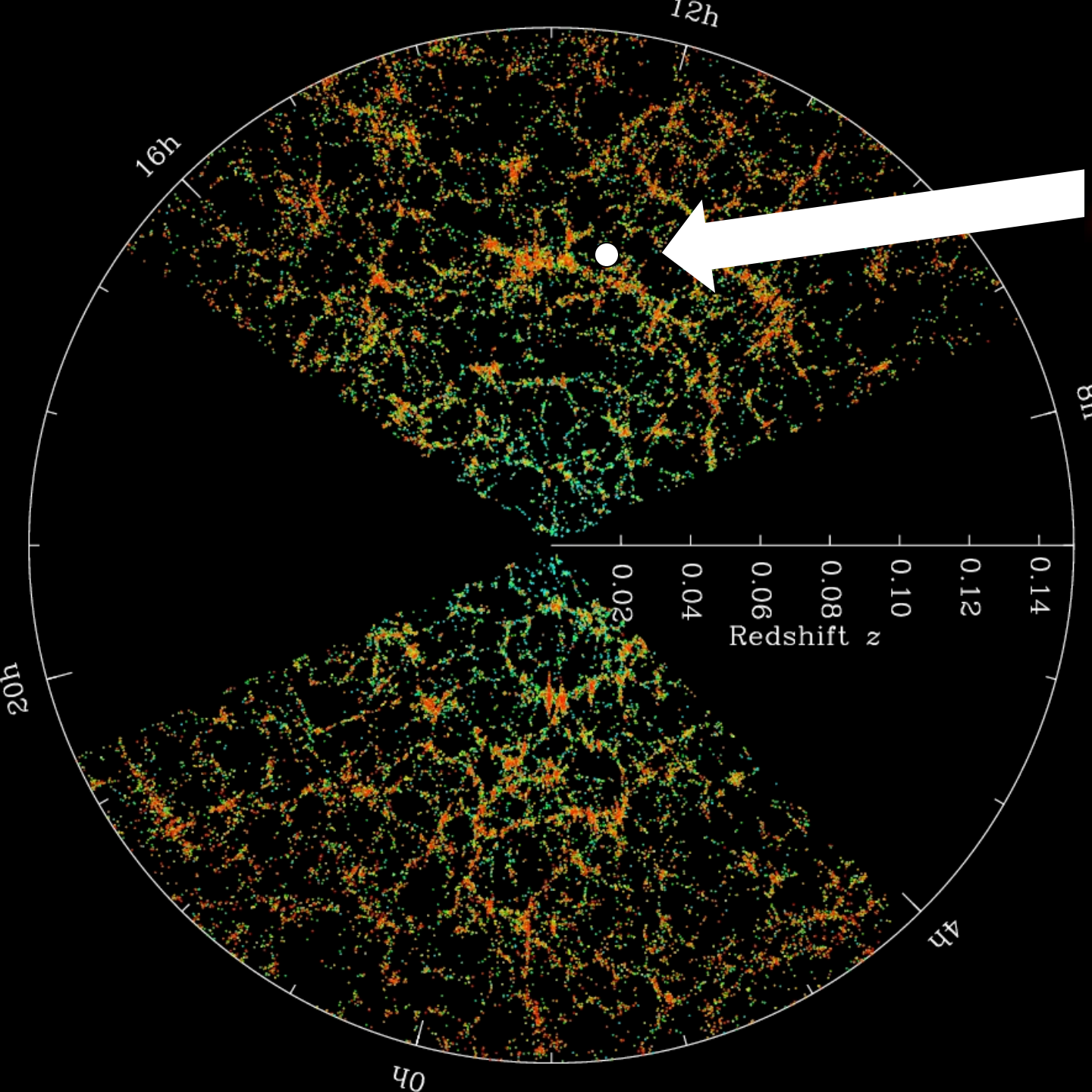
30 Jun 2011

The intergalactic medium was not completely reionized until approximately a billion years after the Big Bang, as revealed<sup>[1]</sup> by observations of quasars with redshifts of less than 6.5. It has been difficult to probe to higher redshifts, however, because quasars have historically been identified<sup>[2-5]</sup> in optical surveys, which are insensitive to sources at redshifts exceeding 6.5. Here we report observations of a quasar (ULAS J112001.48+064124.3) at a red-

shift of  $z = 7.085$ , spectroscopically confirmed to have even higher redshifts, two are faint  $J_{\text{AB}} \geq 26$  galaxies<sup>[10,11]</sup> and the other is a  $\gamma$ -ray burst which has since faded<sup>[12]</sup>. Indeed, it has not been possible to obtain high signal-to-noise ratio spectroscopy of any sources beyond the most distant quasars previously known: CFHQS J0210-0456<sup>[13]</sup> ( $z = 6.44$ ), SDSS 1148+5251<sup>[8]</sup> ( $z = 6.42$ ) and CFHQS J2329+0301<sup>[14]</sup> ( $z = 6.42$ ). Follow-up measurements of ULAS J1120+0641 will provide the first opportunity to explore the 0.1 Gyr between  $z = 7.08$  and  $z = 6.44$ .







1 quasar in  
this volume



The first 600 million years ( $z=7+$ ):

Do we expect such extreme mass BH?

Problem 2:





# How do the first massive black holes grow

## 1. BH Seed:



Pop III stellar  
Remnant

Gas/stellar Dynamical  
collapse



$$M_{\text{BH,seed}} = 100 - 10^5 M_{\text{sun}}$$

High gas density

THEORY

# How do MBH seeds grow at early time?

$z=6-7$  quasars imply  
 $M_{\text{BH}} = 10^9 M_{\text{sun}} =$   
= mass of most  
massive  
black holes today

First billion years  
requires extremely  
large accretion rates

$$L_{\text{Edd}} = \frac{4\pi G c m_p}{\sigma_T} M_{\text{BH}} = \epsilon \dot{M} c^2$$

$$M_{\text{BH}} = M_{\text{seed}} e^{\frac{t}{t_{\text{Edd}}}}$$

$$t_{\text{Edd}} = 450 \text{ Myr} \frac{\epsilon}{1 - \epsilon}; \epsilon = 0.1$$

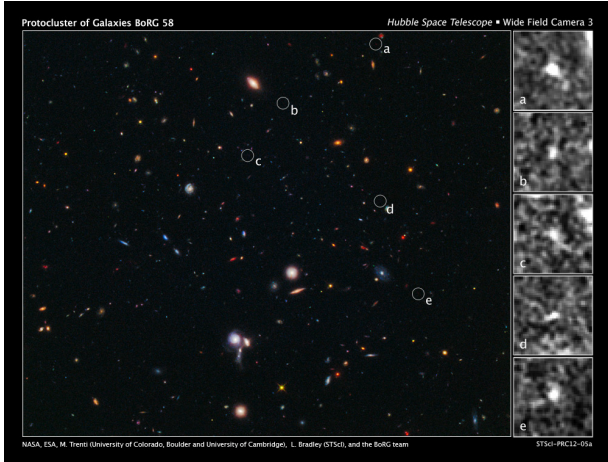
$$\begin{aligned} \ln(M_{\text{BH}}/M_{\text{seed}}) &= \ln[10^9 / (100 - 1e5)] \\ &= 10 - 17 \text{ e-foldings} \end{aligned}$$

Is it feasible?



# At $z > 7$ :

## Plenty of (‘smudgy’) galaxies in tiny volumes



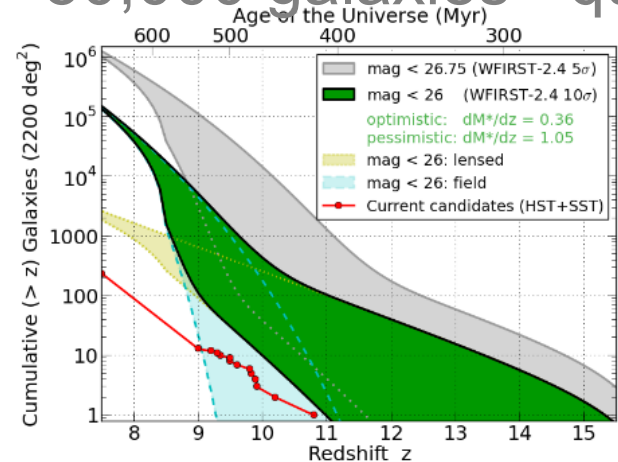
Hubble Legacy Fields



Room for discovery...

## WFIRST

> 50,000 galaxies + quasars



JWST spectra etc..

## 1 supermassive quasar

LETTER TO NATURE

A luminous quasar at a redshift of  $z = 7.085$

Daniel J. Mortlock<sup>1</sup>, Stephen J. Warren<sup>1</sup>, Bram P. Venemans<sup>2</sup>, Mitesh Patel<sup>1</sup>, Paul C. Hewett<sup>3</sup>, Richard G. McMahon<sup>3</sup>, Chris Simpson<sup>4</sup>, Tom Theuns<sup>5,6</sup>, Eduardo A. González-Solares<sup>3</sup>, Andy Adamson<sup>7</sup>, Simon Dye<sup>8</sup>, Nigel C. Hambly<sup>9</sup>, Paul Hirst<sup>10</sup>, Mike J. Irwin<sup>3</sup>, Ernst Kuiper<sup>11</sup>, Andy Lawrence<sup>9</sup> & Huub J. A. Röttgering<sup>11</sup>

The intergalactic medium was not completely re-ionized until approximately a billion years after the Big Bang, as revealed<sup>1</sup> by observations of quasars with redshifts of less than 6.5. It has been difficult to probe to higher redshifts, however, because quasars have historically been identified<sup>2,3,4</sup> in optical surveys, which are insensitive to sources at redshifts exceeding 6.5. Here we report observations of a quasar (ULAS J112001.48+064124.3) at a redshift of 7.085, which is 0.77 billion years after the Big Bang. ULAS J1120+0641 had been previously

trioscopically confirmed to have even higher redshifts, two are faint  $J_{AB} \geq 26$  galaxies<sup>10,11</sup> and the other is a  $\gamma$ -ray burst which has since faded<sup>12</sup>. Indeed, it has not been possible to obtain high signal-to-noise ratio spectroscopy of any sources beyond the most distant quasars previously known: CFHQS J0210-0456<sup>13</sup> ( $z = 6.44$ ), SDSS J148+5251<sup>14</sup> ( $z = 6.42$ ) and CFHQS J2329+0301<sup>15</sup> ( $z = 6.42$ ). Follow-up measurements of ULAS J1120+0641 will provide the first opportunity to explore the 0.1 Gyr between  $z = 7.08$  and  $z = 6.44$ , a significant cosmological epoch about which little is cur-

01 30 Jun 2011

The first 600 million years ( $z=7+$ ):

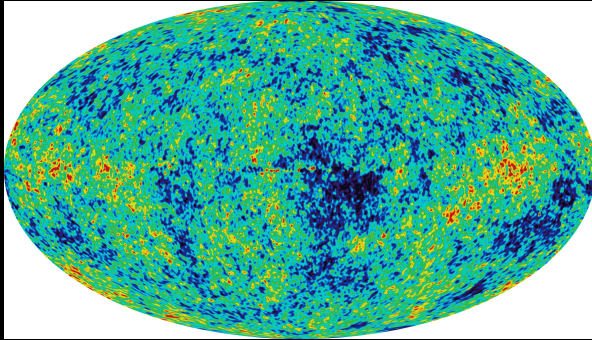
- ② What are galaxies like?
- ② How do the first massive black hole grow?

**Predictions:**

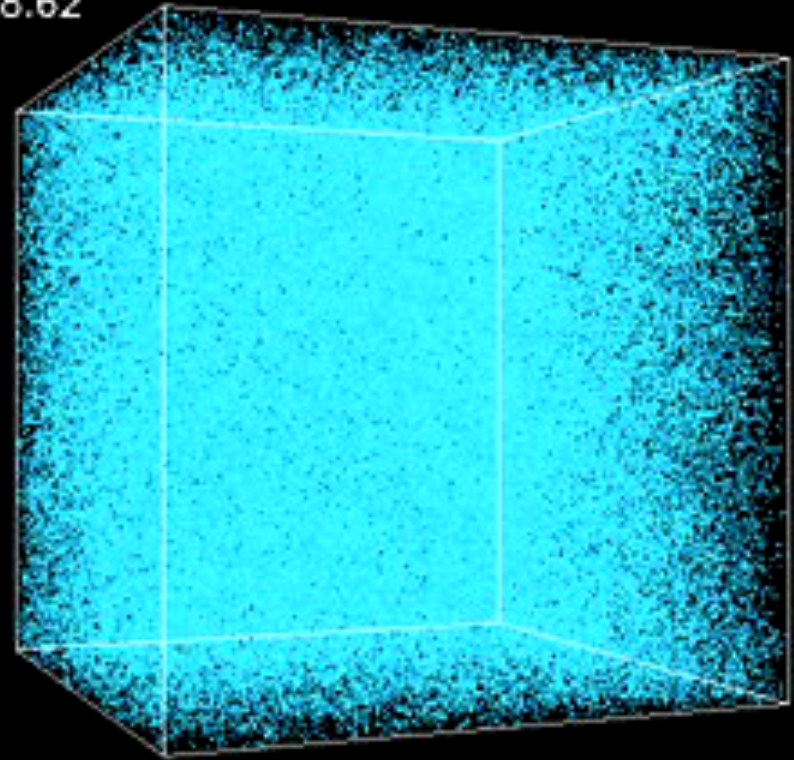


# Simulations of structure formation in the Universe

# The ICs

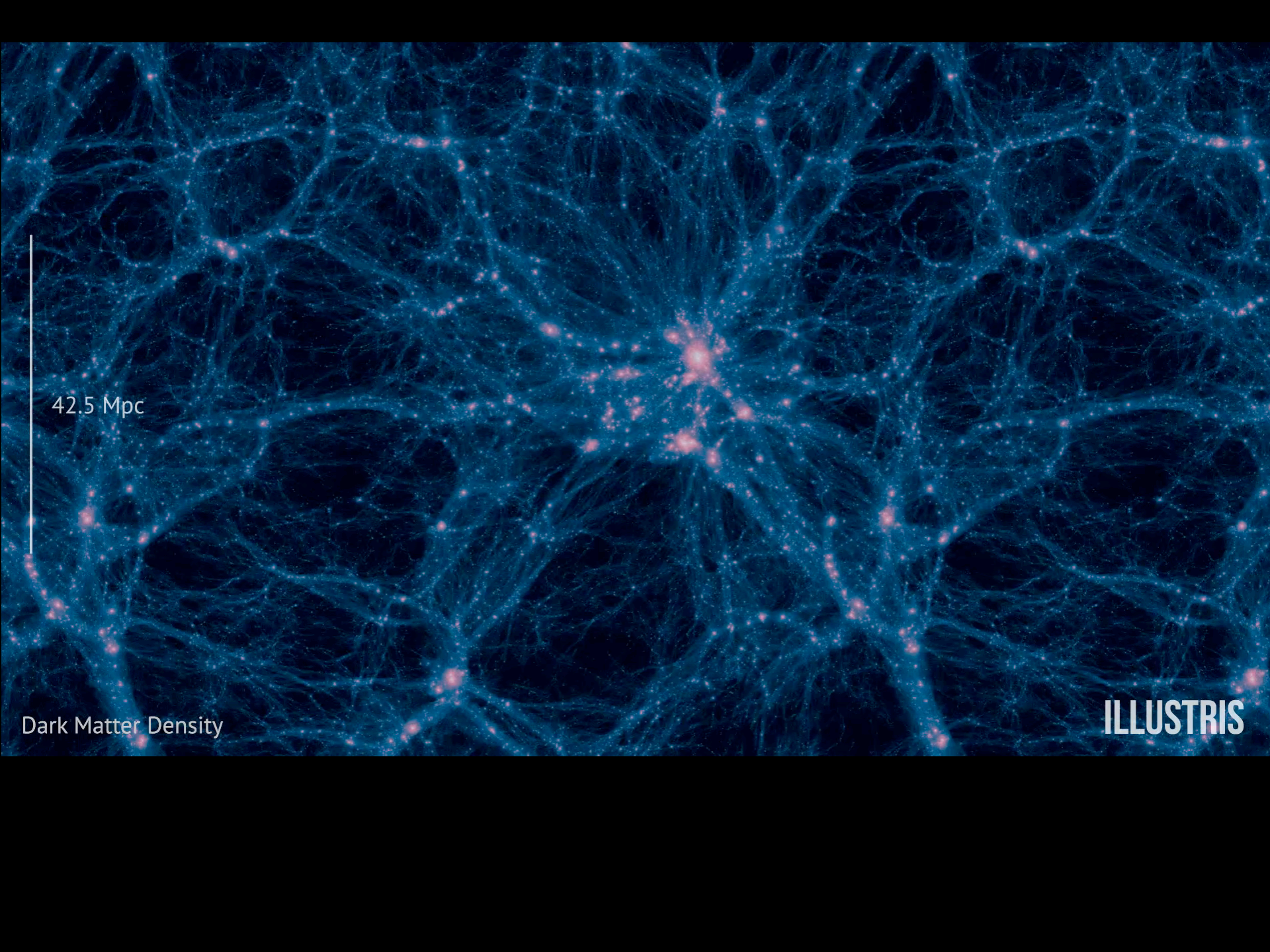


$z=28.62$



Gravitational instability





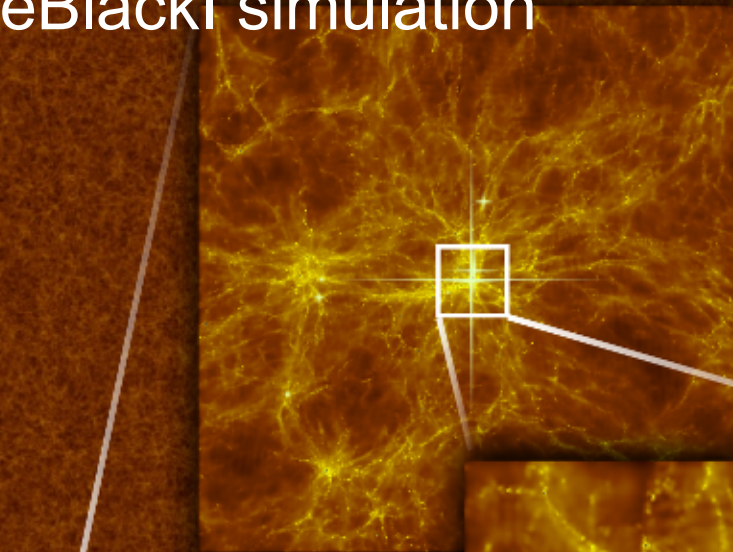
42.5 Mpc

Dark Matter Density

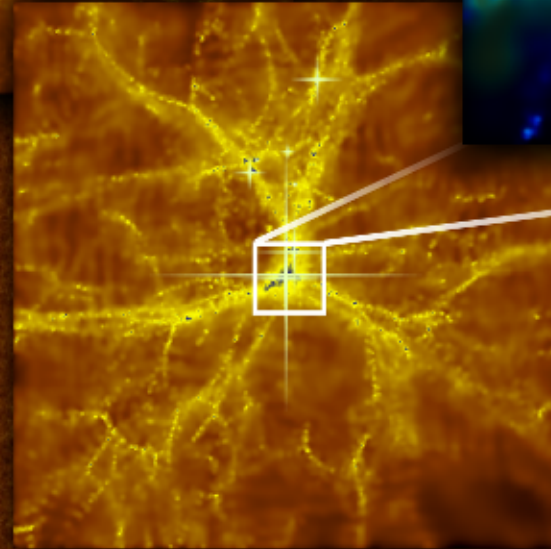
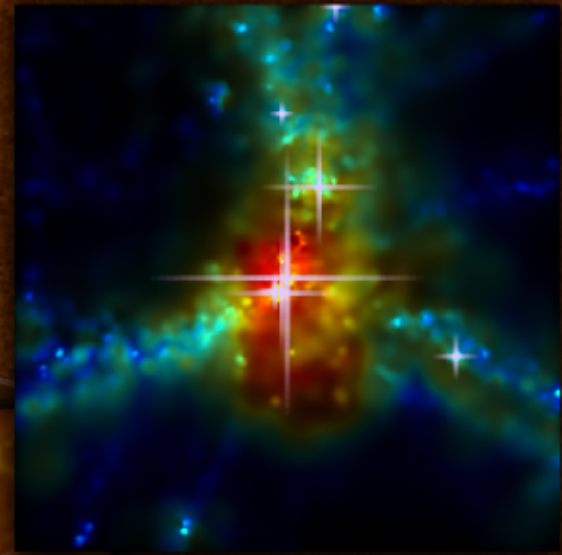
**ILLUSTRIS**



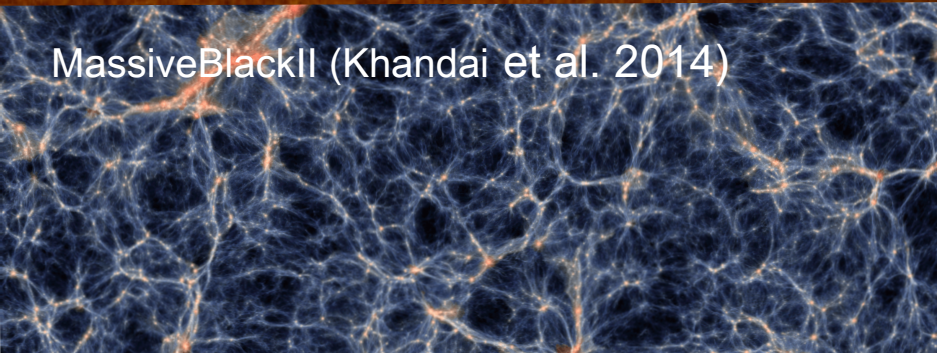
MassiveBlackI simulation



MassiveBlackI (DiMatteo et al. 2011)



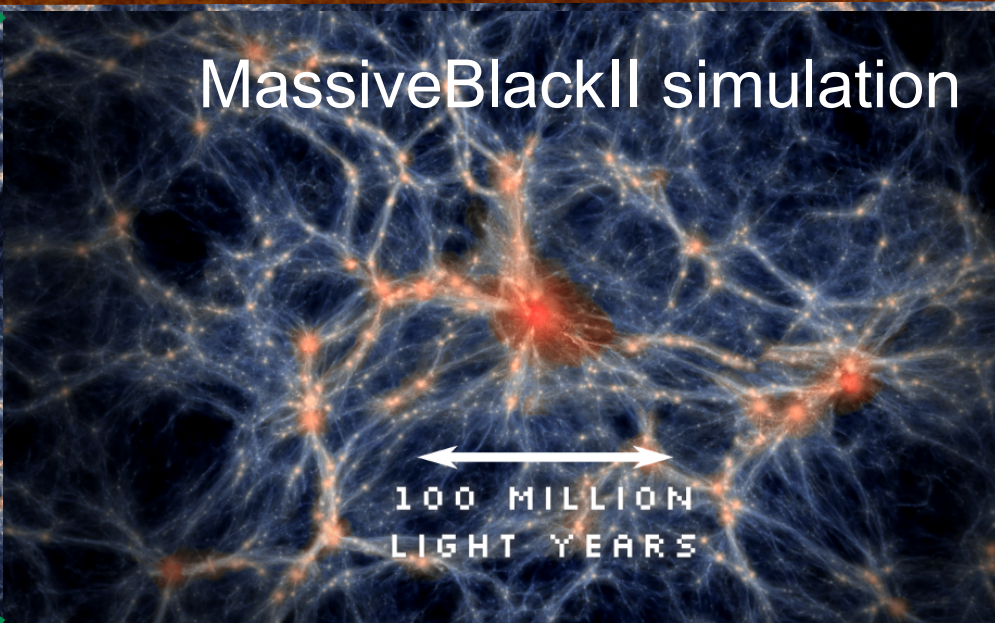
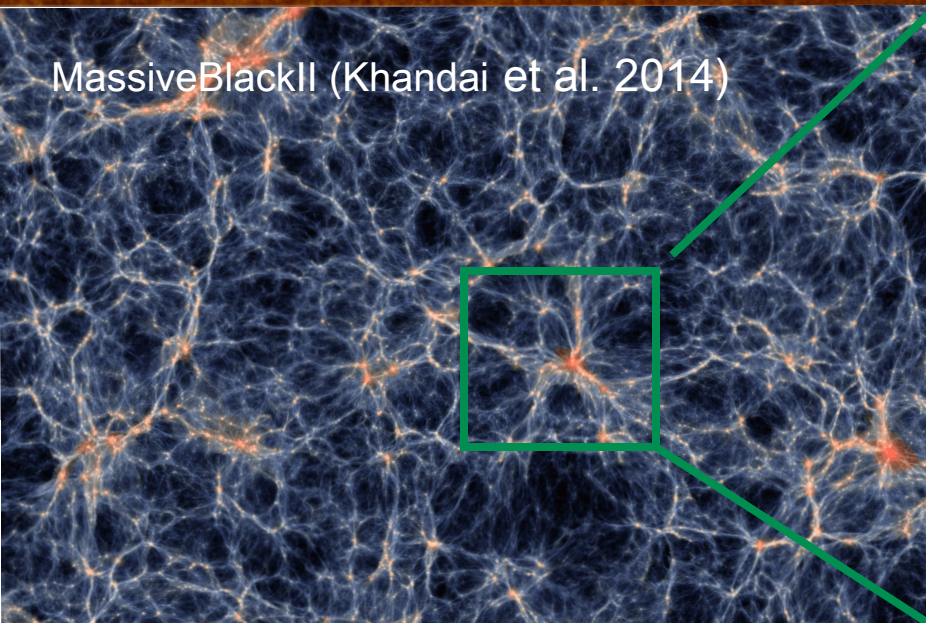
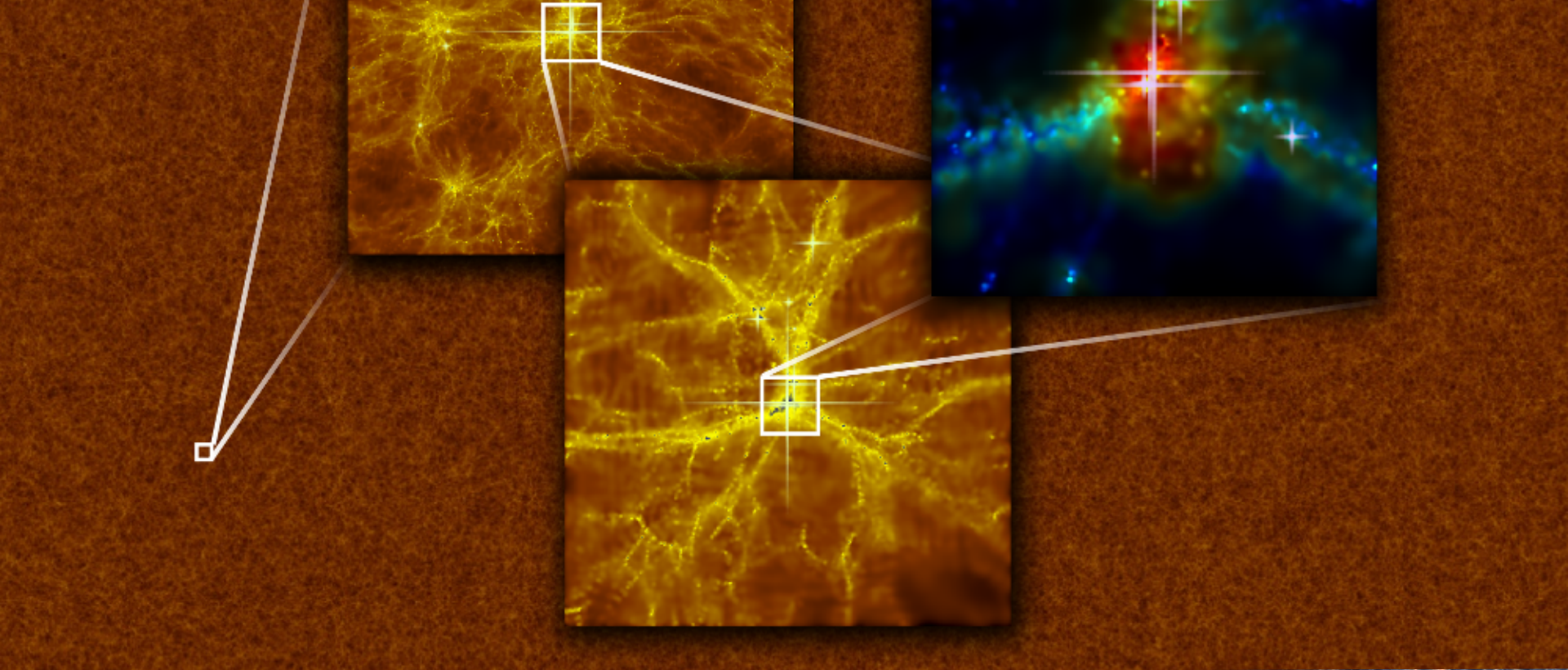
MassiveBlackII (Khandai et al. 2014)



MassiveBlackII simulation

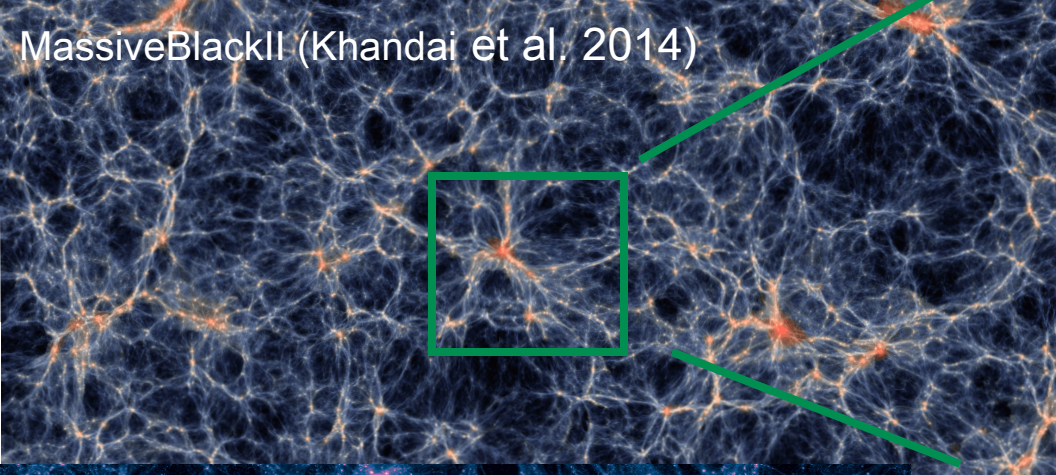




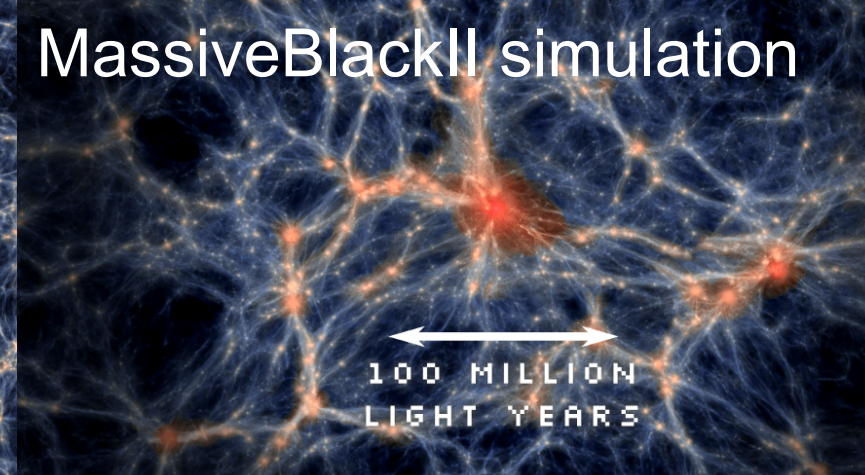




MassiveBlackII (Khandai et al, 2014)

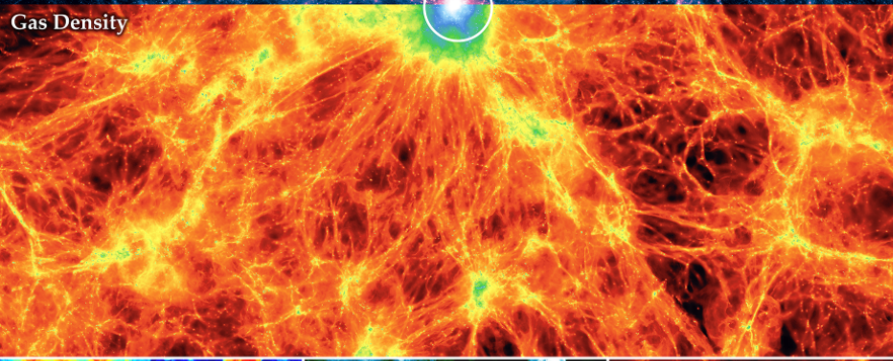
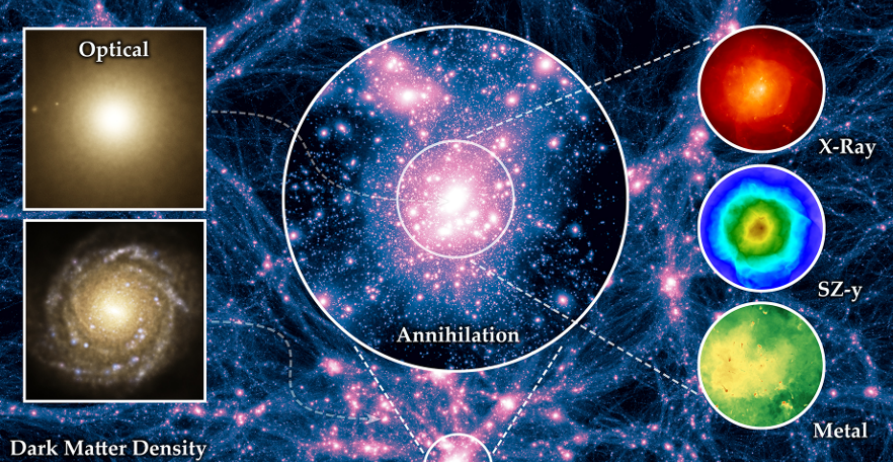


MassiveBlackII simulation

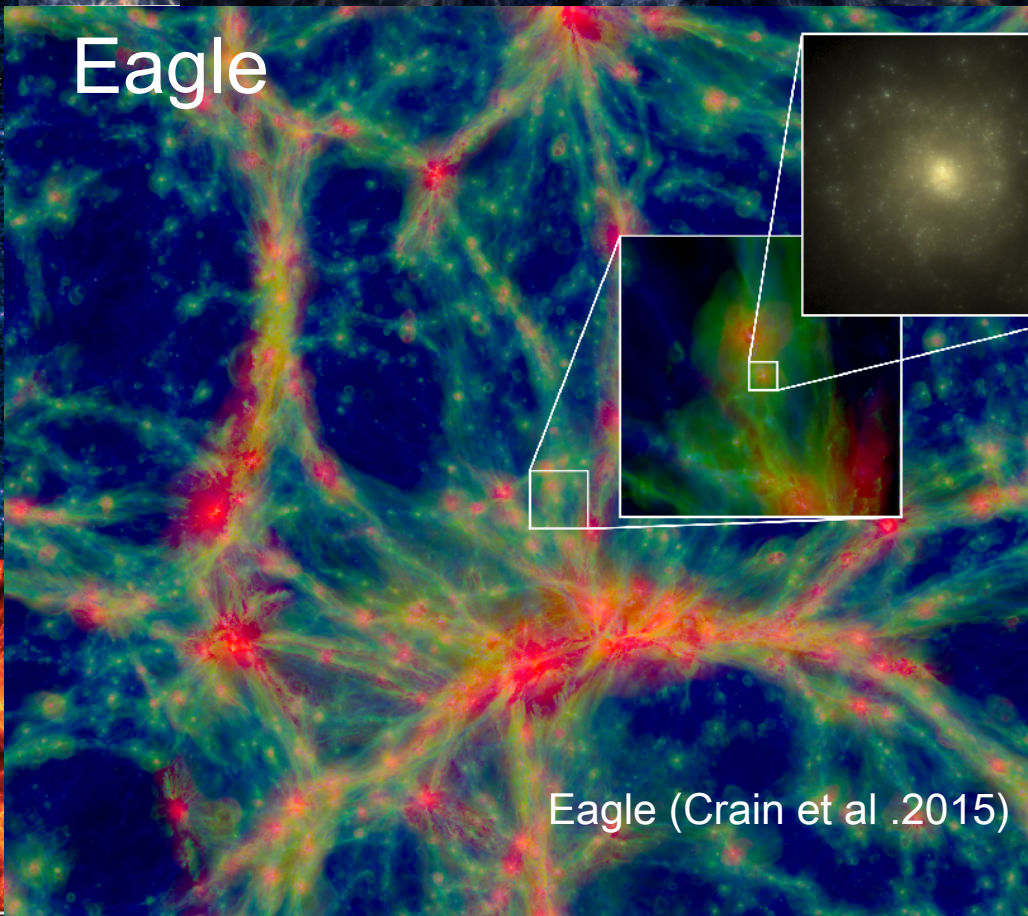


### The Illustris Simulation

M. Vogelsberger S. Genel V. Springel P. Torrey D. Sijacki D. Xu G. Snyder S. Bird D. Nelson L. Hernquist

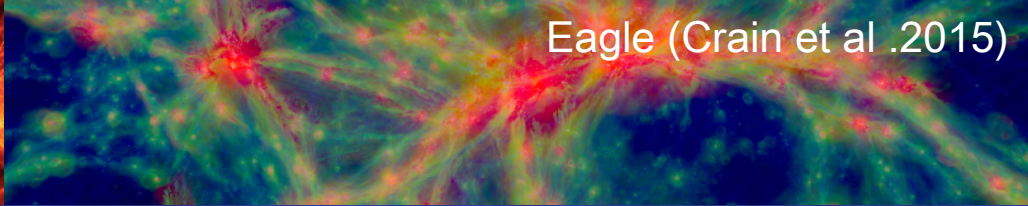
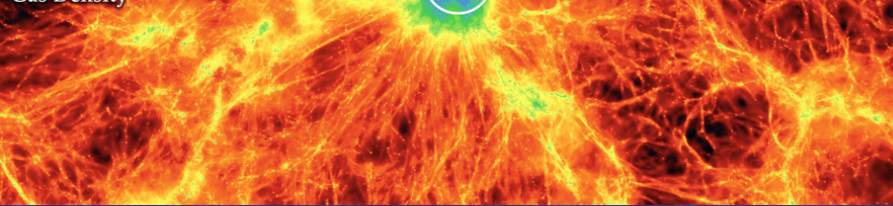


### Eagle

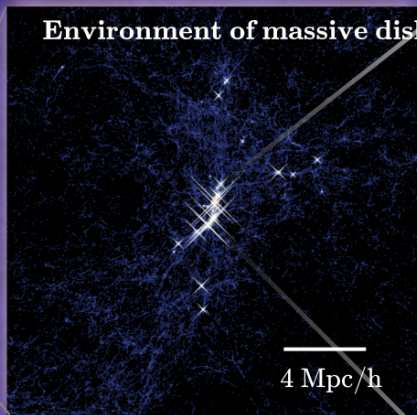


Eagle (Crain et al .2015)



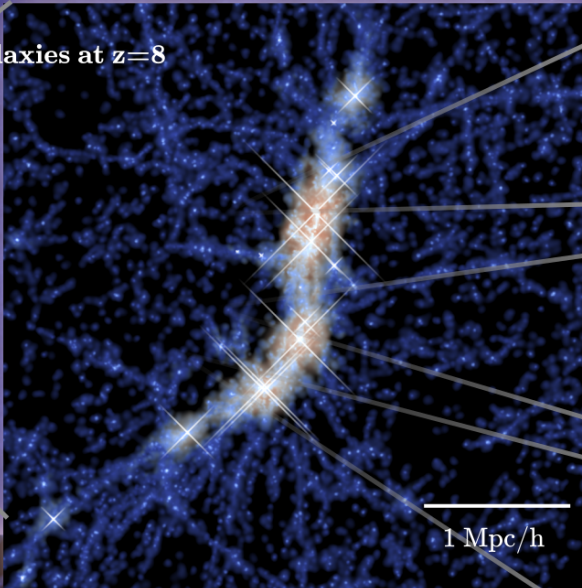


Eagle (Crain et al .2015)



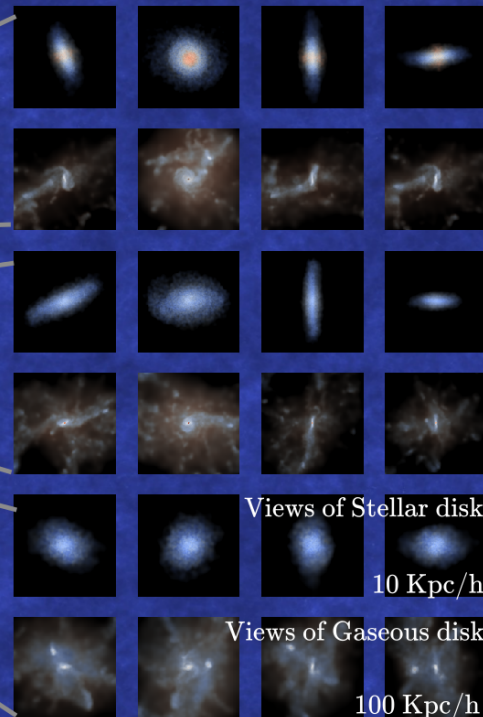
Environment of massive disk galaxies at  $z=8$

4 Mpc/h



Environment of most massive blackhole at  $z=8$

1 Mpc/h

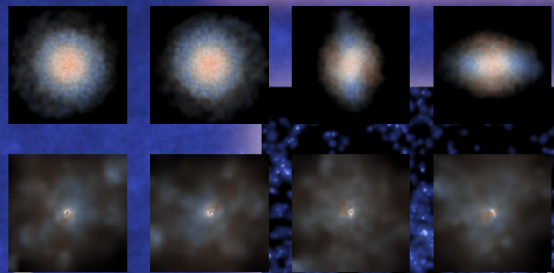


Views of Stellar disk

10 Kpc/h

Views of Gaseous disk

100 Kpc/h



### The **BlueTides** Simulation

0.7 trillion particles

0.65 million cores



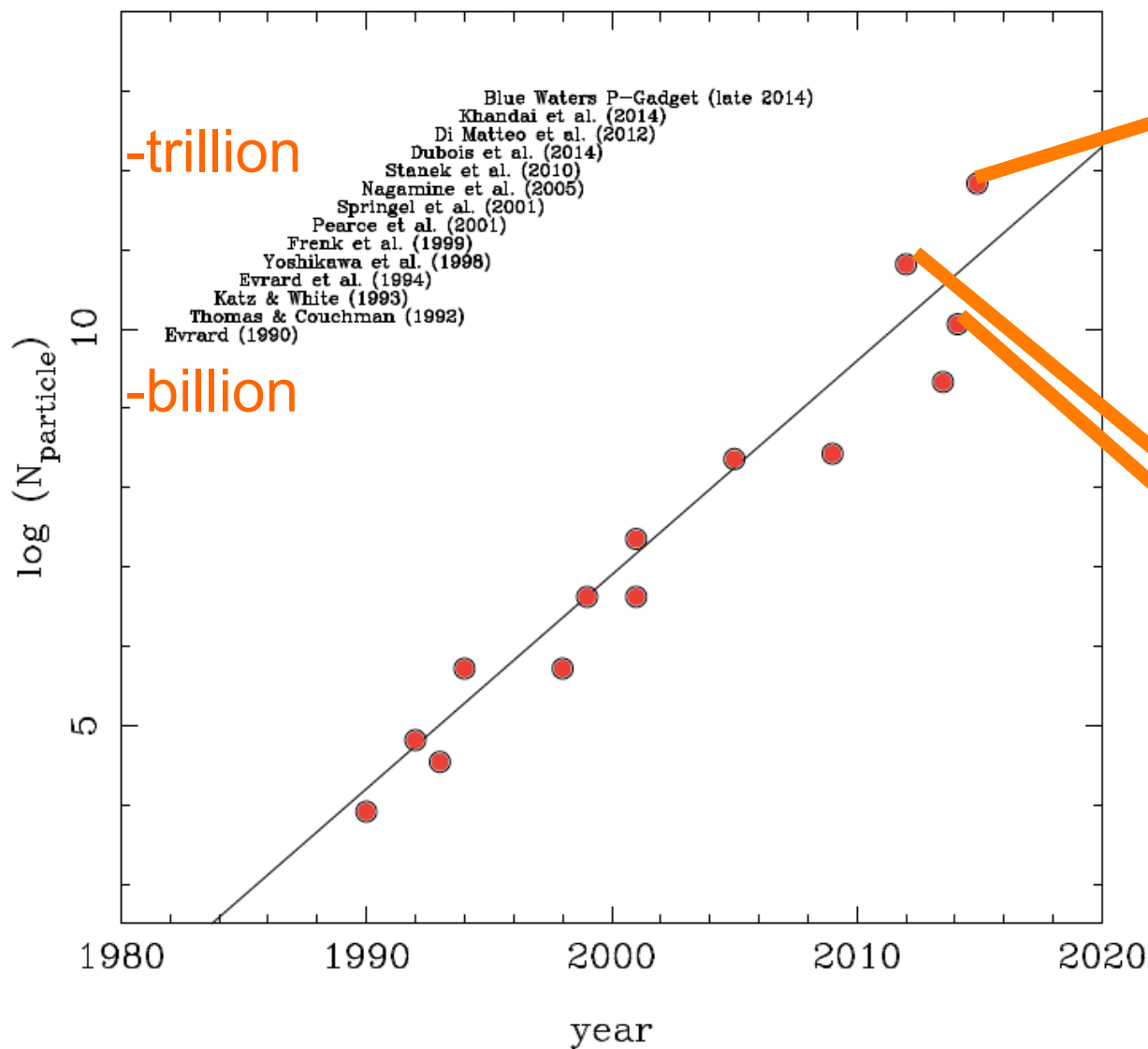
# bluetides

Feng et al. 2015

40 Mpc/h

# Algorithms keep up with computational power

## Hydro simulations:



On 0.72m cores  
NCSA Cray XE6  
Blue Waters



On 112k cores  
NICS Cray XT5  
Kraken



# We resolve galaxies across the full mass function

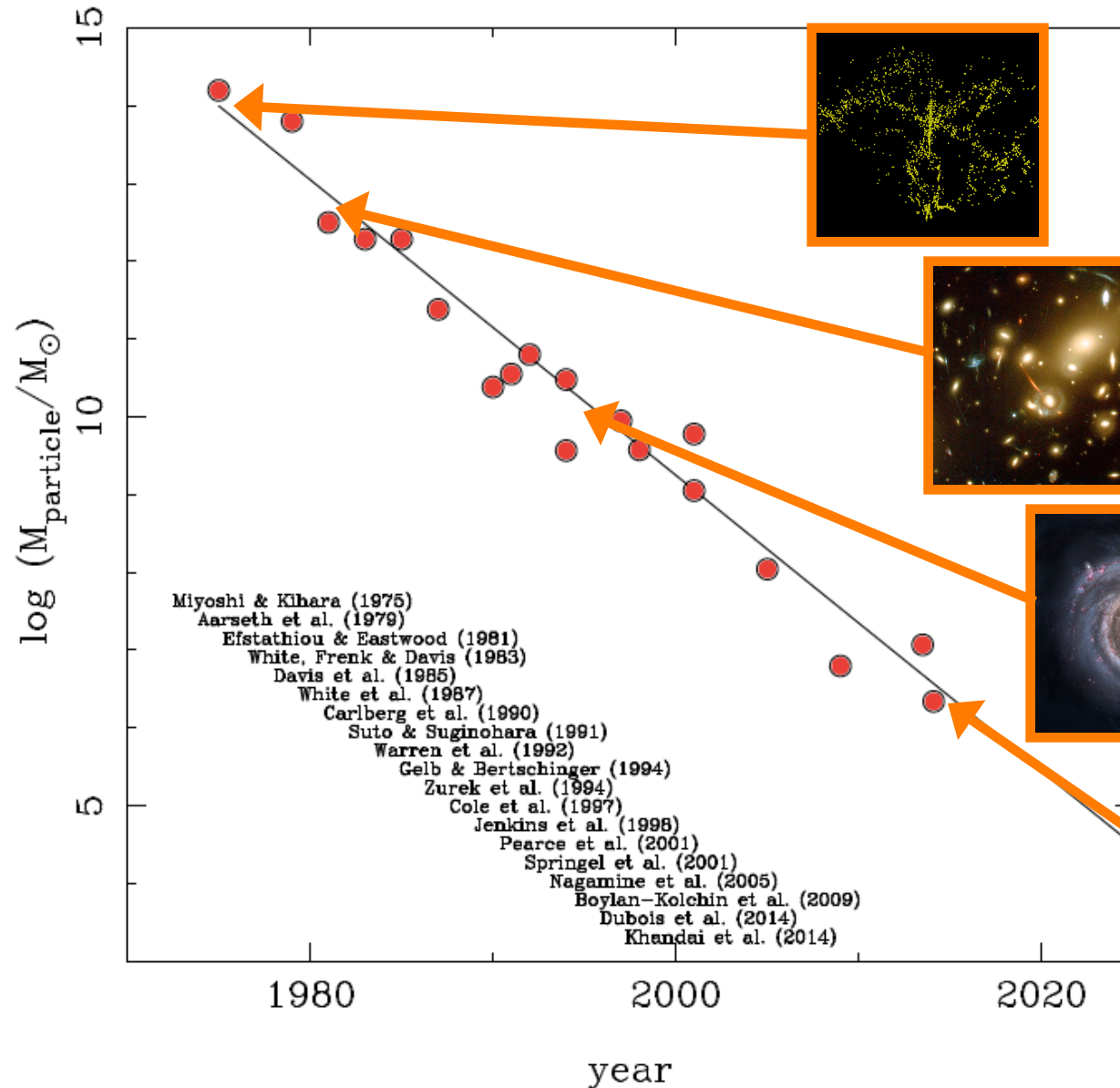
What we can resolve with 100 particles:

Superclusters of galaxies

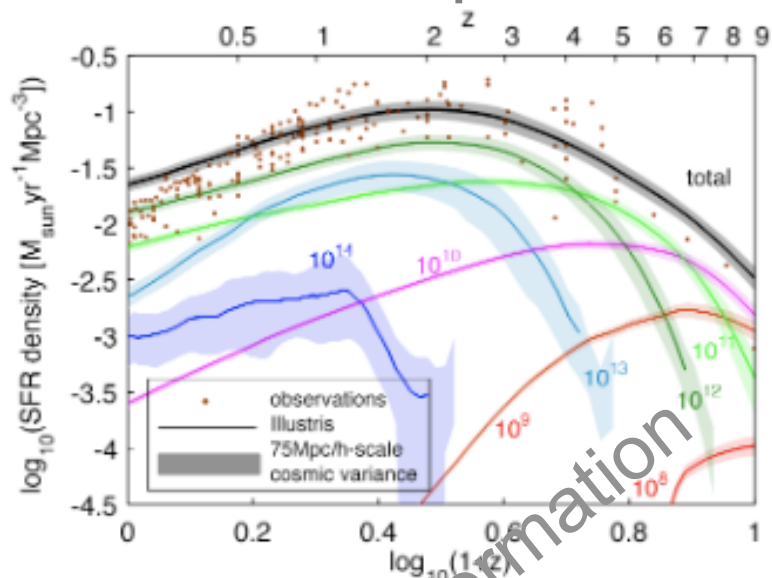
Clusters of galaxies

Milky way-sized galaxies

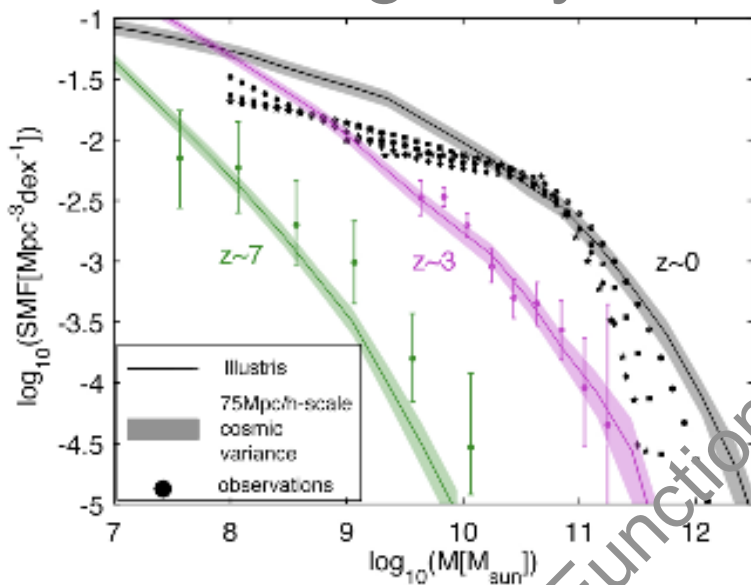
Dwarf galaxies



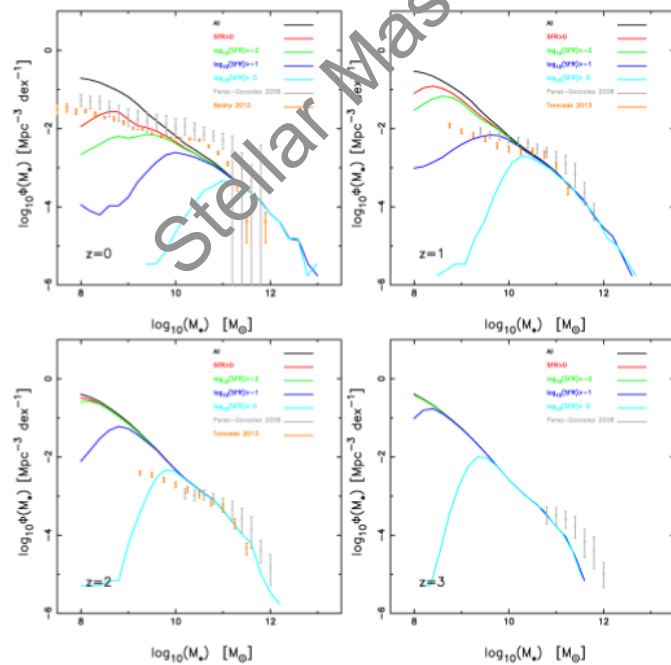
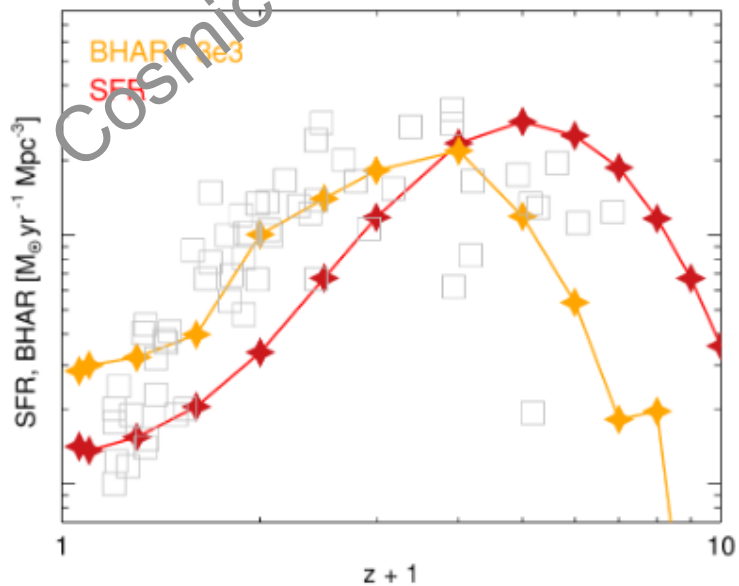
# Simulations reproduce statistics of galaxy formation



(c) Cosmic SFR density, 106.5 Mpc cosmic variance



(d) Stellar mass functions, 106.5 Mpc cosmic variance





The first 600 million years ( $z=6+$ ):

Can massive BH grow to billion solar mass?

## Problem 2:



Example:

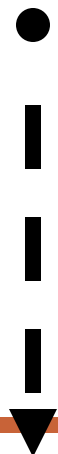
# How do the first massive black holes grow



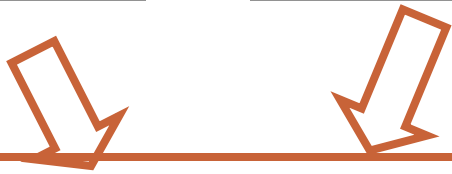
1. BH Seed:

Pop III stellar  
Remnant

Gas/stellar Dynamical  
collapse

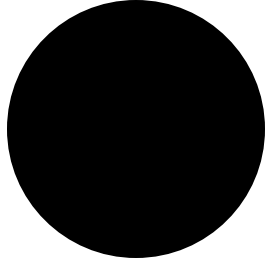


$t < 1$  billion yrs



High **gas** density

2. Gas accretion: Eddington (sustained)

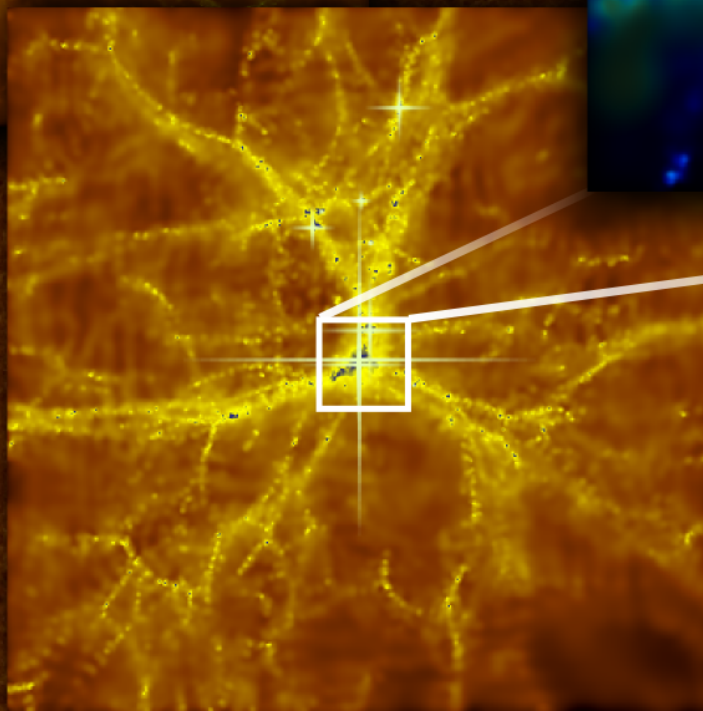
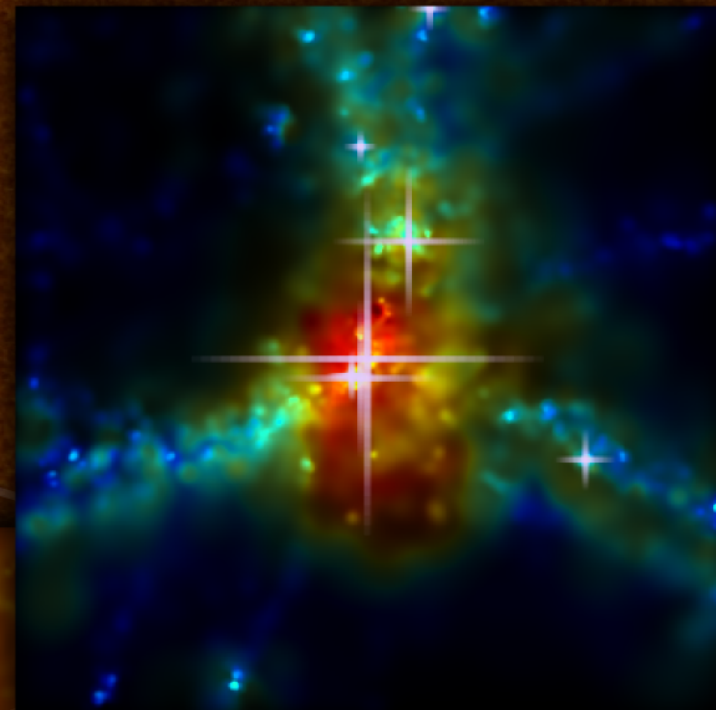
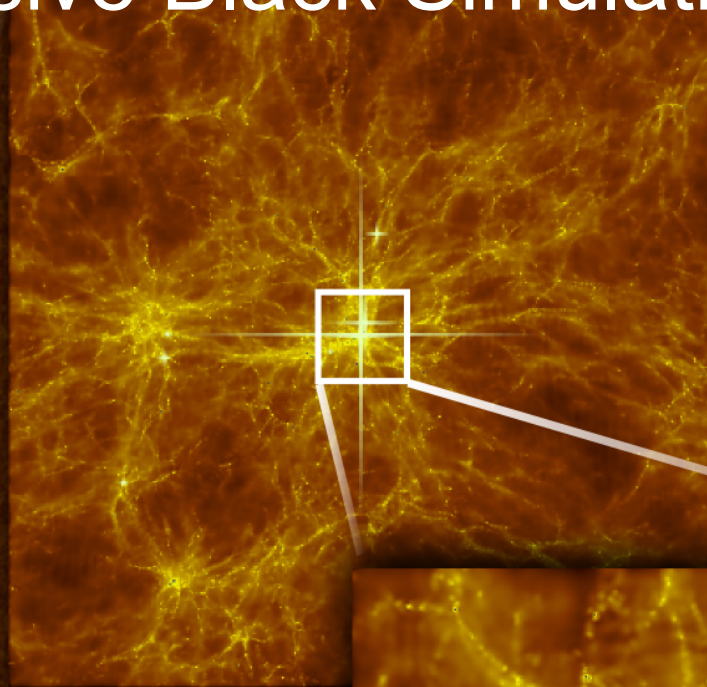


Highest **gas** inflow rates

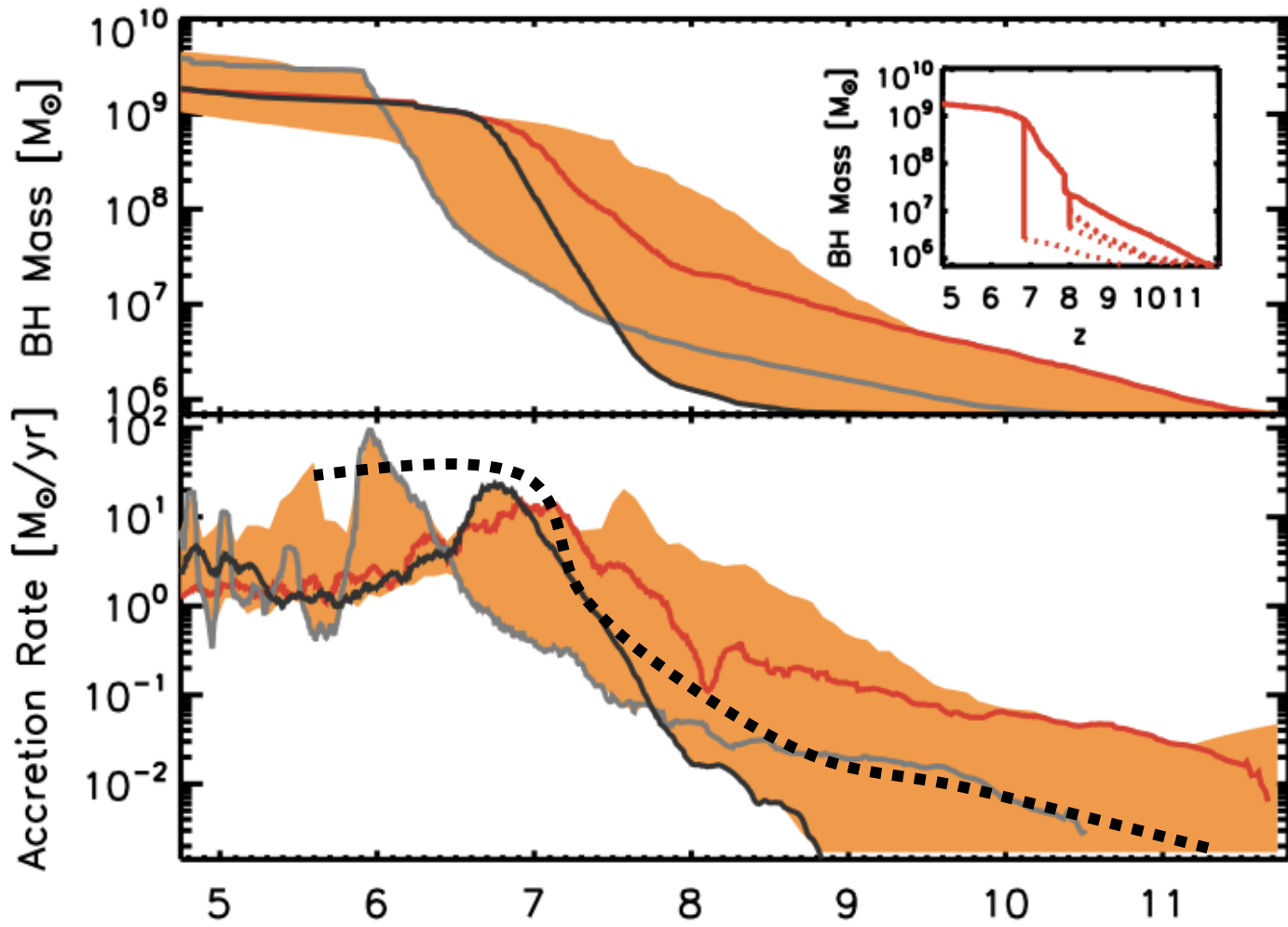
Feasible? if so how, where?



# Massive Black Simulation, Uniform $\sim 1 \text{ Gpc}^3 \text{ Vo}$



# Black holes grow to $10^9 M_{\text{sun}}$ by $z=7-6$



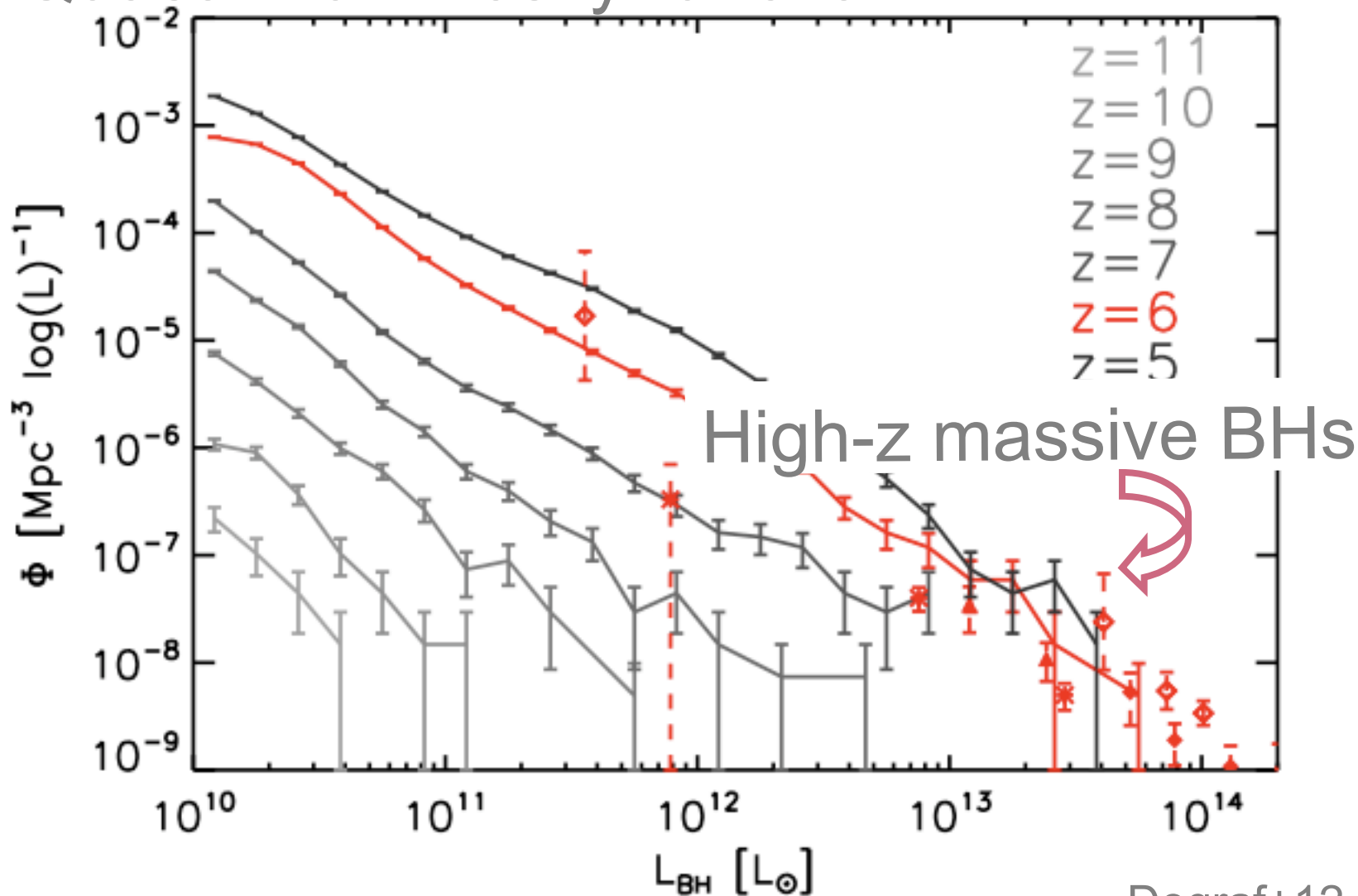
DM et al . 2012

Eddington rates sustained long enough before AGN feedback able to act

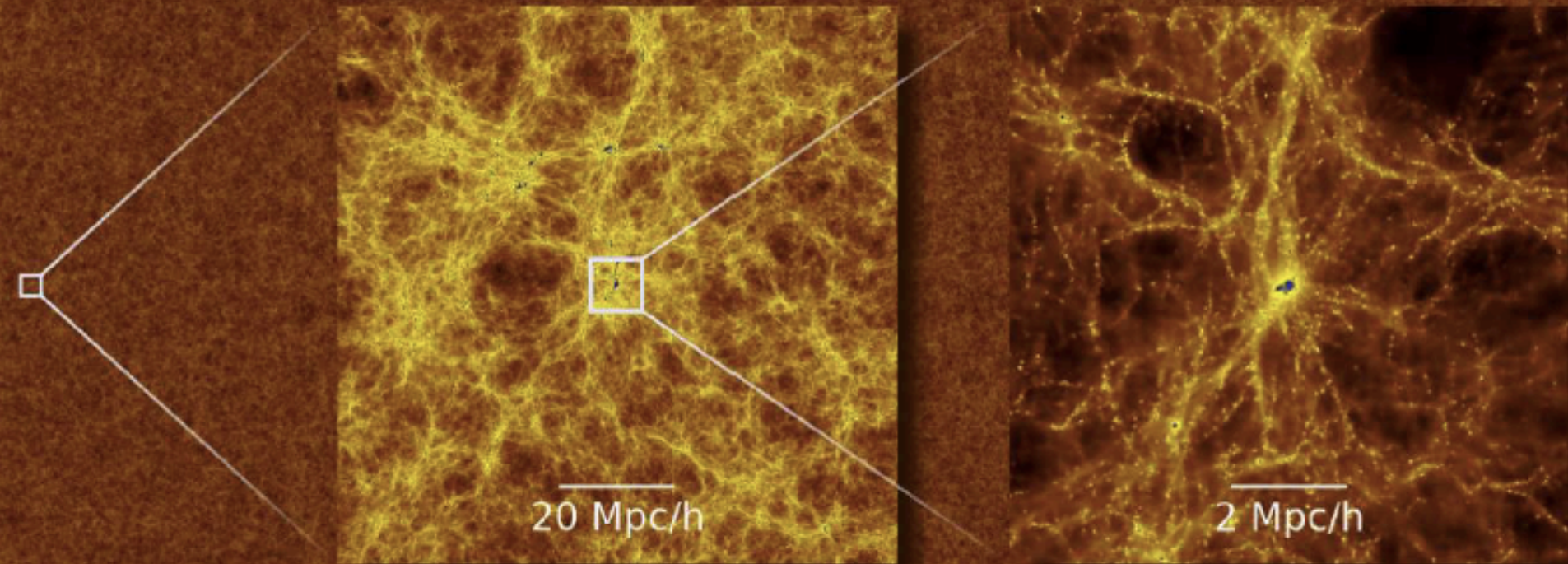


# MB cosmological simulation predicts the right number of Sloan Quasars

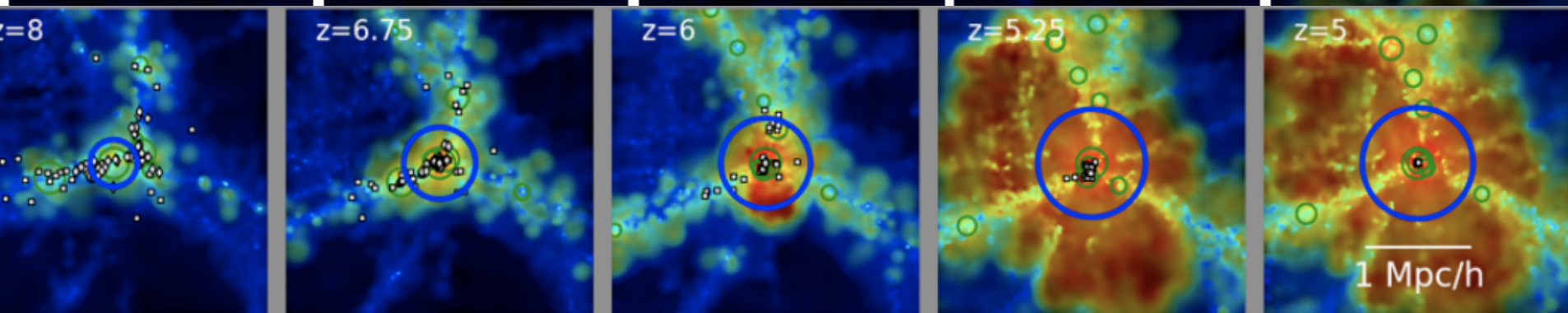
## High-z Quasar Luminosity Function



# Simulations show that in high density regions



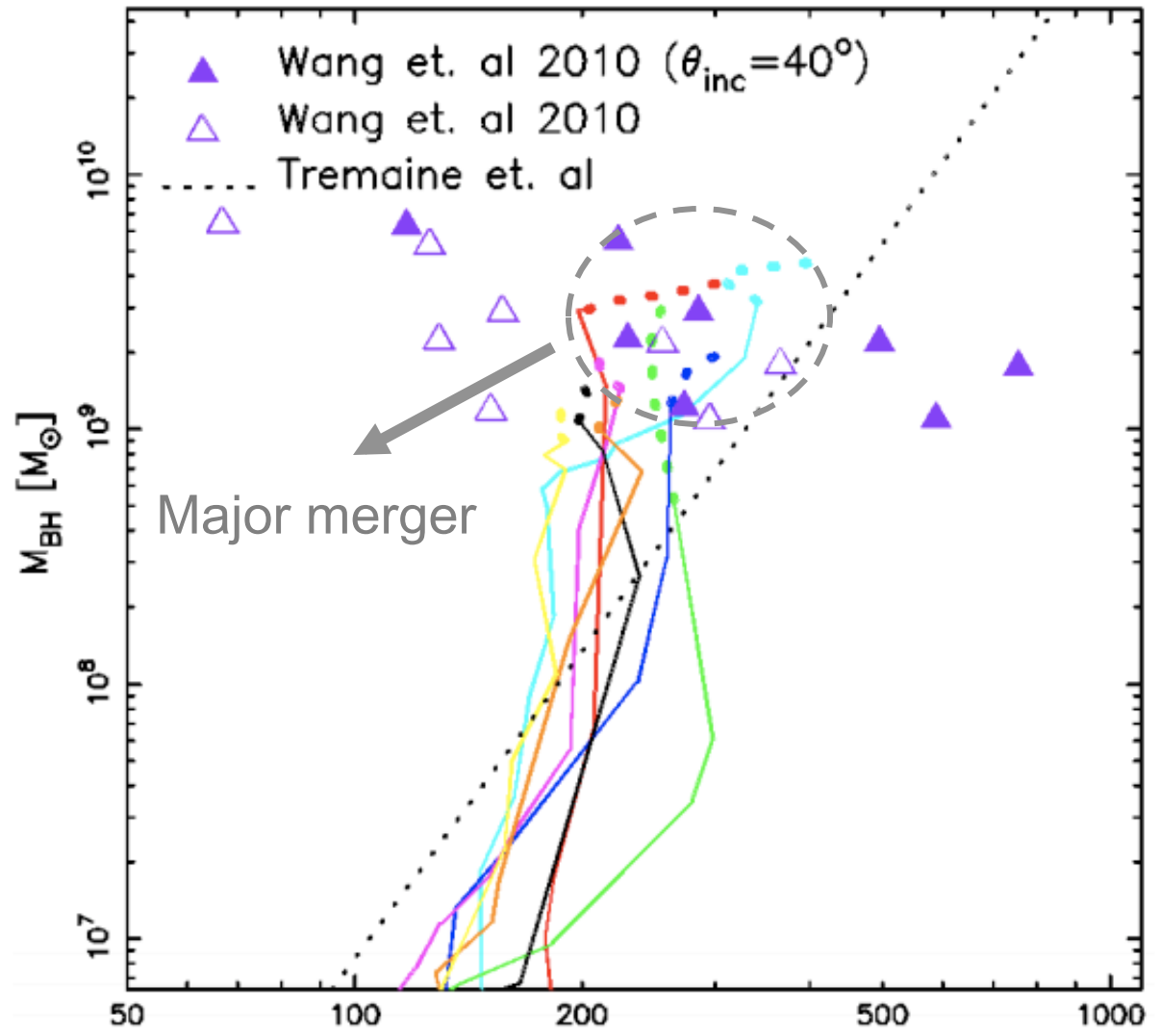
“Cold gas flows” stream to the heart of galaxies and grow black holes





# Z=6 Quasar Hosts: The $M_{\text{BH}}\text{-}\sigma$ relation

Khandai et al. 2012



**Black Holes grow first!**

$\sigma$  (km/s) Proxy for galaxy mass

However, theoretical predictions lacking at  $z=7+$  simulations have either: **insufficient resolution or too small volumes for massive objects**

MassiveBlack I (DiMatteo et al 2011) 533 Mpc/h & 5.5 Kpc/h

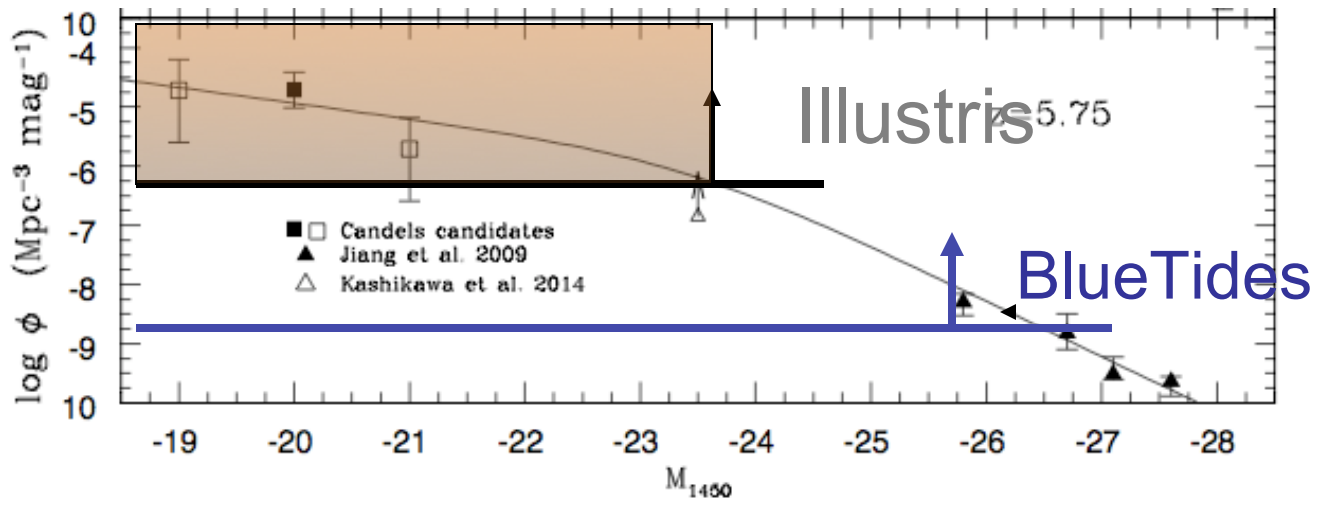
MassiveBlack II (Khandai et al 2014) 100 Mpc/h & 1.8 Kpc/h

100Mpc/h ~0.2 sq degrees at  $z=8-9$

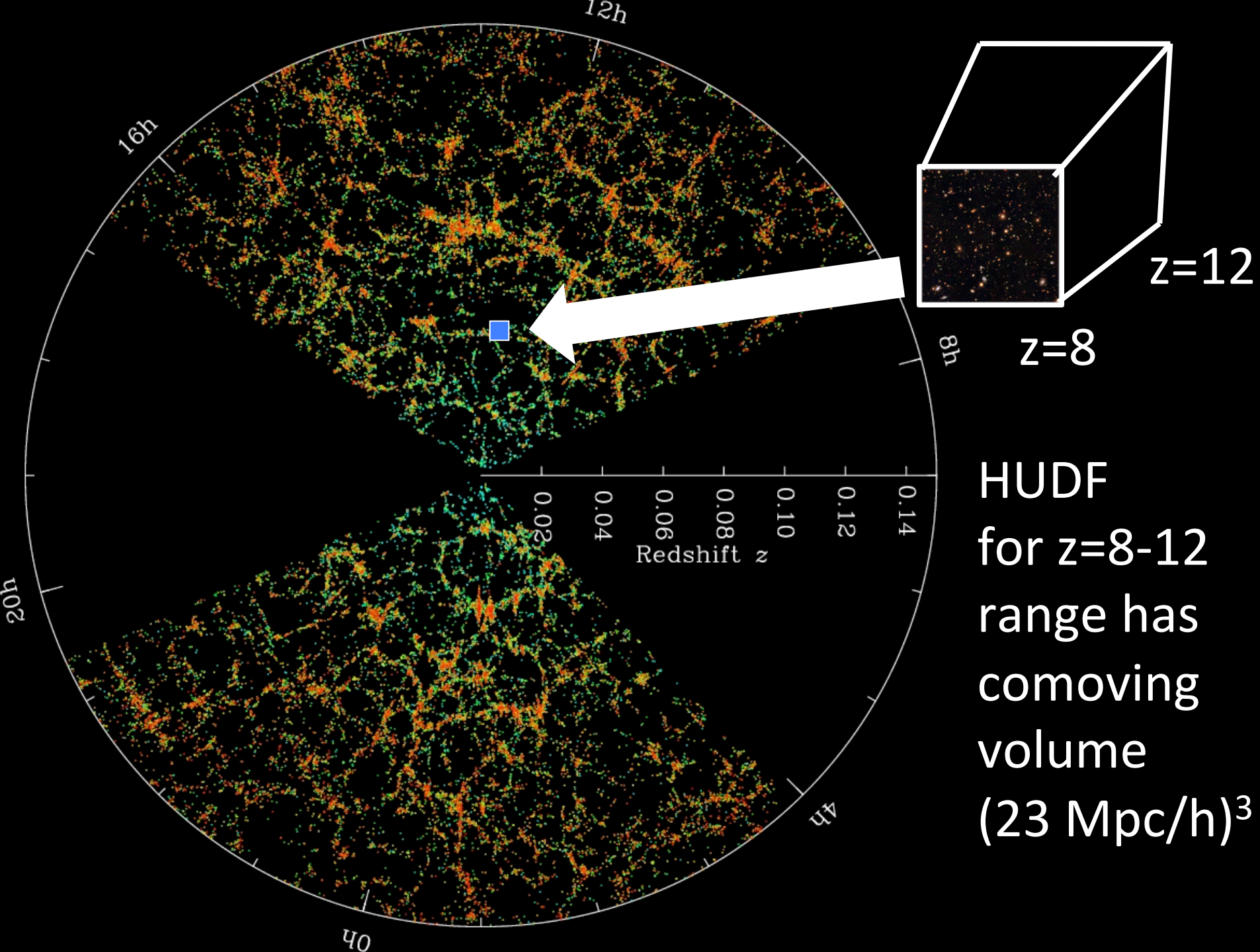
Illustris (Vogelsberger et al 2014) : 72.5 Mpc/h & 1 Kpc/h

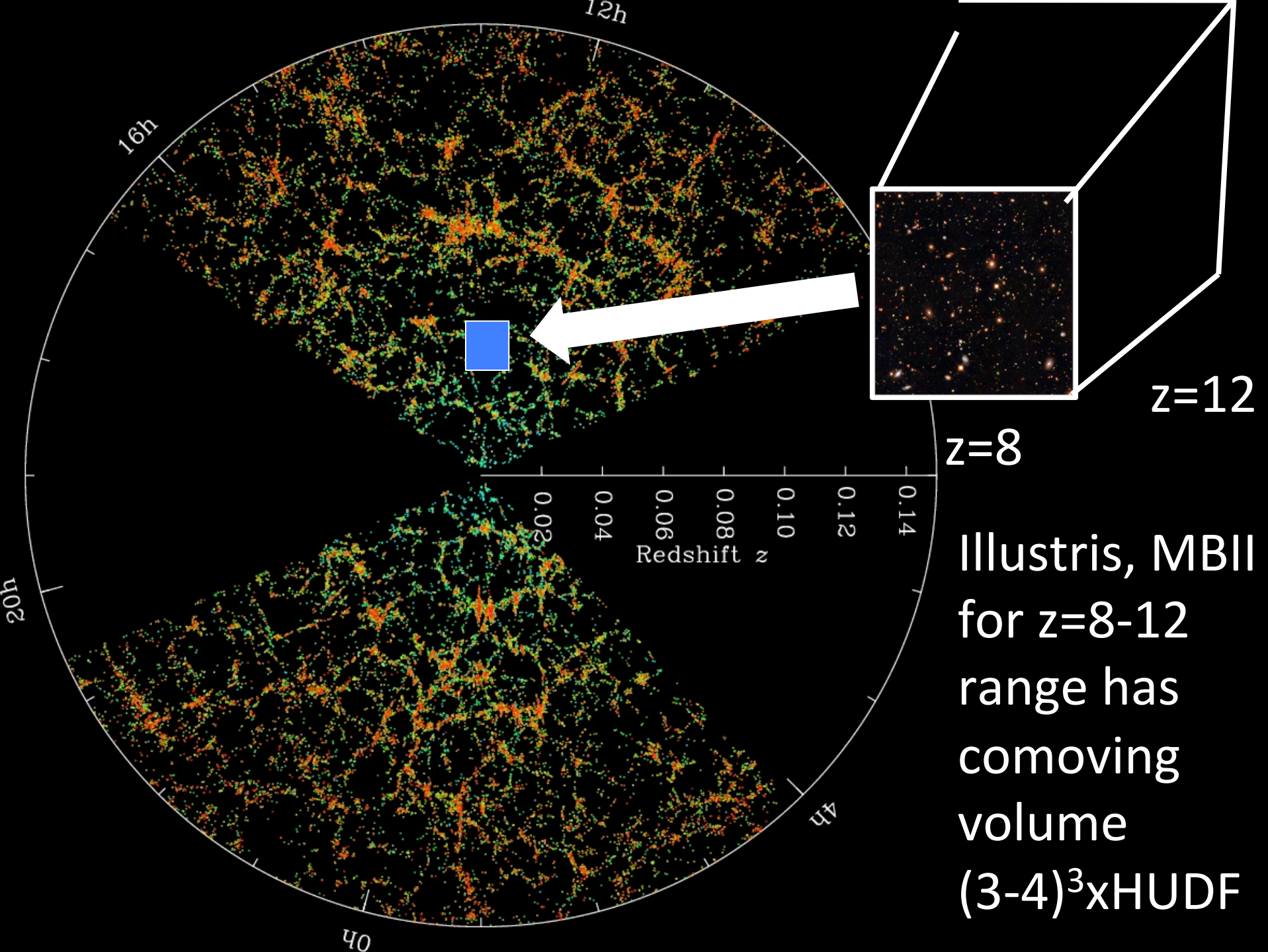
Eagles (Crain et al 2015) : 100 Mpc/h 2.6 Kpc/h

e.g: High-z QSO lum. Function (Giallongo 15)

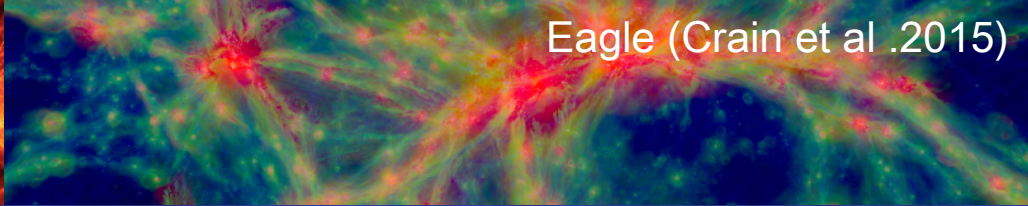
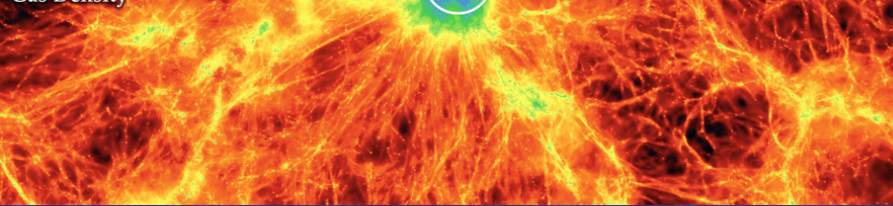




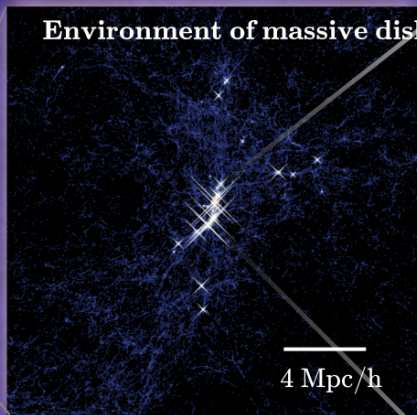






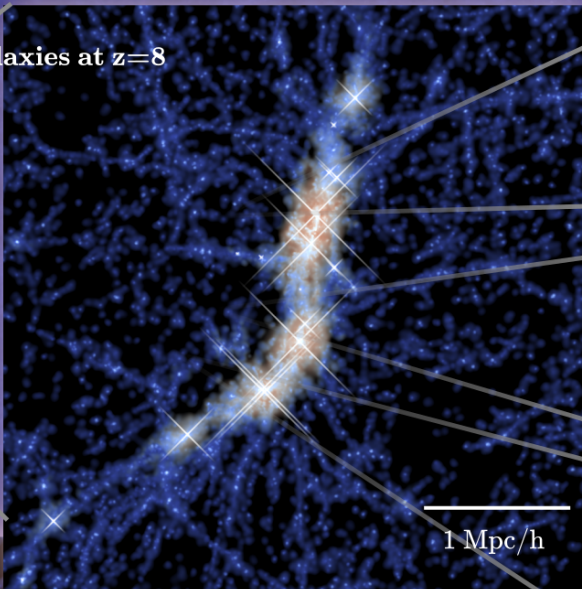


Eagle (Crain et al .2015)



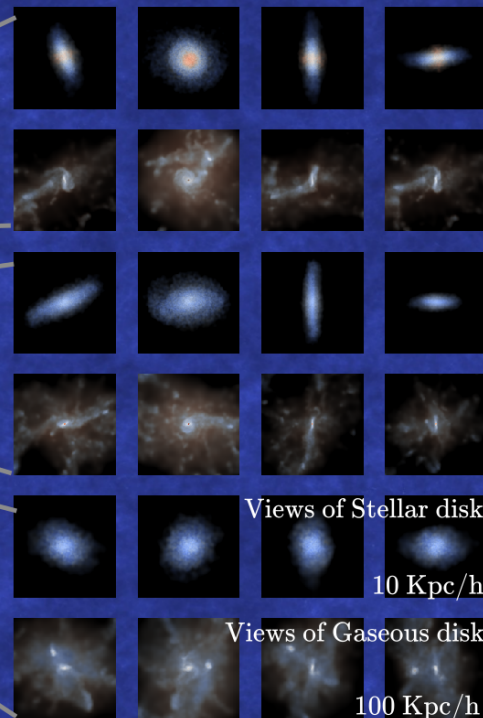
Environment of massive disk galaxies at  $z=8$

4 Mpc/h



Environment of most massive blackhole at  $z=8$

1 Mpc/h

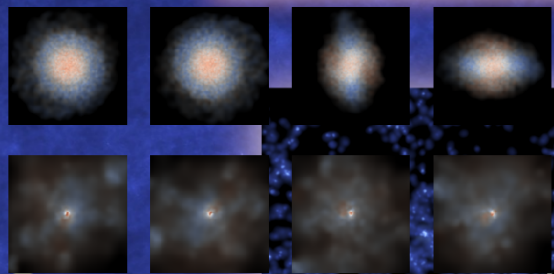


Views of Stellar disk

10 Kpc/h

Views of Gaseous disk

100 Kpc/h



### The **BlueTides** Simulation

0.7 trillion particles

0.65 million cores



# bluetides

Feng et al. 2015

40 Mpc/h

# BlueTides Simulation:

0.7 million cores on NCSA [BlueWaters](#)



## Goals:

- Technology Path Finder for future hydro simulation
- Predictions for high-redshift surveys

714 Mpc on the side

200 pc resolution at  $z=9$

$2 \times 7040^3$  (0.7 trillion) particles

Star formation/ AGN model compatible with Illustris

**50 times bigger volume** with highest resolution



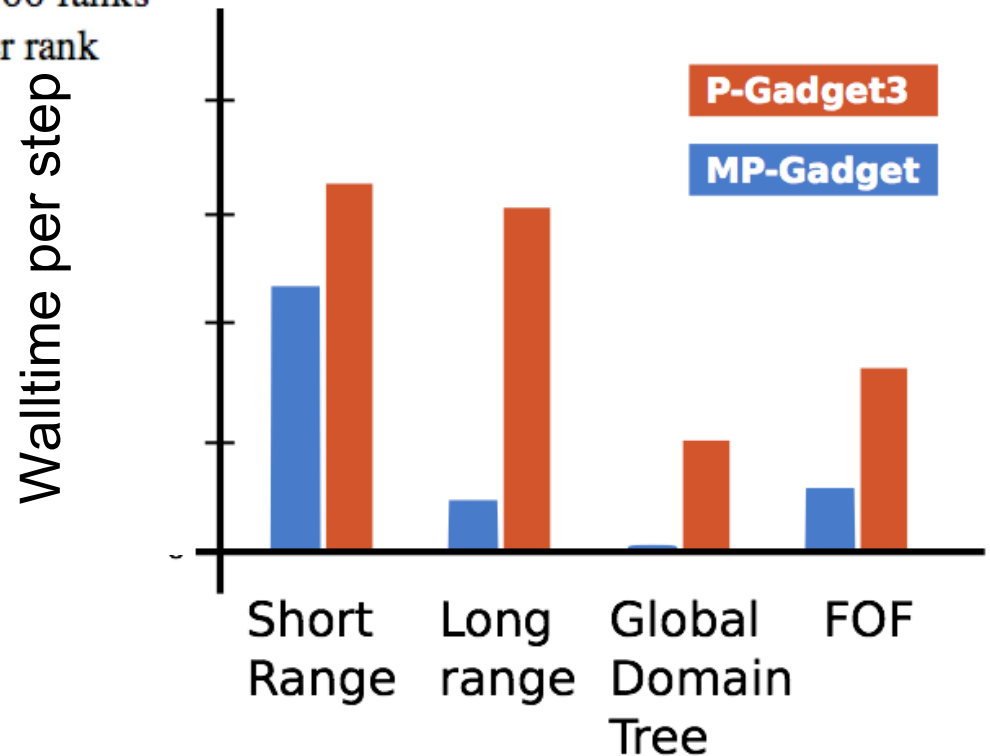
# BlueTides Simulation: Technology

## Technology Path-Finder:

- From P-Gadget to MP-Gadget
- 81000 MPI ranks, 8 OpenMP threads per rank
- Large, distributed FFT: 10,000 mesh on 81000 ranks
- Efficient thread parallelism up to 32 threads per rank

## Spinoffs and Open source contributions:

- bigfile : hierarchical snapshot format
- MP-sort : practical parallel sorting
- sharedmem : parallel data analysis
  
- PFFT : large-scale distributed FFT



# Long range force calculation (PM): New solver:

E.g. 8 processes:

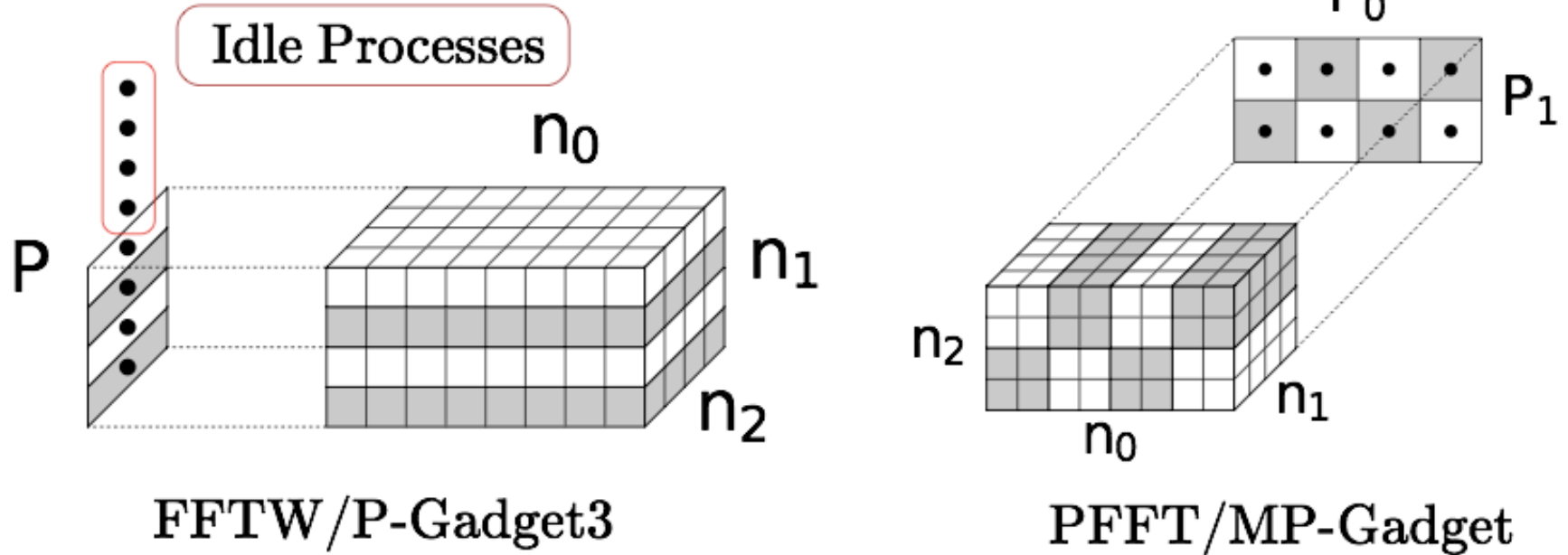
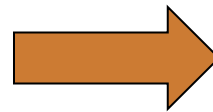


Figure from M.Pippig 2013

## Blue Tides:

N= 10000 slabs  
on 81000 MPI ranks



Pencil beam domain  
decomposition

**8 x speed-up**

Open Source: Added new Array-execution interface and python binding to PFFT  
(<http://github.com/mpip/pfft>)



# BlueTides Simulation: Science

## Physics modelling in BlueTides

- Hydrodynamics (pSPH)
- Primordial cooling
- Multi-phase medium star-formation
- SN wind feedback
- H<sub>2</sub> molecule fraction
- AGN feedback
- Metal enrichment and cooling
- Non-uniform UV background calibrated from rad. Hydro sims (Battaglia+13)

## Science of high redshift galaxy

- a statistical sample of high redshift galaxies, accessible only via uniform simulations
- reionization
- morphology
- mock surveys
- high redshift AGNs
- ....

The first 500 million years ( $z=7+$ ):

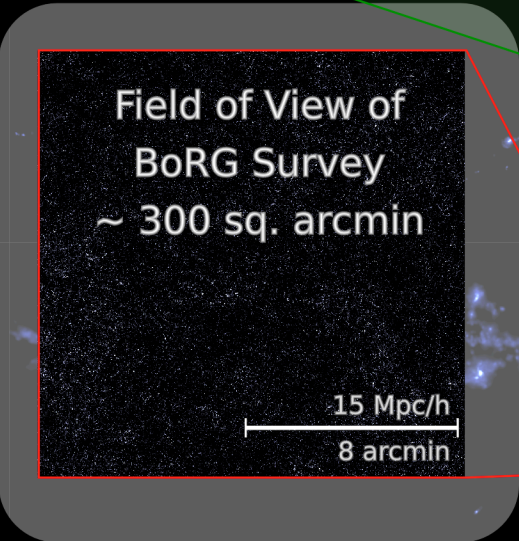
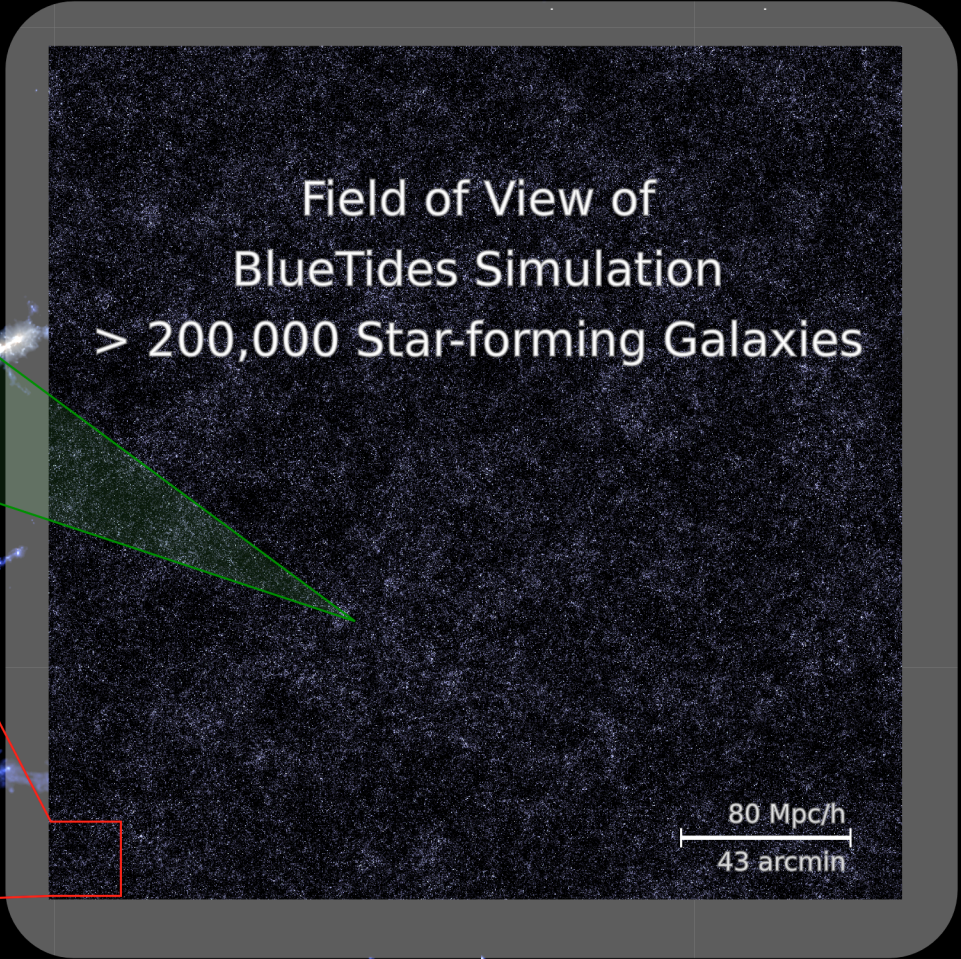
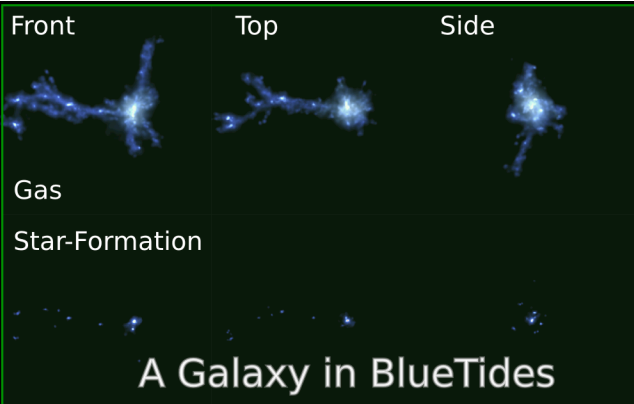
What are galaxies like?

Problem 1:

Predictions from  
BlueTides:

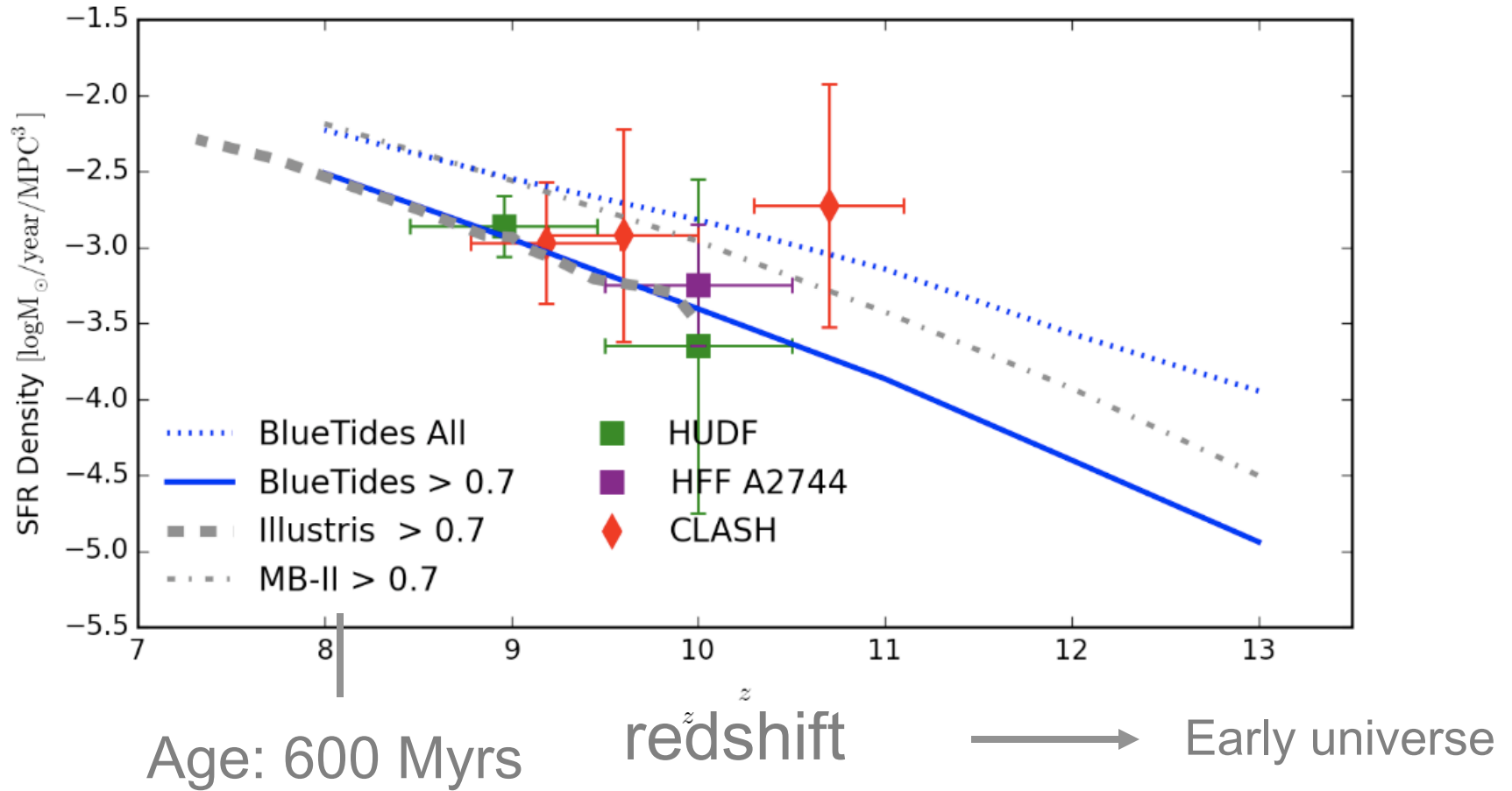


# BlueTides 400 x volume of HUDF



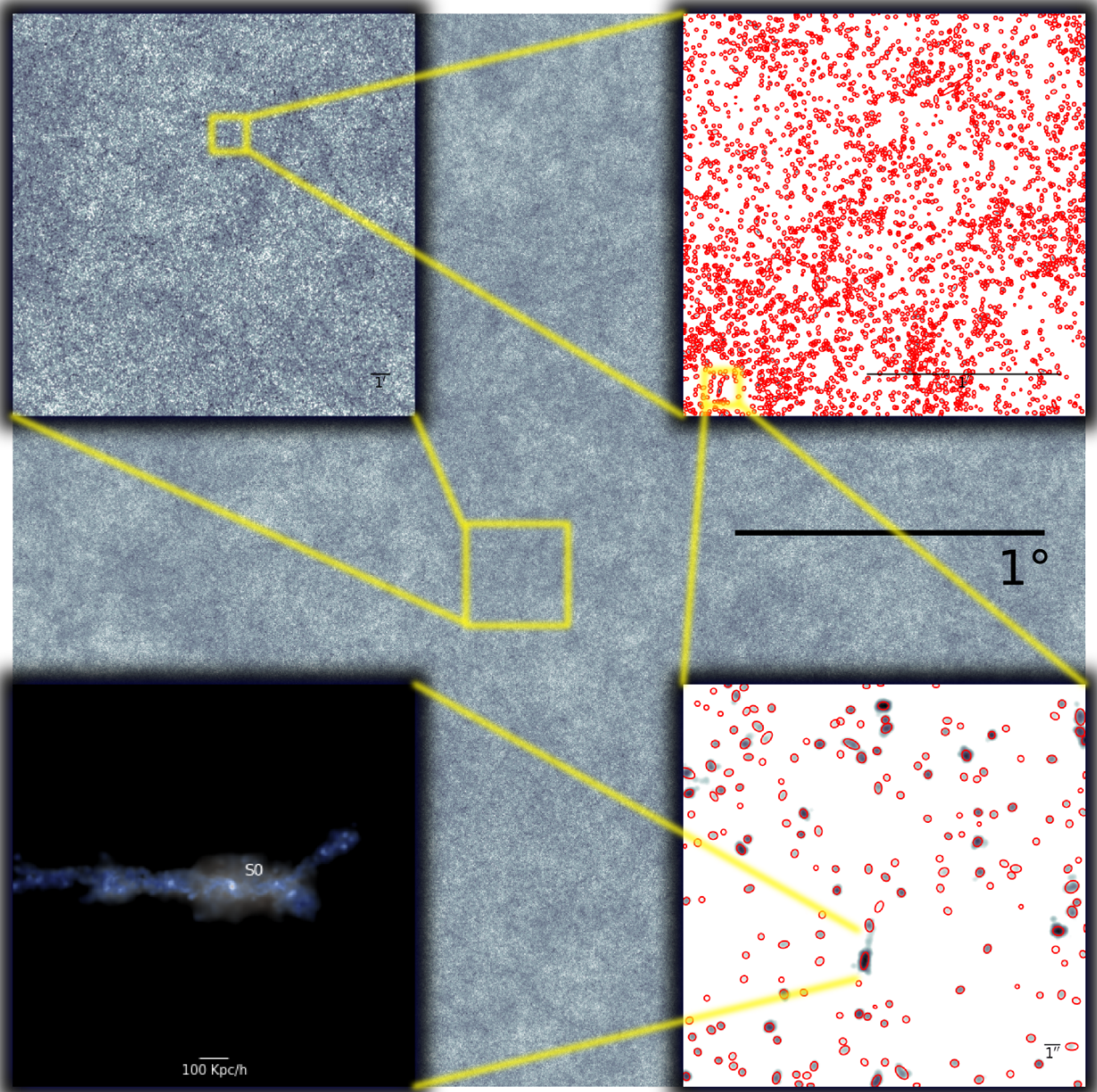
# BlueTides Simulation: Global SFRD is consistent with current observational limits.

Star formation rate density



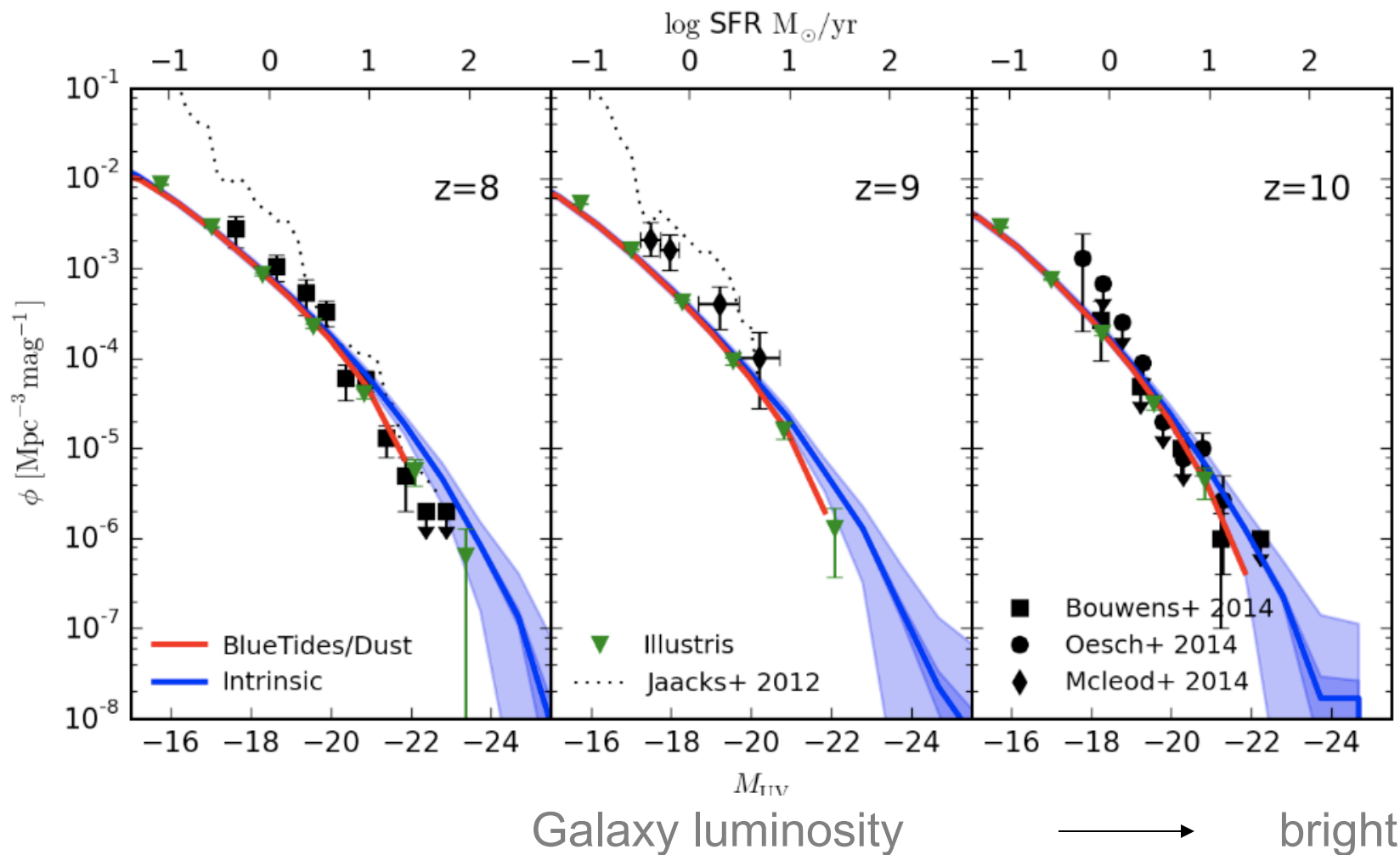


# Simulations like Observations: Create Mock Fields. Source extract detection to find galaxies

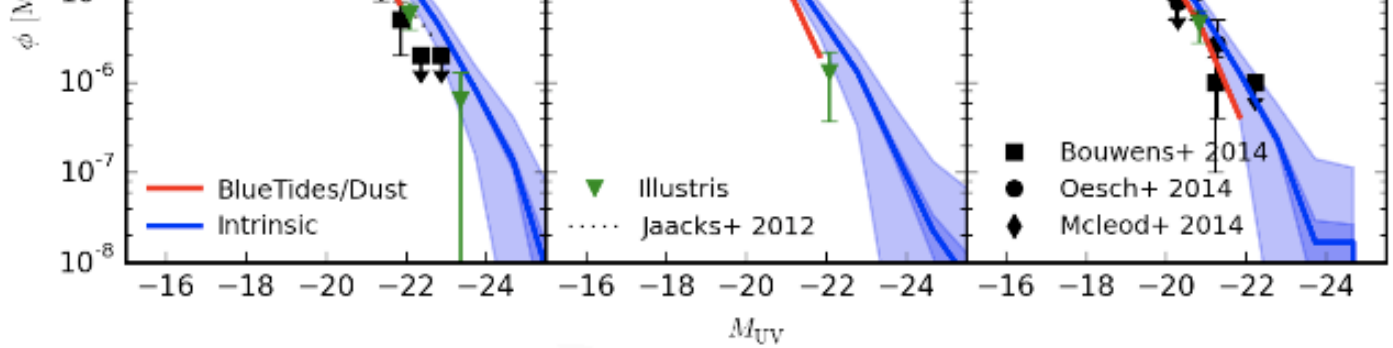


# Galaxy Luminosity Function in BlueTides consistent with Hubble Legacy Fields

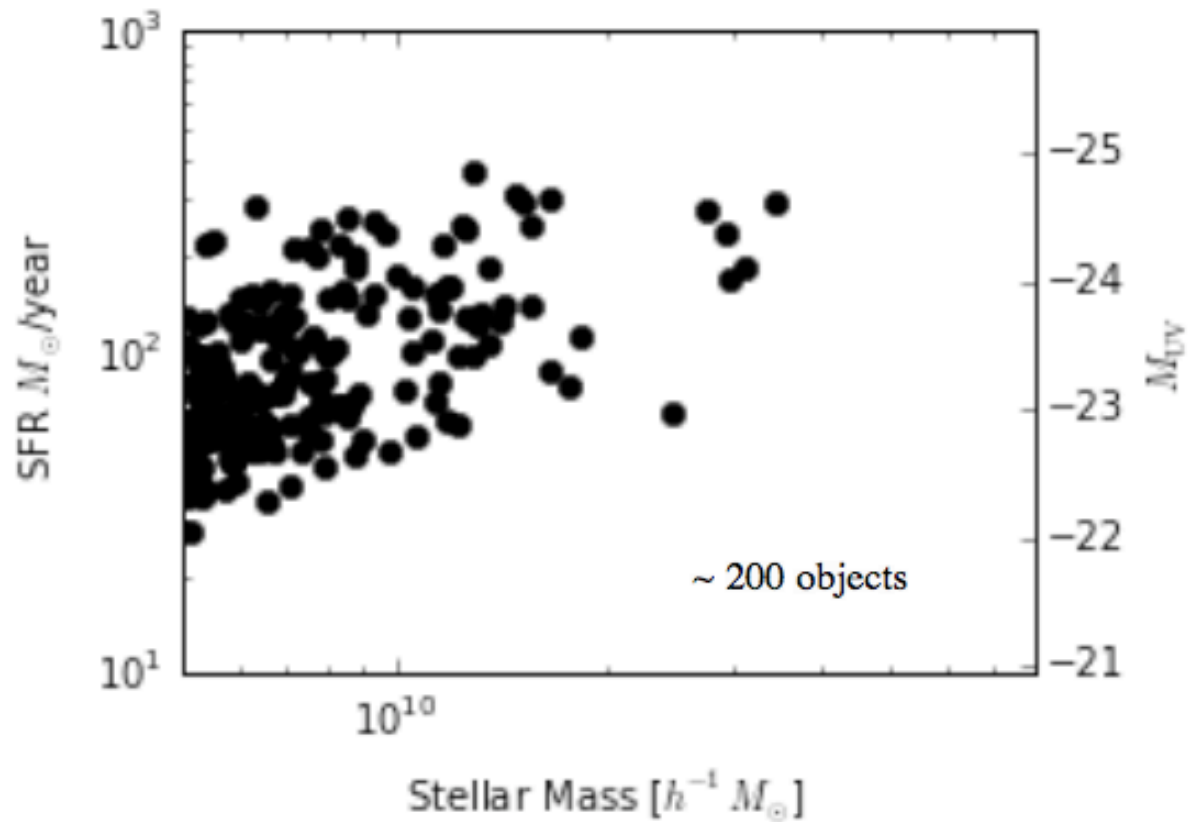
Diff. Number density of galaxies



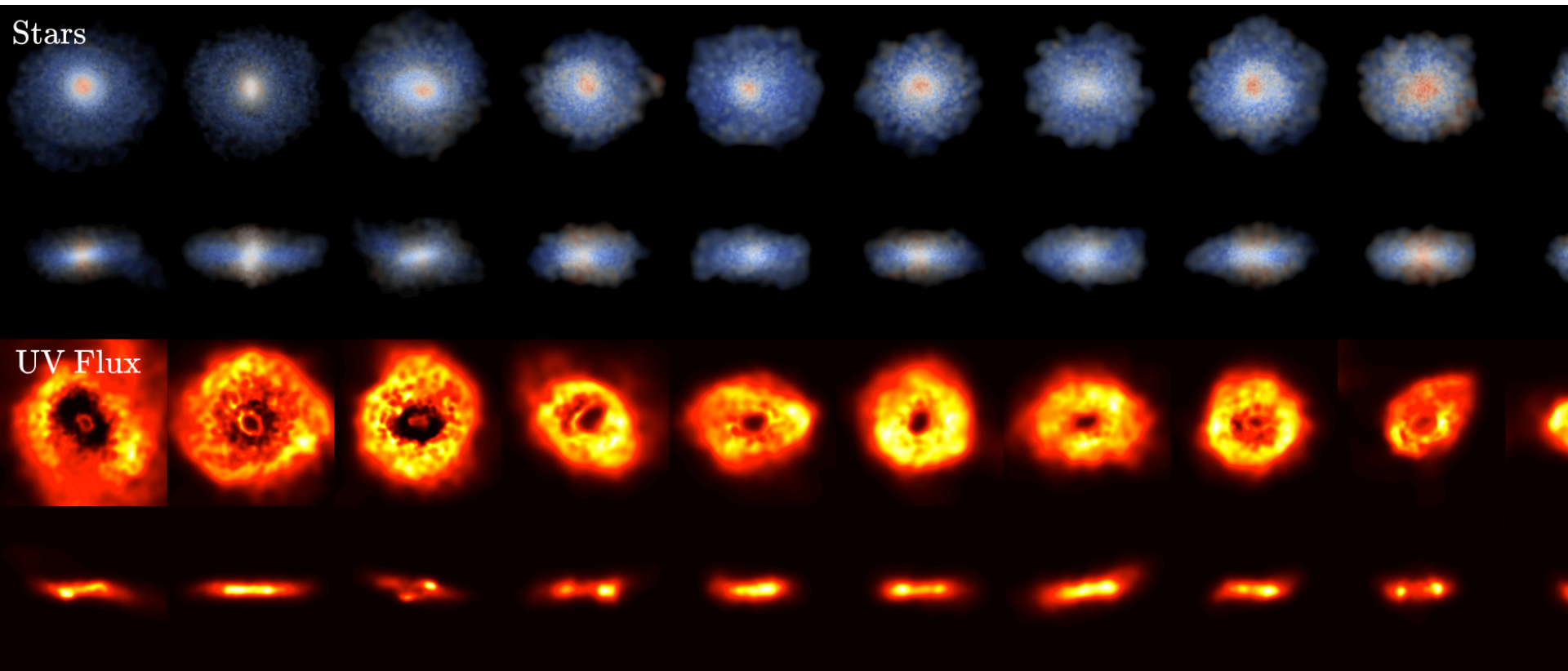




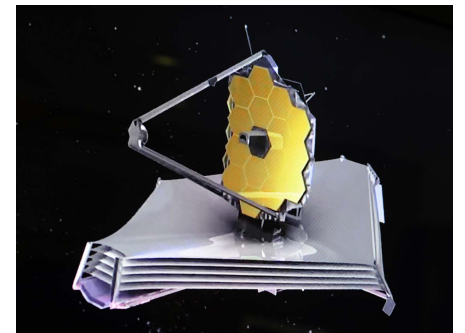
Predictions  
 for the brightest  
 Galaxies at  $z=8$ :



# $z=8$ Milky Way (/Massive) galaxies are **disks!**

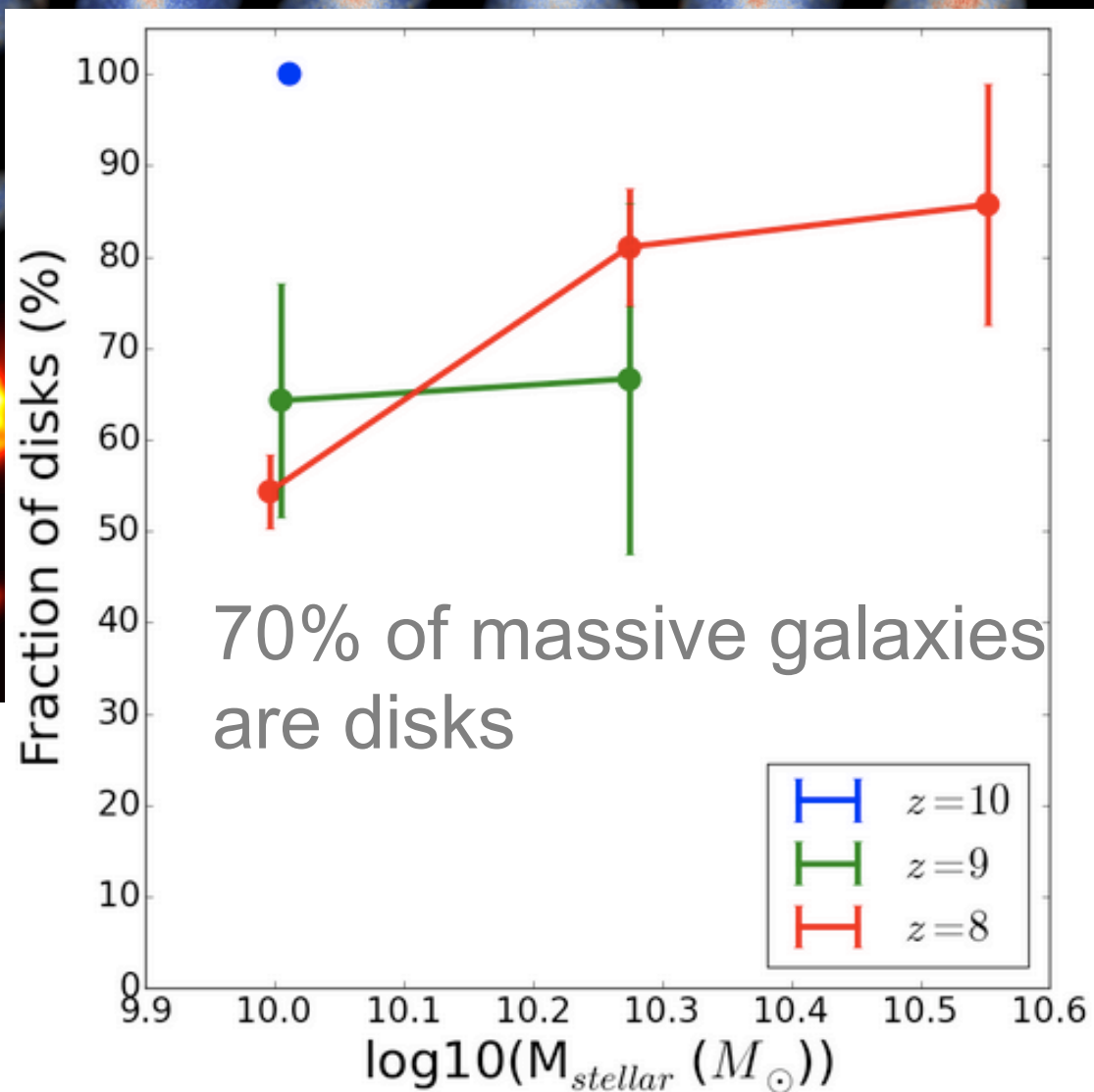
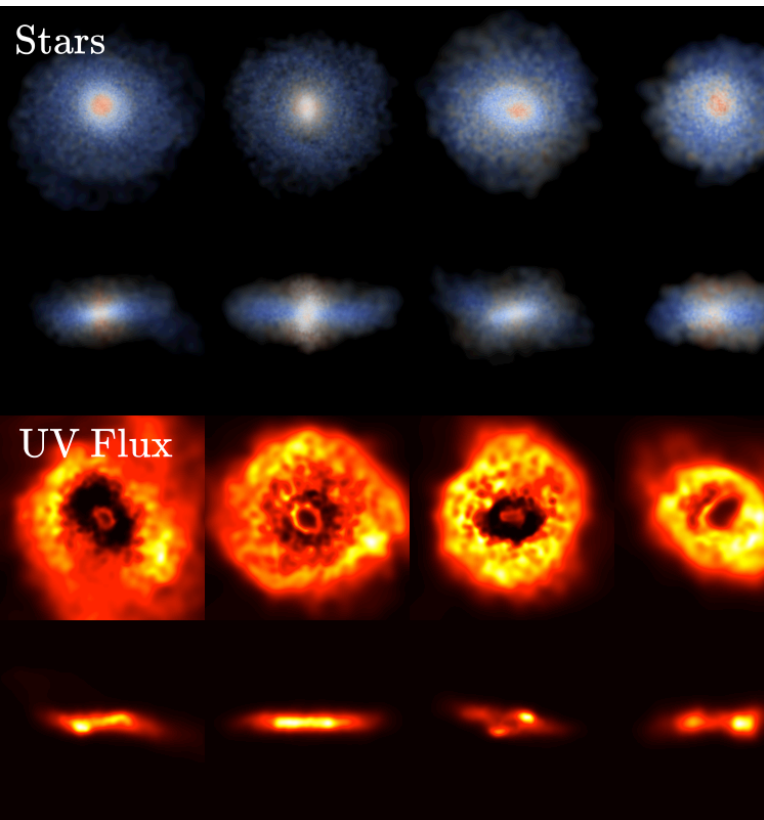


JWST

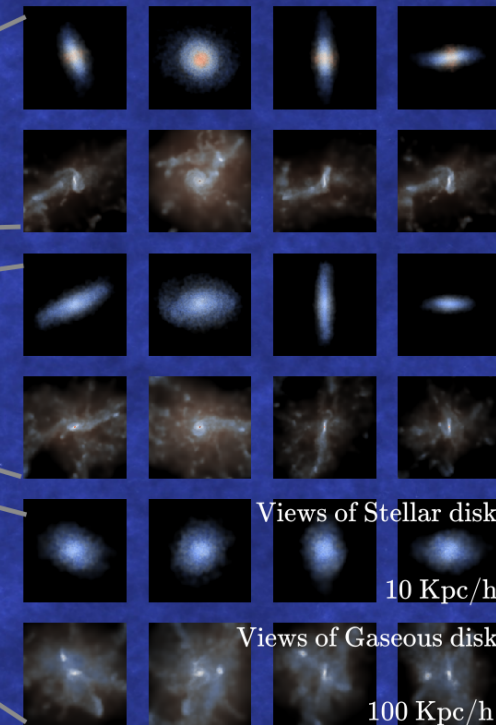
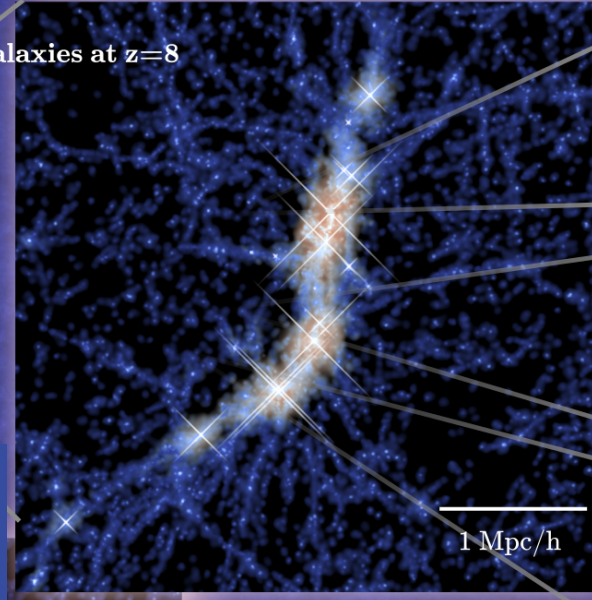
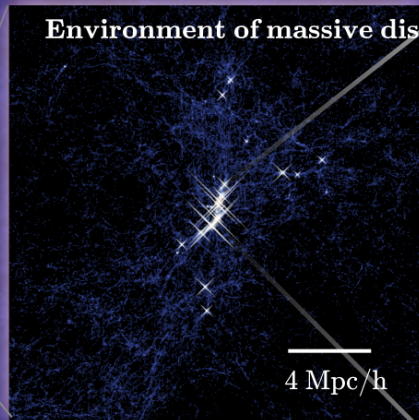




# $z=8$ Milky Way (/Massive) Halos look like **disks!**



Environment of massive disk galaxies at  $z=8$



WFIRST  
should detect  
~ 8000 Milky Way  
mass disks  
at  $z=7-8$



The **BlueTides** Simulation

0.7 trillion particles

0.65 million cores

BLUE WATERS  
SUSTAINED PETASCALE COMPUTING



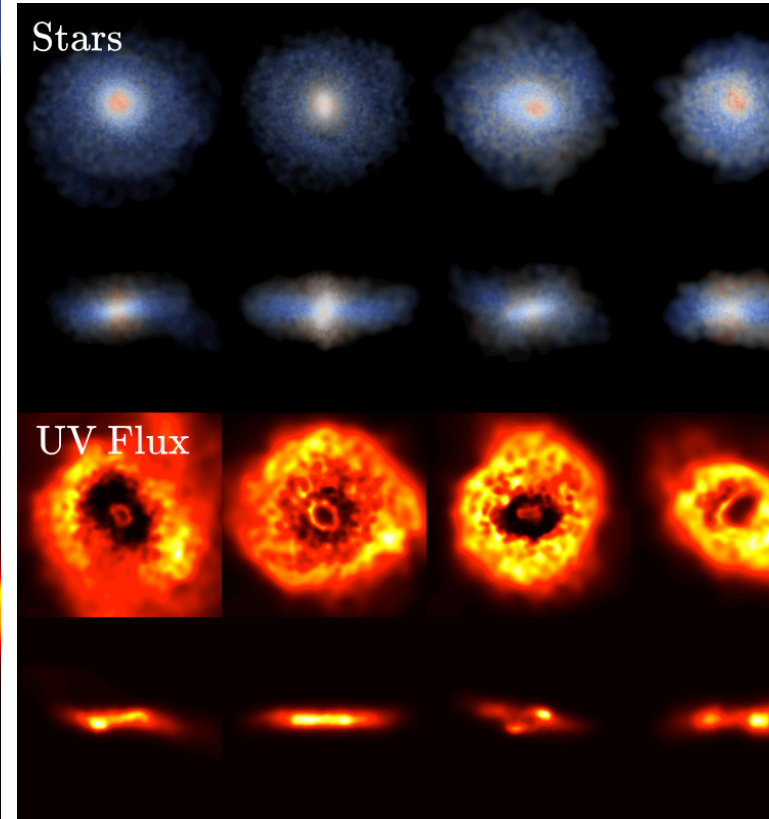
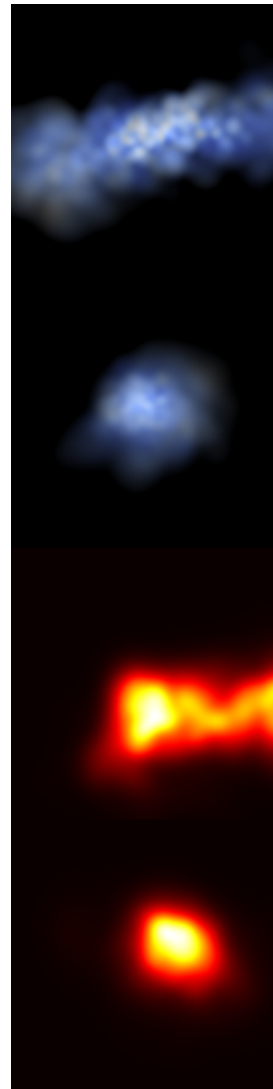
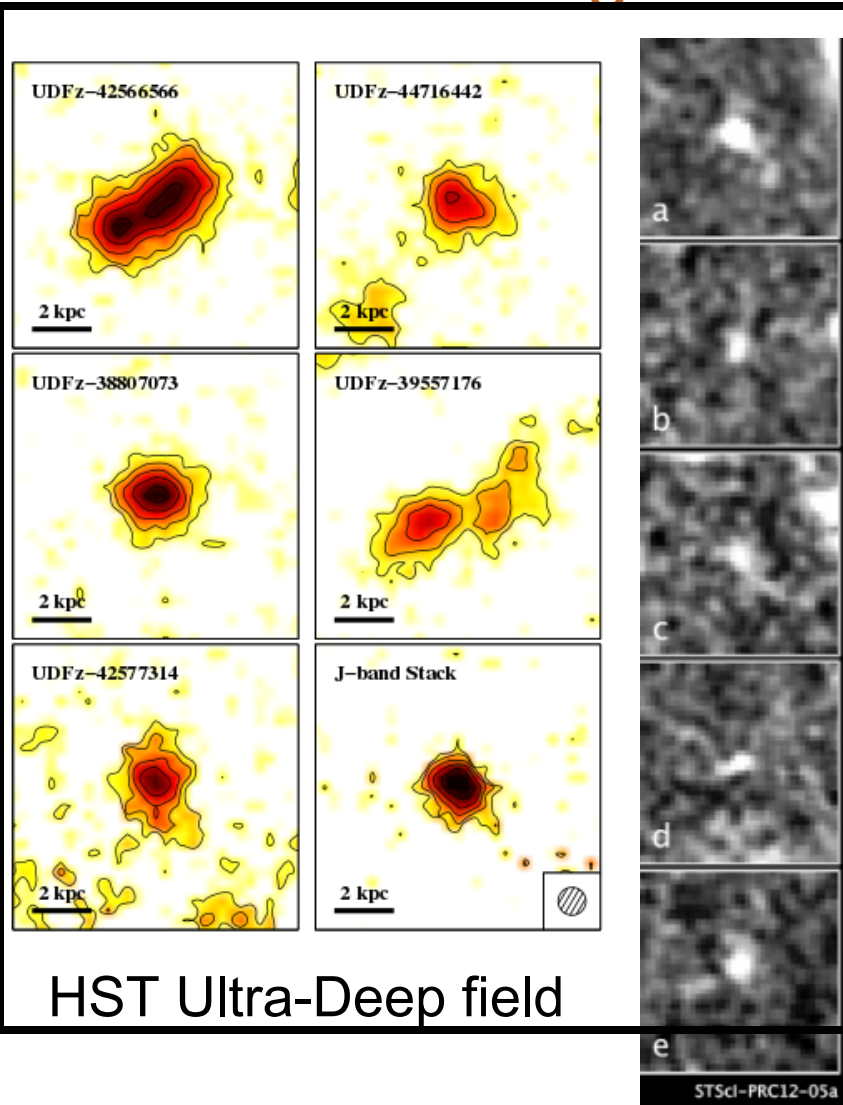
bluetides

Feng et al. 2015



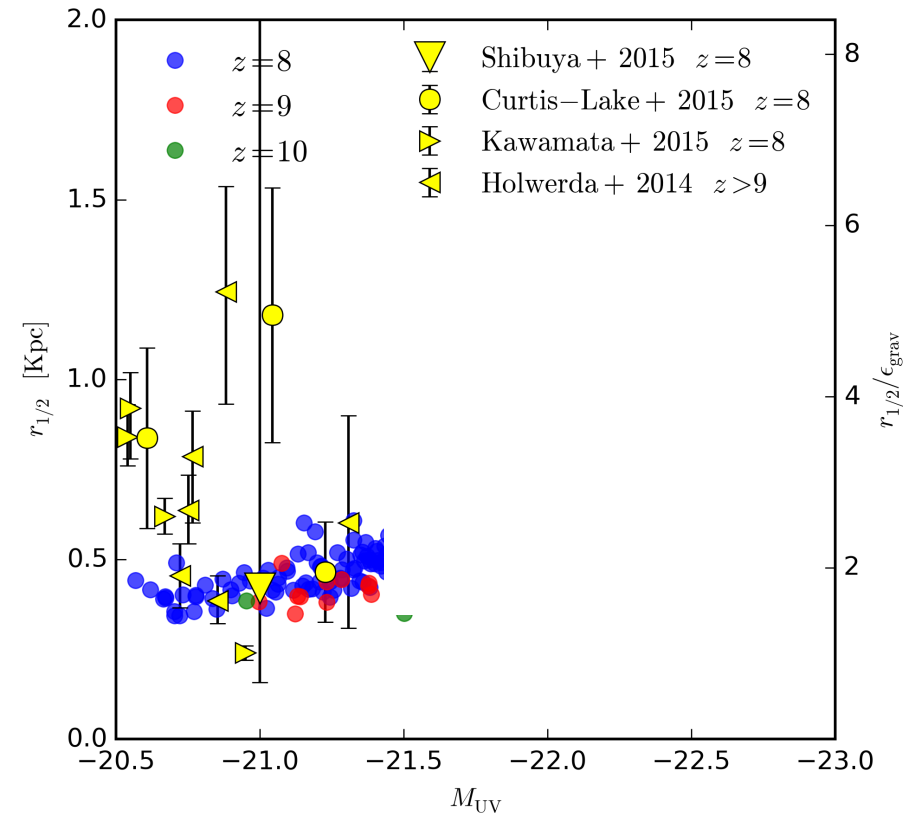
$z=8$  Milky Way (Massive) galaxies are **disks!**

But small faint galaxies are irregular ....



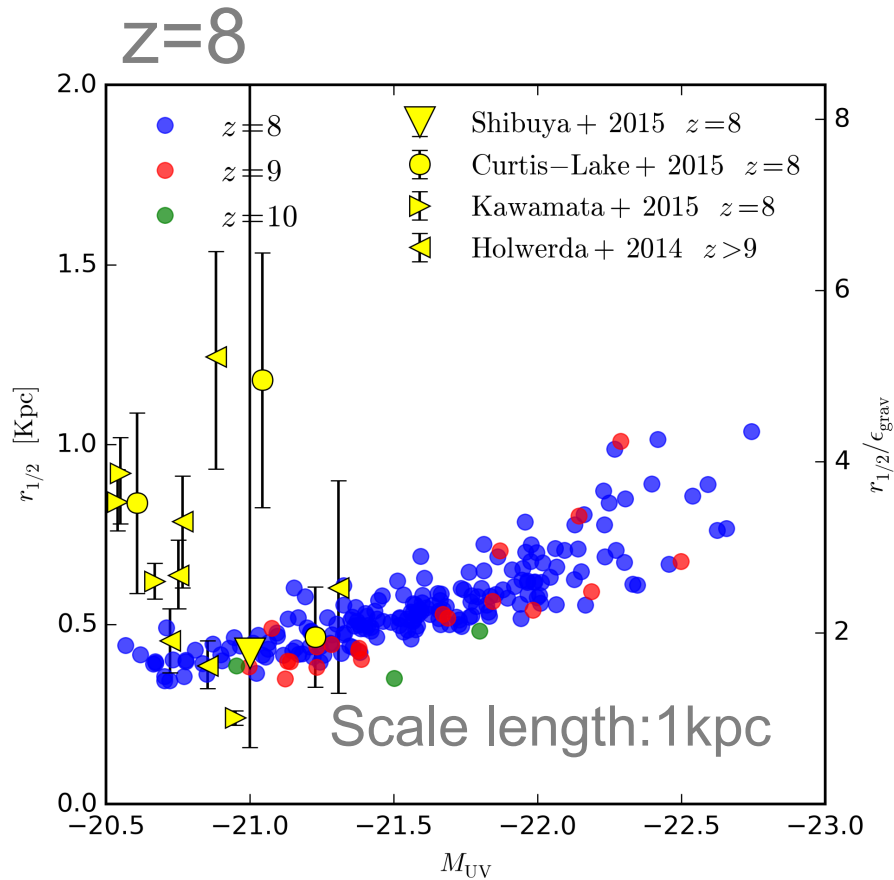
# The sizes of galaxies in BlueTides are consistent with HST observations

$z=8$

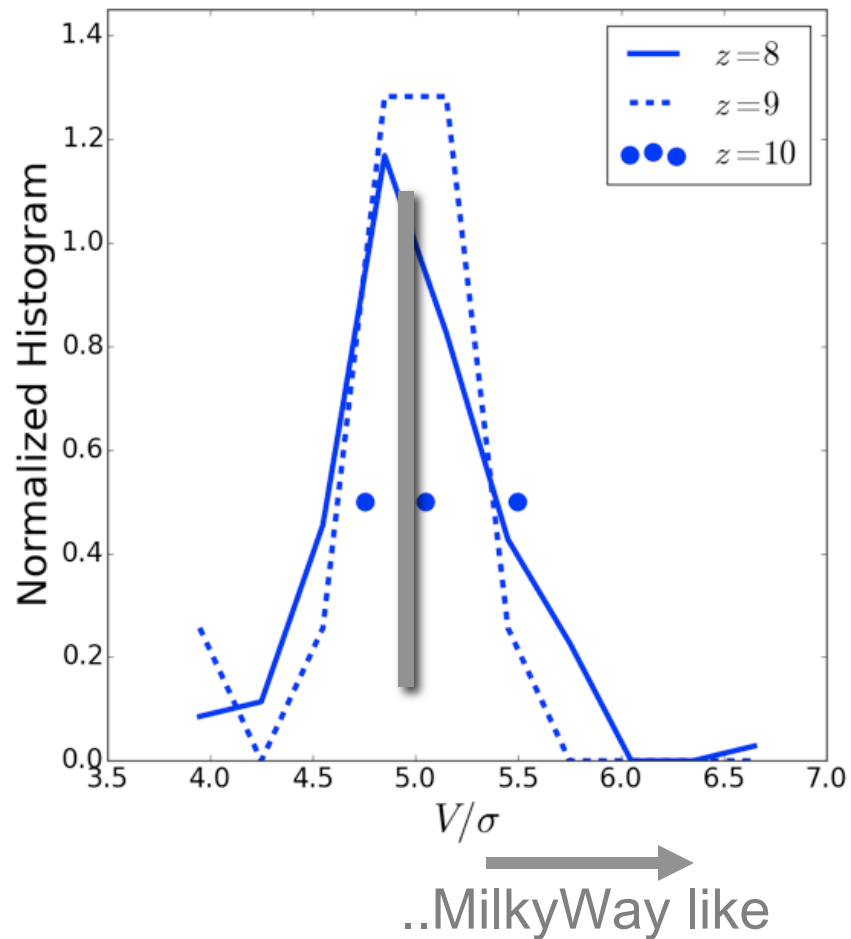
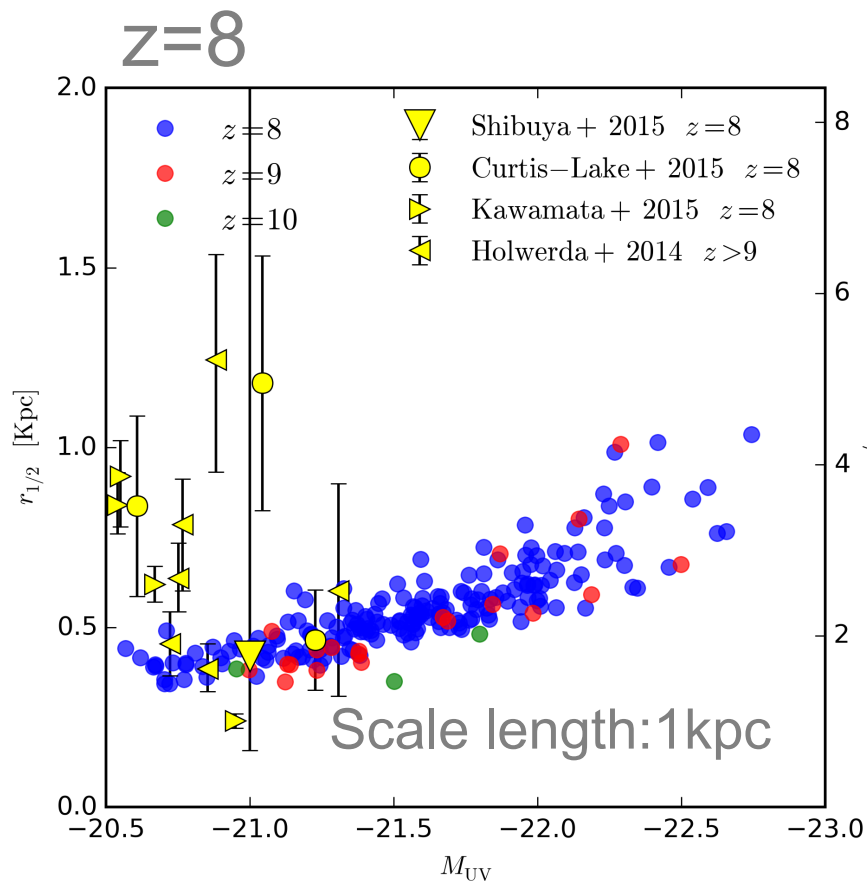




The sizes of galaxies in BlueTides are consistent with HST observations --> **‘massive’ disks in bright galaxies are compact**



The sizes of galaxies in BlueTides are consistent with HST observations --> **‘massive’ disks in bright galaxies are compact rotationally supported**

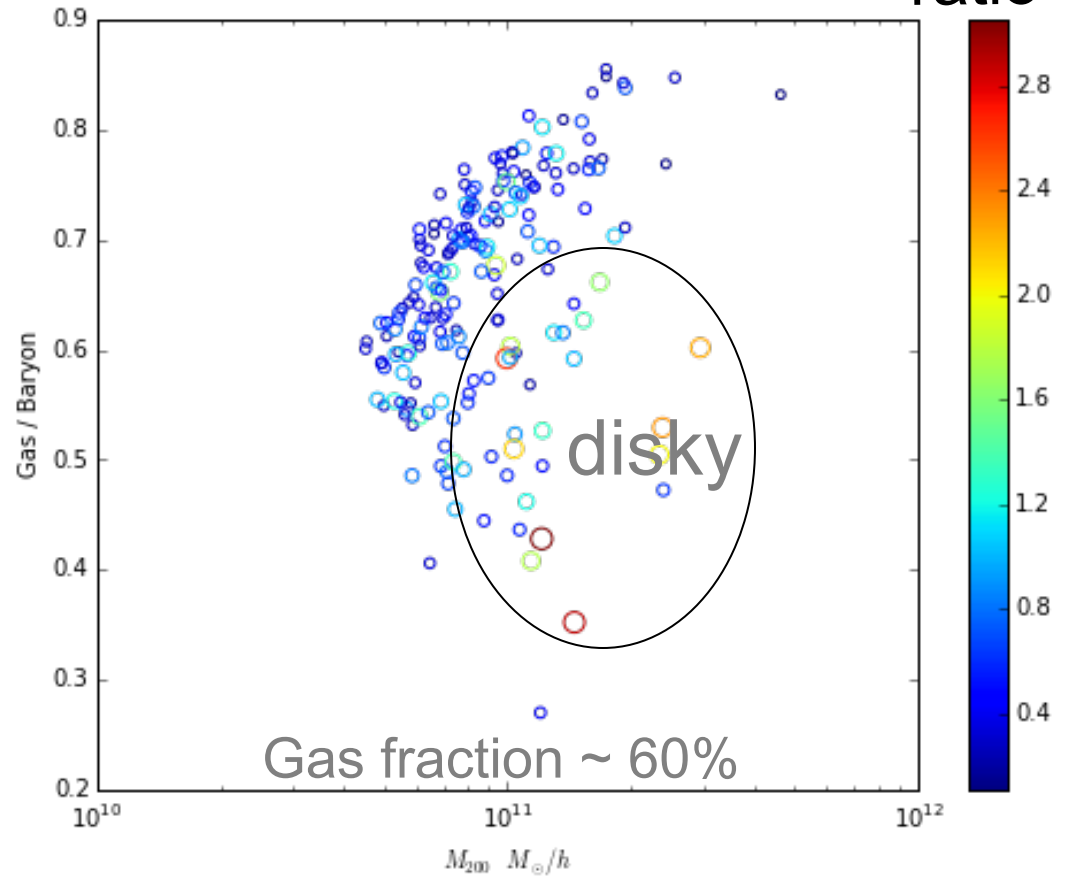
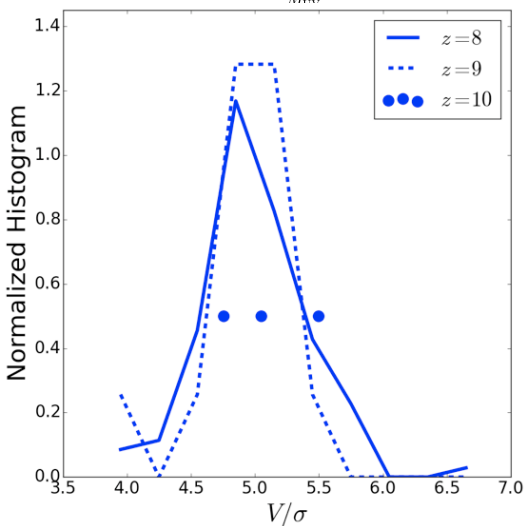
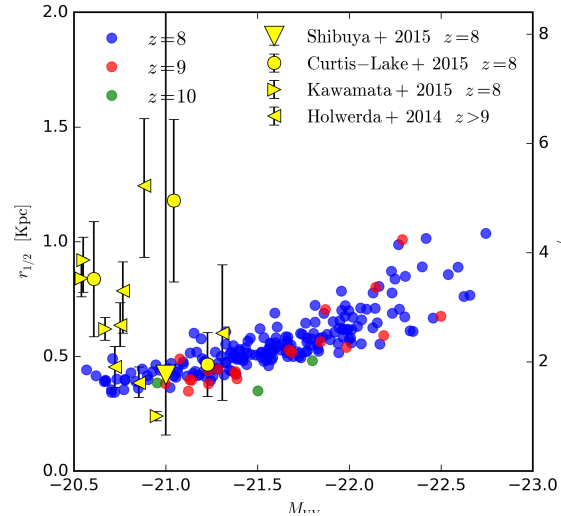




# Massive Disks at $z=8$ , more compact gas rich

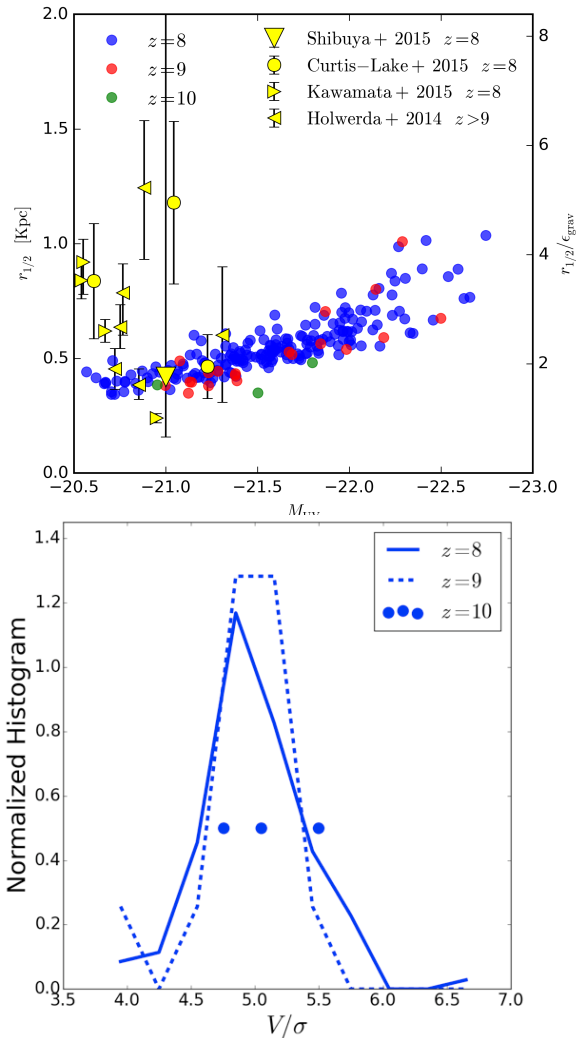
$z=8$

Disk/Bulge ratio

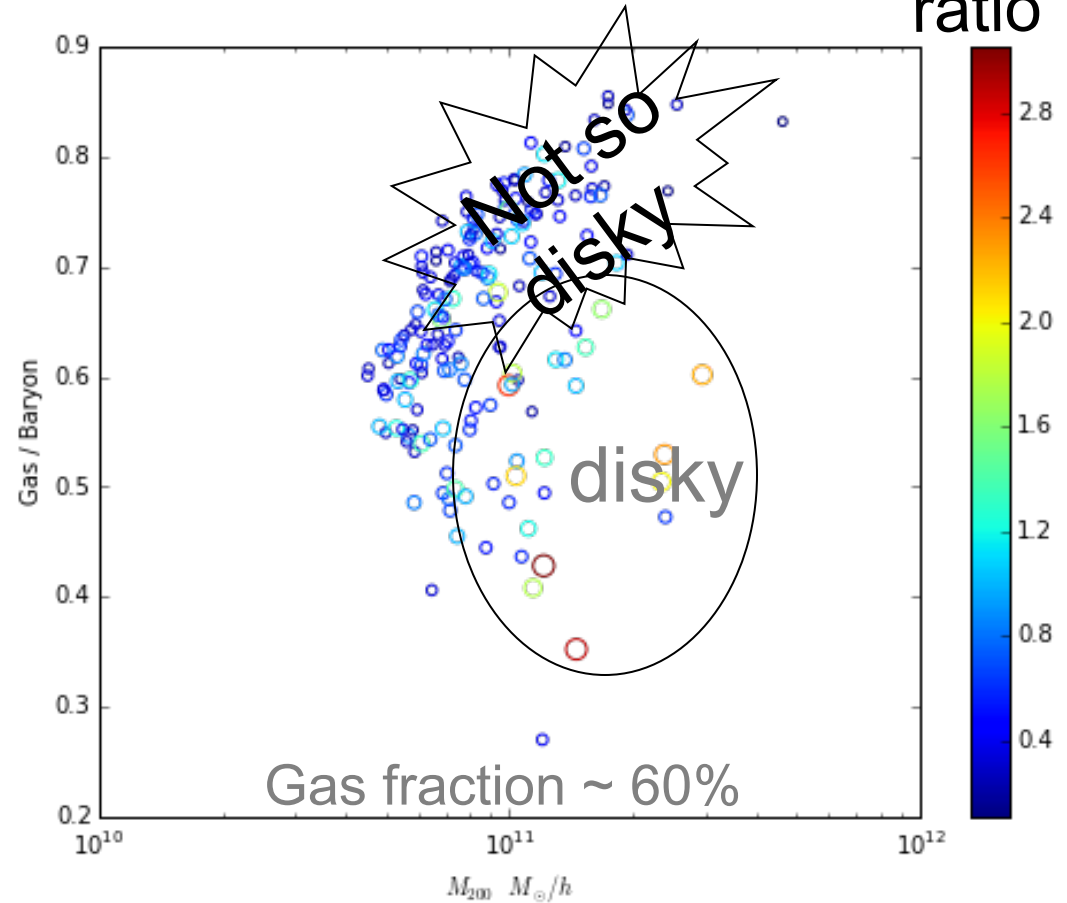


# Gas richest systems have less dominant disk....

$z=8$



Disk/Bulge  
ratio



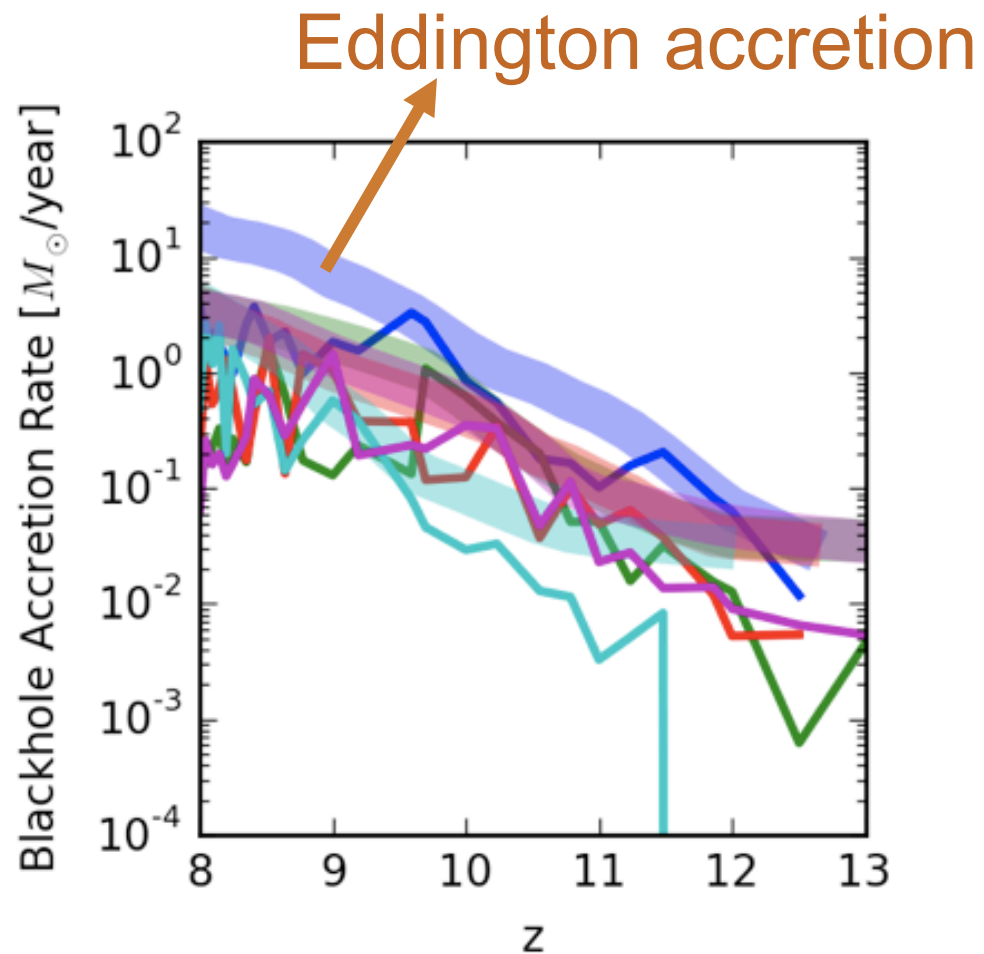
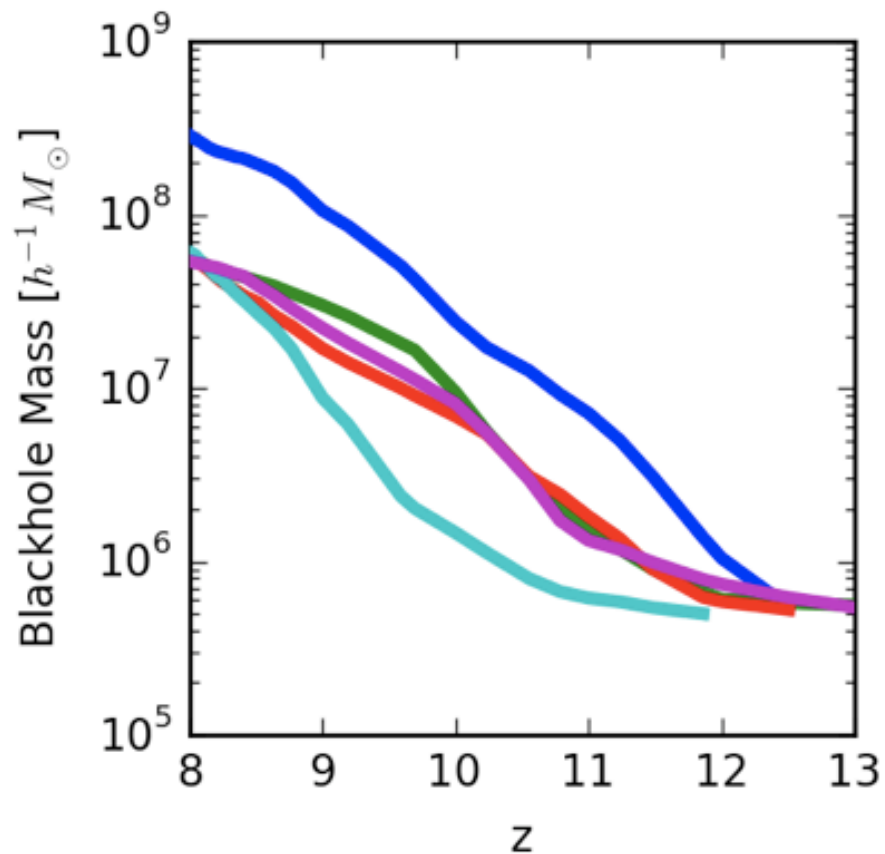
The first 600 million years ( $z=7+$ ):

What are the most massive BH/  
quasars like?

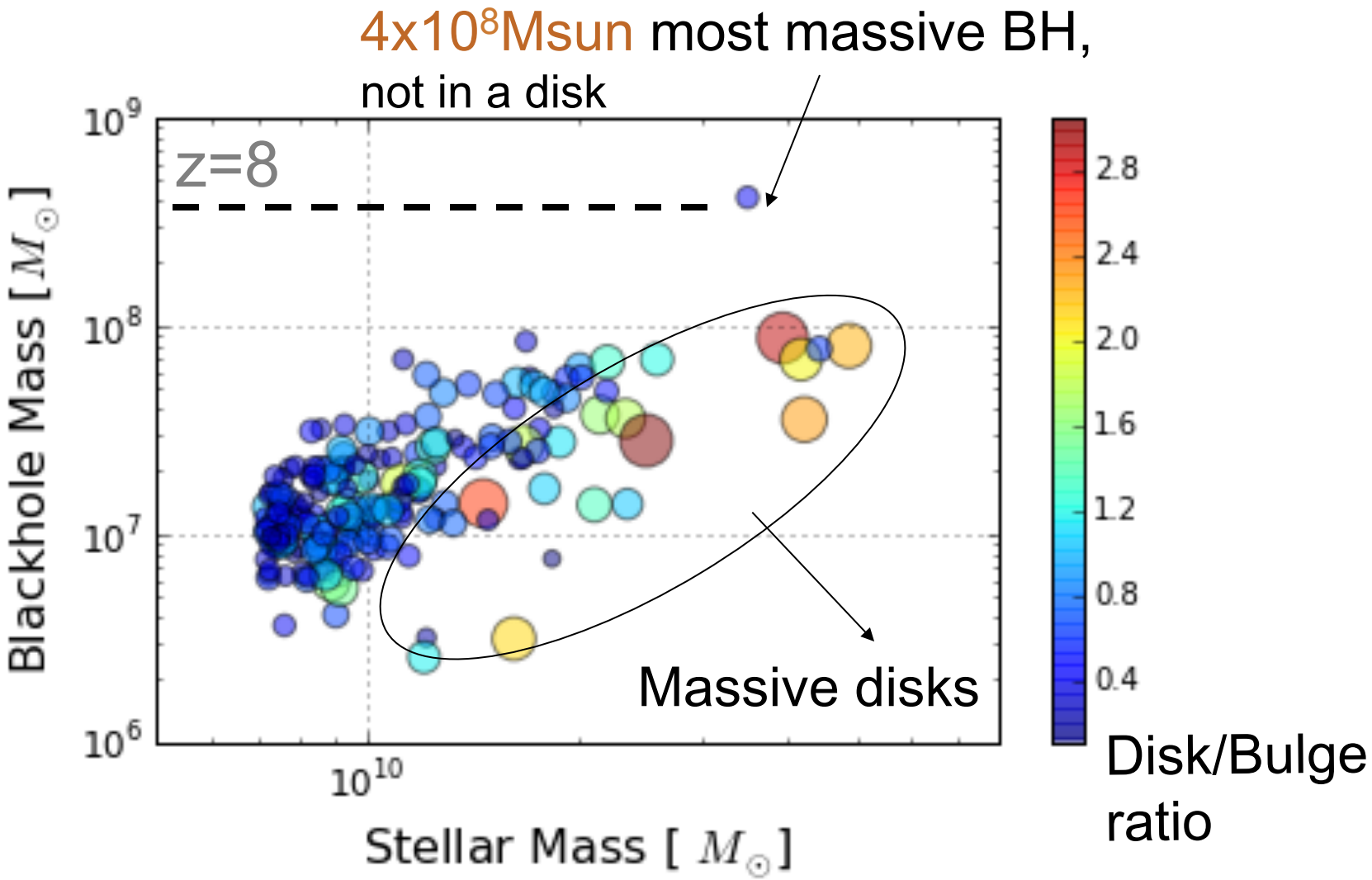
Predictions from  
BlueTides:



# Assembly of largest BHs at $z=8$ ,

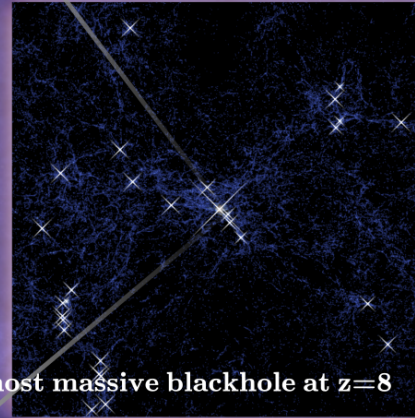
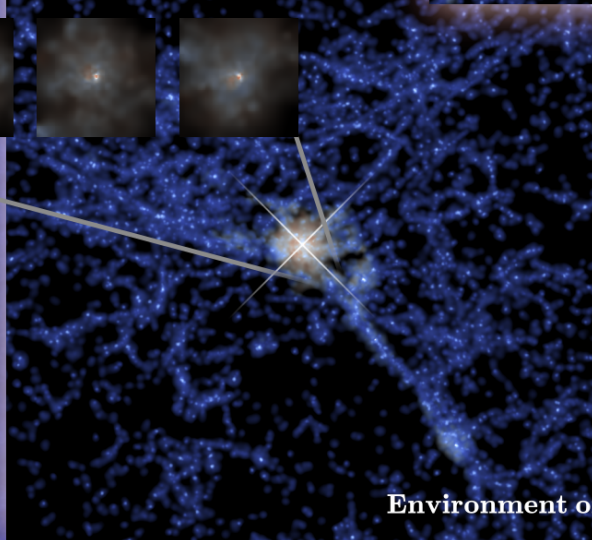
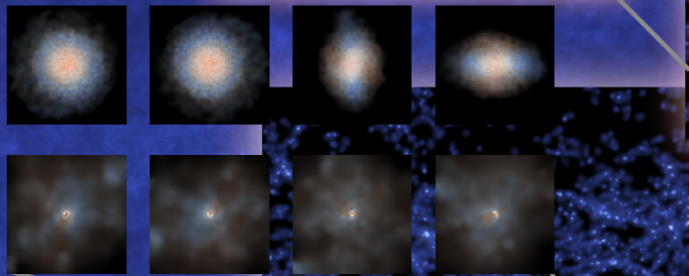


# Most massive black holes at $z=8$ , about $10^8 M_{\odot}$



Fastest growing, massive black holes are not found in the massive disk galaxies!

# The environment of the most massive black hole: radial inflow, compact, elliptical host galaxy



Environment of most massive blackhole at  $z=8$

40 Mpc/h

The **BlueTides** Simulation

0.7 trillion particles

0.65 million cores

**BLUE WATERS**  
SUSTAINED PETASCALE COMPUTING

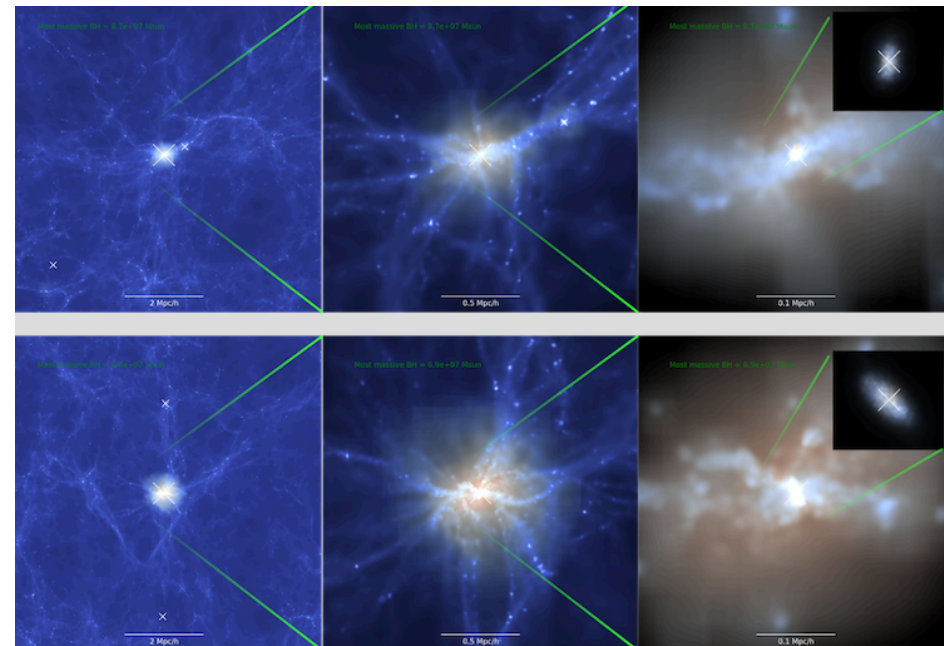
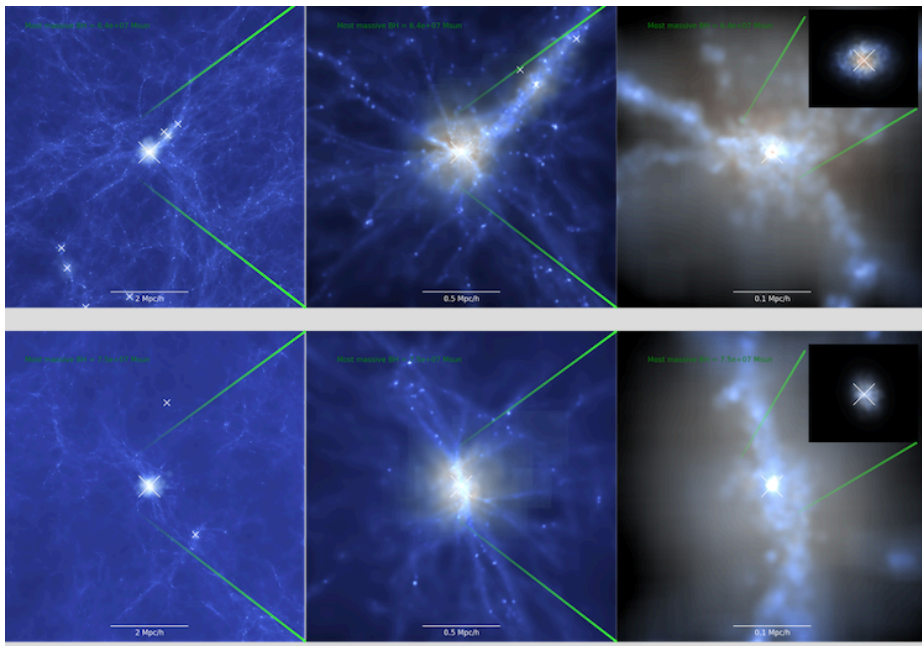


**bluetides**

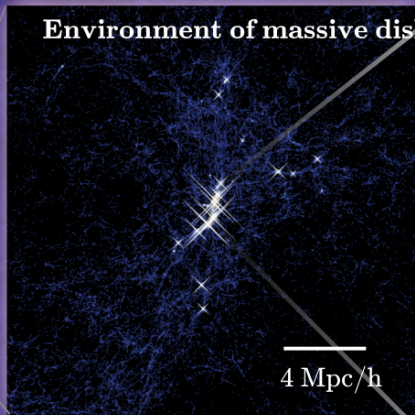
Feng et al. 2015



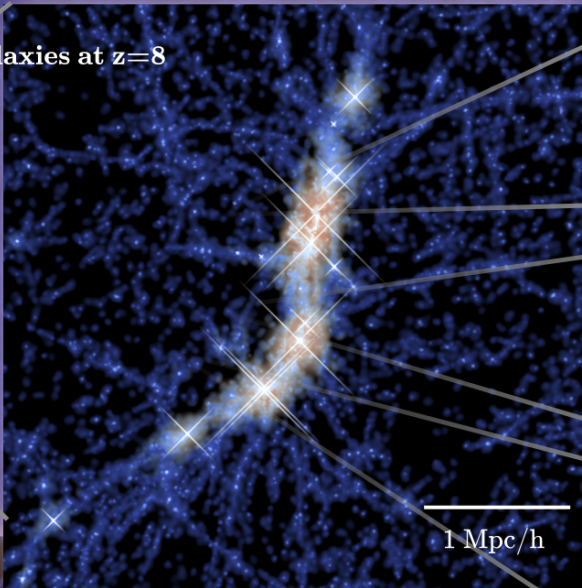
# The most massive BHs in isolated overdensities in supercompact spheroidal hosts



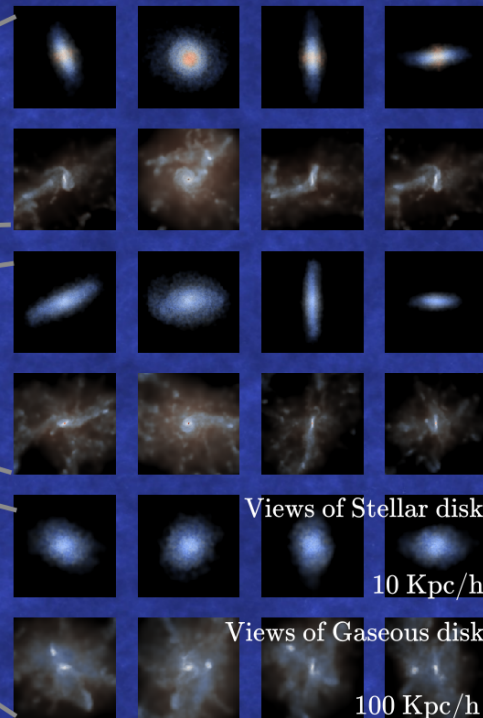
Environment of massive disk galaxies at  $z=8$



4 Mpc/h



1 Mpc/h

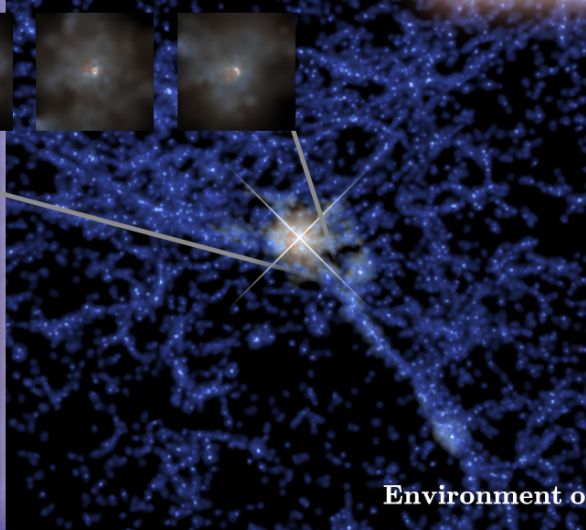
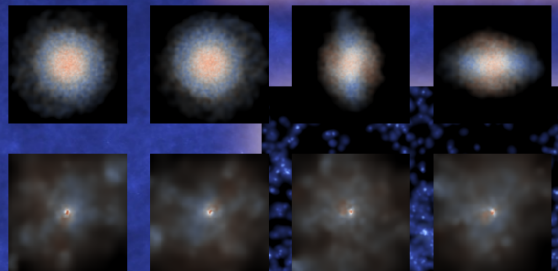


Views of Stellar disk

10 Kpc/h

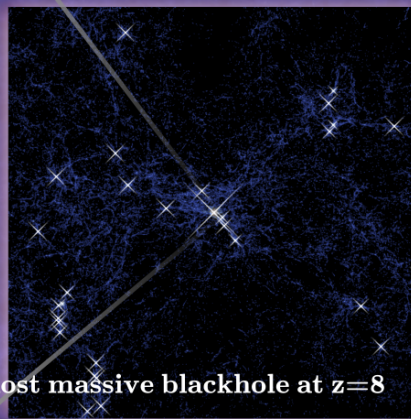
Views of Gaseous disk

100 Kpc/h



Environment of most massive blackhole at  $z=8$

40 Mpc/h



## The **BlueTides** Simulation

0.7 trillion particles

0.65 million cores

**BLUE WATERS**  
SUSTAINED PETASCALE COMPUTING



**bluetides**

Feng et al. 2015

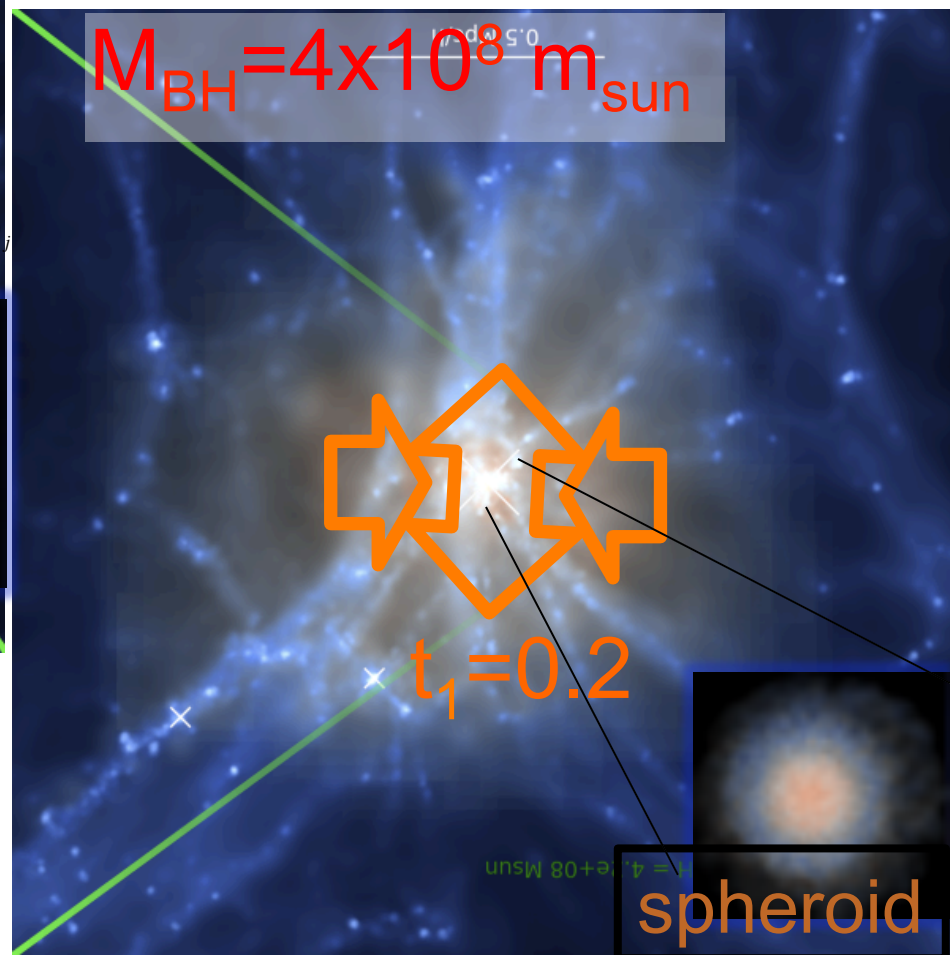
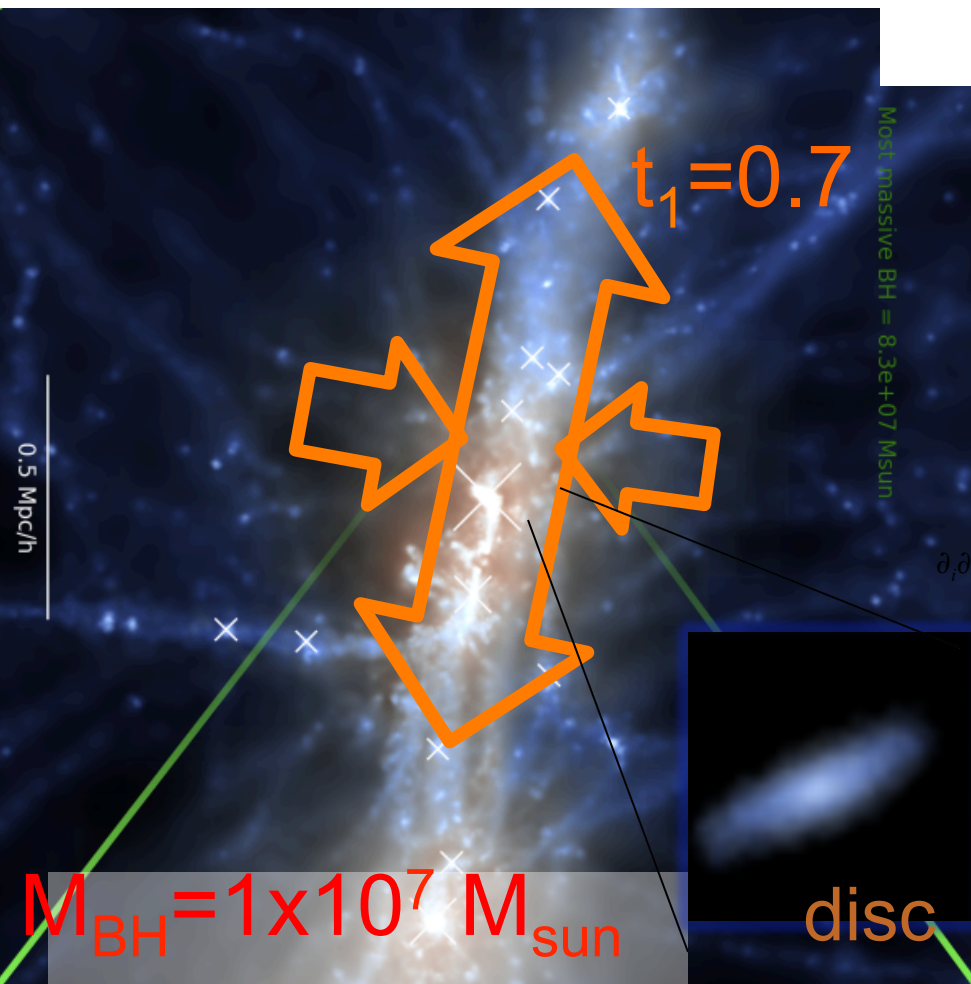


tidal tensor

$$T_{ij}(\mathbf{x}) = \frac{\partial^2 \phi}{\partial x_i \partial x_j},$$

weak tidal field:

Thin filaments  
radial motions,  
cold accretion

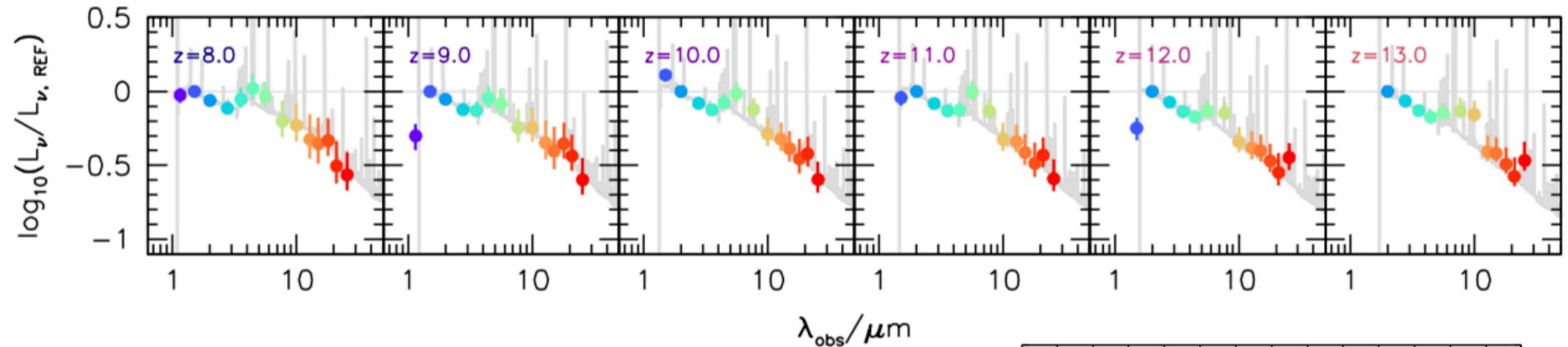


Large tidal field:  
Large filaments,  
Accretion perp. to  $t_1$

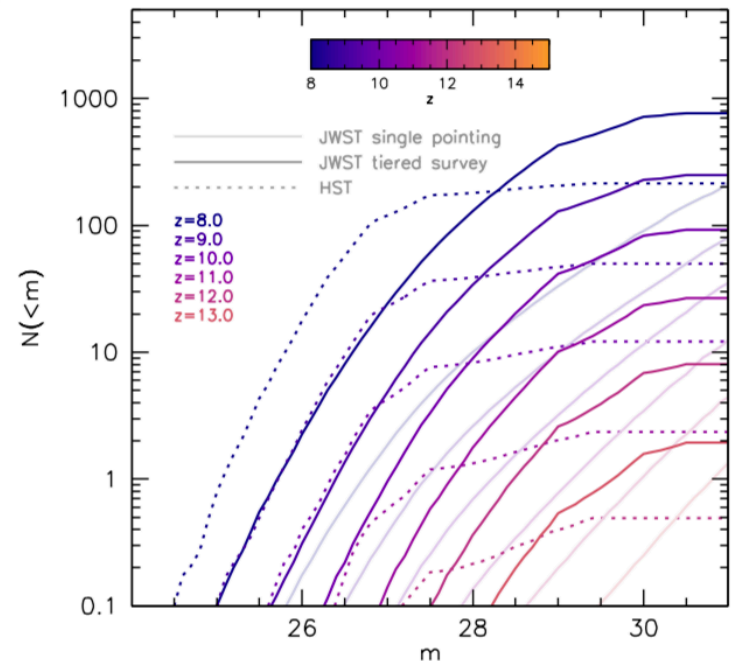
spheroid



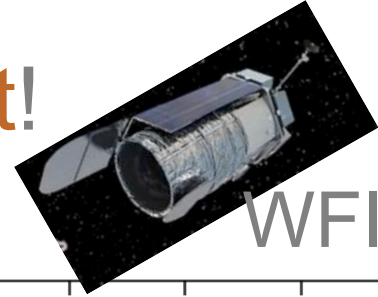
# Predictions for JWST



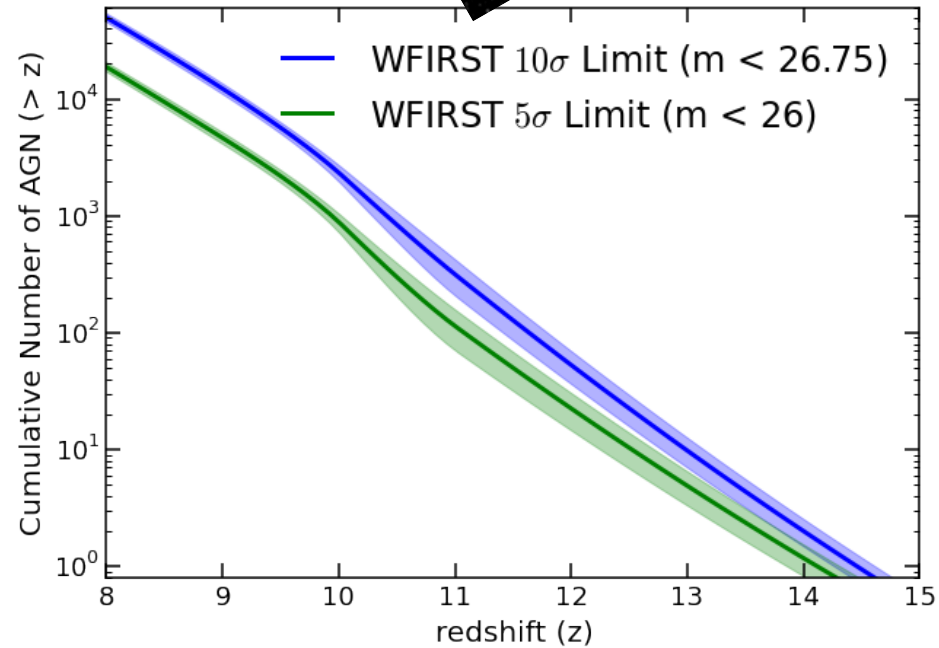
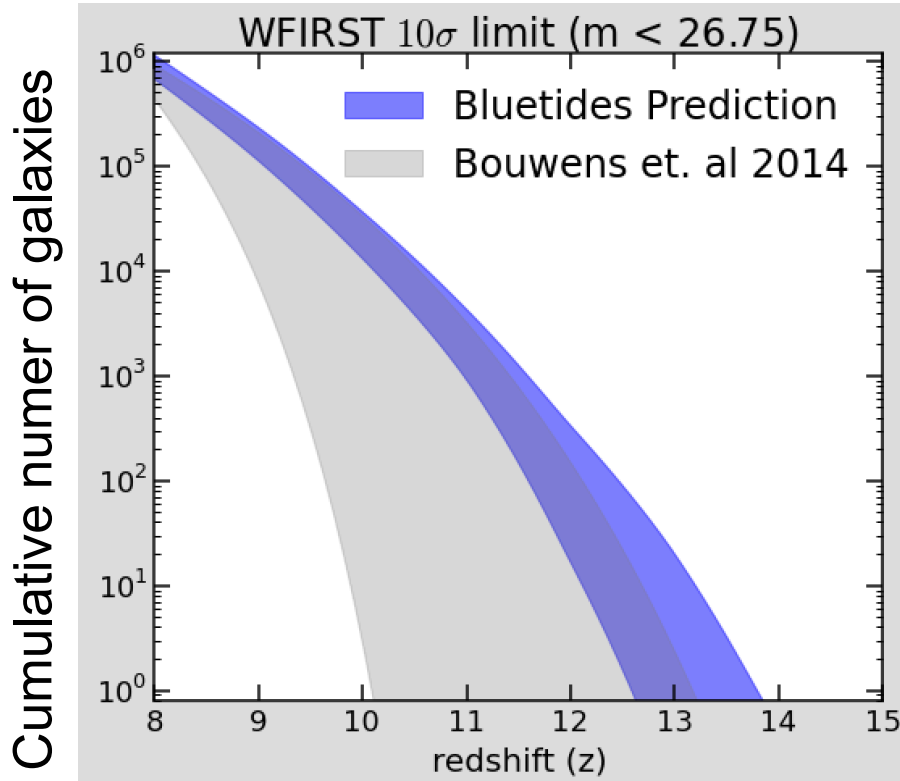
4 times as many galaxies  
as in current HST fields



The end of the dark ages is **bright!**



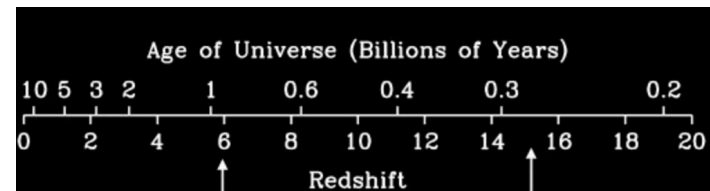
WFIRST



Waters, DM+, 16, in prep

Up to 1M galaxies,  $1e4$  of quasars at  $t = 600\text{Myr}$

**First galaxies at  $t=300\text{Myr}$**



# The end of the dark ages is **bright!**

News Release Number: STScI-2016-07

March 3, 2016 12:00 PM (EST)

## Hubble Team Breaks Cosmic Distance Record

Introduction

The Full Story

Release Images

Release Videos

Fast Facts

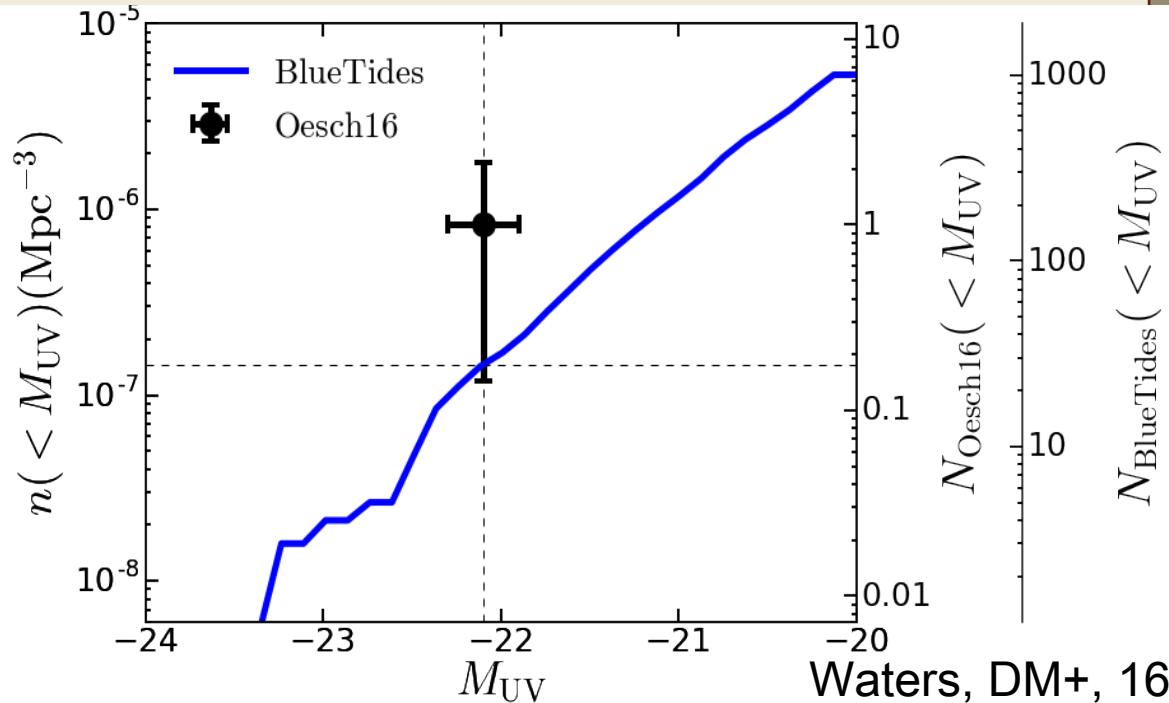
Related Links

### The full news release story:



By pushing NASA's Hubble Space Telescope to its limits, an international team of astronomers has shattered the cosmic distance record by measuring the farthest galaxy ever seen in the universe. This surprisingly bright, infant galaxy, named GN-z11, is seen as it was 13.4 billion years in the past, just 400 million years after the big bang. GN-z11 is located in the direction of the constellation of Ursa Major.

**BlueTides**  
predicts discovery  
of 'bright'  
400 million  
years old galaxy



Waters, DM+, 16

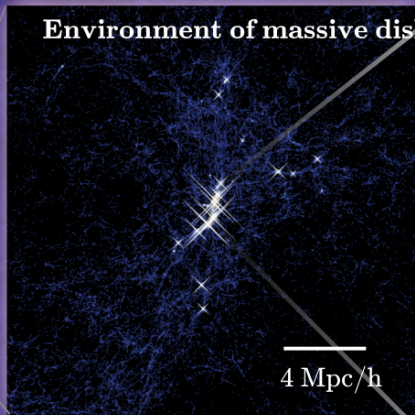


# Conclusions:

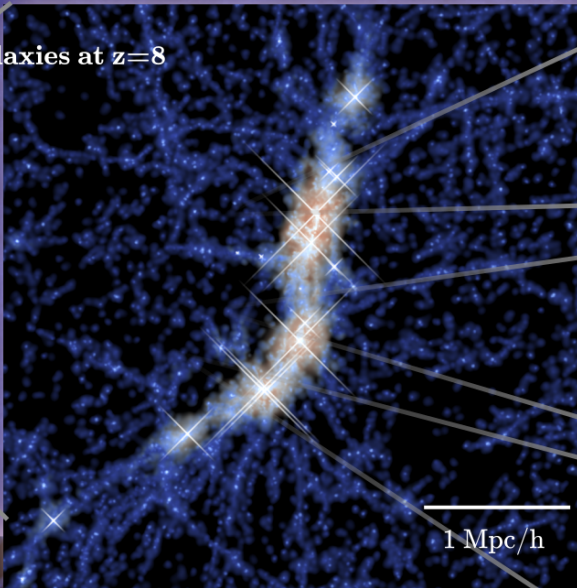
New large volume high res **BlueTides** Simulation predicts (at  $z > 7$ ):

- Population of massive, compact disks --> **WFIRST** should detect thousands
- Most massive black holes,  $10^8 M_{\text{sun}}$  --> **WFIRST** tens of objects

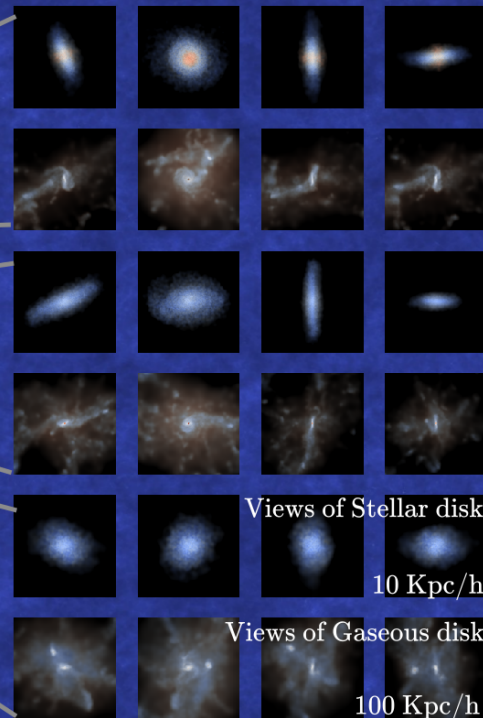
Environment of massive disk galaxies at  $z=8$



4 Mpc/h



1 Mpc/h

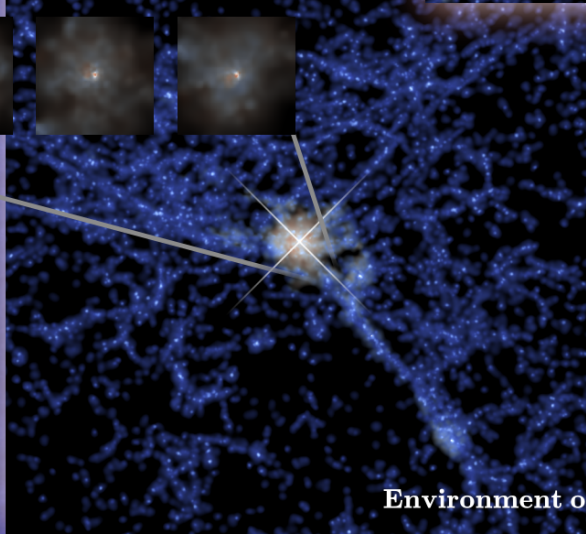
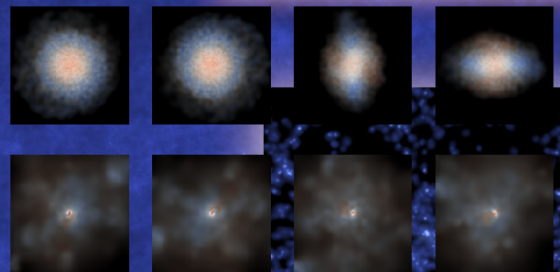


Views of Stellar disk

10 Kpc/h

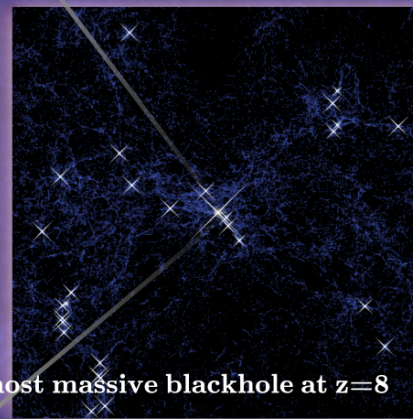
Views of Gaseous disk

100 Kpc/h



Environment of most massive blackhole at  $z=8$

40 Mpc/h



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0.7 trillion particles

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# bluetides

Feng et al. 2015